Shift operators and momentum-space conformal field theory

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D'autant que je m'intéresse moins aux mathématiques, q'aux mathématiciens, comme en tout autre domaine.

Simone Weil [1]

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Abstract

A momentum-space approach to conformal field theory offers a new perspective on cosmological correlators and better reveals the underlying connections to scattering amplitudes. While correlation functions at up to three points are well understood, the form of higherpoint functions is still under active study and few explicit results are available.

A representation for the general *n*-point function of scalar operators was recently proposed in the form of a Feynman integral with the topology of an (n-1)-simplex, featuring an arbitrary function of momentum-space cross ratios. In this thesis, we show the graph polynomials for this integral can all be expressed in terms of the first and second minors of the Laplacian matrix for the simplex. Computing the effective resistance between nodes of the corresponding electrical network, an inverse parametrisation is found in terms of the determinant and first minors of the Cayley-Menger matrix. These parametrisations reveal new families of weight-shifting operators expressible as determinants that connect n-point functions in spacetime dimensions differing by two. Furthermore, they enable the validity of the conformal Ward identities to be established directly without recourse to recursion in the number of points.

We then analyse the representation of conformal, and more general, Feynman integrals through a class of multivariable hypergeometric functions proposed by Gelfand, Kapranov & Zelevinsky. Among other advantages, this formalism enables the systematic construction of highly non-trivial weight-shifting operators known as "creation" operators. We discuss these operators from a physics perspective emphasising their close connection to the spectral singularities that arise for special parameter values, and their relationship to the Newton polytope of the integrand. Via these methods we construct novel weightshifting operators connecting contact Witten diagrams of different operator and spacetime dimensions, as well as exchange diagrams with purely non-derivative vertices.

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Part I Introduction

Chapter 1

Introduction

1.1 Conformal field theory

Conformal symmetry arises in many different physical contexts. Historically, it was first applied in the 70s to the study of critical phenomena [2] and soon after started to play a major role in string theory [3]. In the late 90s, conformal field theories (CFTs) were found to be dual to gravitational theories in Anti de Sitter (AdS) space through a paradigm known as holography [4, 5].

Among the key observables of quantum field theory, cosmology, and condensed matter are correlation functions of operators. These are scale-invariant at critical points, and surprisingly different physical systems are sometimes found to share the same set of critical exponents. For example, the critical exponent for ferromagnets is the same as for water (liquid-vapor transition) [6]. This property is called *universality* and reveals a common underlying conformal symmetry. Polyakov showed that correlators at critical points are indeed invariant under the full conformal group (which also includes special conformal transformations) and opened the path for applications of conformal symmetry in physics and quantum field theory [7]. A program to study the implications of conformal symmetry for scalar and tensorial operators in general spacetime dimension was developed in [8–13]. The idea was to derive the form of correlators by symmetry principles, and this led to looking for solutions of conformal Ward identities in a spacetime dimension d > 2 for general scaling dimensions. These analyses were carried out in position space, where conformal transformations act directly.

The study of conformal anomalies [14–16] and the application of holography to cosmology [17–32] motivated the development of momentum-space conformal field theory. The inflationary epoch is described by an approximately de Sitter spacetime geometry, and the symmetries of this spacetime act on late-time slices as conformal transformations. Therefore, inflationary correlators can equivalently be regarded as CFT correlators. Moreover, momentum-space CFT also found a central role in the study of renormalisation [33–36] and scattering amplitudes, revealing features such as double-copy structure and colour/kinematic duality [37–41].

The analysis of the implications of conformal symmetry in momentum space started ten years ago. The form of 2- and 3-point functions of scalar and tensorial operators in d > 2 are strictly constrained by the symmetry and their unique form was found [42, 43]. Equivalent representations for the 3-point function have been analysed: this can be expressed as an integral over three Bessel functions of the second kind, as a multivariable hypergeometric function Appell F_4 , or as a one-loop triangle Feynman diagram. Except for some special solutions, evaluating these functions can be difficult. A reduction scheme has been described to construct a class of 3-point functions which are also the building blocks of tensorial correlators [44, 33, 34]. This reduction is performed via the action of shift operators that connect different solutions with shifted parameters. Moreover, a full understanding of their singularities and renormalisation has also been discussed [45]. Less complete and understood is instead the form of 4- and higher-point correlators for general values of their parameters. A representation for the general *n*-point function of scalar operators was recently proposed in the form of a Feynman integral with the topology of an (n-1)-simplex, featuring an arbitrary function of momentum-space cross ratios [46, 47]. This was shown to be conformally invariant, and a recursive interpretation of its form was given.

The research presented in this thesis takes its starting point from the simplex representation of n-point functions and develops to explore the interplay between integral representations and shift operators.

1.2 Shift operators

Operators that act on a function to shift one or more of its parameters appear in the physical and mathematical literature with different names. Here, we refer to such operators as *shift operators*.

Historically, their relevance in physics was first revealed in quantum mechanics. To solve the Schrödinger problem of the harmonic oscillator, Dirac introduced shift operators known as the creation and annihilation operators [48]. They were useful both on the computational and physical sides. Indeed, such operators act on an eigenfunction of the Hamiltonian to generate a new eigenstate of the same Hamiltonian, but with shifted eigenvalue. Physically, the action of such shift operators makes an energy quantum $\hbar\omega$ appear or disappear. By knowing these operators, Dirac was able to find the ground state and the full set of solutions of the quantum harmonic oscillator: once the ground state is found, it is sufficient to apply the creation operator to find all the remaining eigenstates.

The idea of introducing operators that act on a quantum state to shift a quantum number was also applied to describe the physics of the quantum angular momentum [49]. In the case of a three-dimensional angular momentum, it turned out that it was convenient to define the shift operators, often referred to as raising/lowering operators. They are complex linear combinations of the quantised spatial x and y components of the angular momentum operator. In an analogy with the system of the harmonic oscillator, one can show that these operators act on an eigenstate of the z component of the angular momentum to increase or decrease it by an angular momentum quantum $m\hbar$. Once the quantisation of the angular momentum is defined, several applications show the utility of such shift operators [50]. Among these, we find the solution of the Schrödinger problem of the hydrogen atom [51, 52], and the study of hydrogen-like systems in solid state physics, the effect of a magnetic field on the energy levels of an atom, and the Zeeman effect [53].

In quantum field theory (QFT), creation and annihilation operators appear in the solution of dynamical equations. For instance, a scalar field is expressed as a superposition of normal modes whose amplitudes of oscillation are given by operators analogous to the creation and annihilation operators of the quantum harmonic oscillator, since they satisfy the same commutation relations. It is interesting to note that according to Pauli's exclusion principle [54], fermions and bosons are described by different symmetries and this property results in the fact that while bosons' shift operators obey commutation relations, fermions' obey anticommutation relations [55]. Therefore shift operators are related to the symmetry of a system. This relation is extensively used in particle physics as well as in statistical physics.

Along with the development of QFT, Feynman integrals became key objects in various areas of physics. Their accurate evaluation became central to the understanding of physical phenomena. With the development of experiment and theory, the need for precision and accuracy increased. Higher orders in perturbative QFT are required in the computation of observables via Feynman diagrams, whose number and complexity increase with the order in perturbation. For this reason, Feynman integrals continue to be an active research topic needed for example in scattering processes, perturbative quantum chromodynamics, lattice computations, CFT and cosmology [56–64].

The problem of computing Feynman integrals is often hard to tackle. The main needs are the reduction of tensorial integrals to scalar integrals and the reduction of the latter into a small (finite) number of Feynman integrals known as master integrals. Various techniques have been explored and continue to be discussed [65]. Among these we find integration-by-parts (IBP) identities [66–68], allowing us to express any Feynman integral as a linear combination of master integrals. These identities stem from taking the total derivative of the integrand. More recently the IBP method has been extended by looking at operators that annihilate the integrand [69]. A complementary method to simplify the computation of Feynman integrals is based on recurrence relations, and the shift operators from which they can be derived [70–72]. This method is based on considering the Feynman integral as a function of its parameters, (*i.e.*, the powers of propagators and/or the spacetime dimension) and by acting with appropriate operators on such an integral, one finds a new integral with shifted parameters. This is also useful in tensorial reductions.

Later, when CFT arose in the study of physical phenomena, the methods involving shift operators and recursion relations appeared in this context. Position-space 4-point functions of scalar operators are expressed via the operator product expansion whose terms, known as conformal blocks, depend on the spacetime dimension and the operator dimensions. Dolan and Osborn showed that such conformal blocks satisfy a second-order differential equation and are related to the eigenfunctions of the quadratic Casimir of the conformal group [73, 74]. They define various sets of shift operators that act on the conformal blocks to shift their parameters. These operators were fundamental for the development of numerical bootstrap methods and found applications, for instance, in the study of the three-dimensional Ising model, helping to find bounds on the physical parameters [75].

Subsequently, various techniques for computing blocks of operators with spin have been developed. Such operators played an important role, for example, in finding new results on the Regge limit in CFTs [76, 77], or universal numerical bounds on classes of CFTs [78–80]. More recently it was also useful for the large-N solution of the SYK model [81, 82]. One of the most promising techniques is the method introduced in [83], based on *weight-shifting* operators. These operators act to increase or decrease the parameters of an operator, for

instance, they can shift the spin. Recently, these shift operators found applications in the computation of inflationary correlators [84]. Moreover, in momentum-space CFT, a set of shift operators was also introduced to compute 3-point functions of tensorial operators [85]. In a similar fashion to Feynman integrals, these operators allowed the computation of tensorial correlators requiring only the knowledge of one (master) scalar correlator.

The search for new shift operators for momentum-space CFT correlators and Feynman integrals that is presented in this thesis was inspired by the works summarised above.

1.3 Outline

The outline of the thesis is as follows. The first part is devoted to illustrating some aspects of the state of the art of conformal field theory. In Chapter 2, we present the essential features of conformal symmetry in position space and derive the consequent constraints on correlators up to n-points. In Chapter 3, we give a broader presentation of conformal symmetry in momentum space, where our research is focused. First, we derive conformal Ward identities in momentum space and construct their solutions up to n points. We give a detailed overview of the properties of 3-point functions by discussing equivalent representations, singularities and shift operators. We then present the simplex representation for n-point functions and conclude with an illustration of some special 4-point solutions to the conformal Ward identities, namely the contact and exchange Witten diagrams. All the ingredients necessary to follow the second part of the thesis are then in place. The second part of the thesis is based on the research that appeared in [63] and [86]. Chapter 4 focuses on the simplex integral. We derive parametric integral representations for the simplex integral. By using inverse Schwinger parameters, we find that all graph polynomials for this integral can be expressed in terms of the first and second minors of the Laplacian matrix for the simplex. Inspired by the analogy between the simplicial geometry and electrical circuits, we regard the Schwinger parameters as resistances in an electrical network and re-parametrise the simplex integral by computing the effective resistance between all vertices of the simplex. This gives a representation in terms of the determinant and first minors of the Cayley-Menger matrix. These parametrisations have various advantages. The diagonal structure of the exponential factor in the integrand allows a Fourier-like correspondence. This reveals new families of shift operators, expressible as determinants, that connect solutions of the conformal Ward identities in spacetime dimension d to new solutions in dimension d+2. They are the generalisation to *n*-point of the known 3-point shift operators. Moreover, these novel representations reduce the number of scalar integrals and allow us to verify that the conformal Ward identities are satisfied via direct computation. Different integral representations may give complementary perspectives on the same object. Motivated by the description of some 3- and 4-point conformal correlators in terms of hypergeometric functions, in Chapter 5 we then move to analyse the representation of conformal, and other more general Feynman integrals through a class of multivariable hypergeometric functions proposed by Ge'lfand, Kapranov & Zelevinsky known as GKZ functions. A Feynman integral in GKZ form is characterised by a unique denominator and a higher-dimensional space of variables (in the context of Feynman integrals, these are momenta and masses). The strength of this representation is that all the properties of the function can be encoded in a matrix. From this, we can derive a set of partial differential equations satisfied by the integral and the singularities in the parameters (for

physical integrals, these are represented by the spacetime dimension and the generalised propagators). These spectral singularities are given by an infinite number of hyperplanes parallel to the facets of the Newton polytope associated with the matrix. We discuss how the knowledge of these singularities is the starting point for a systematic construction of non-trivial shift operators known as creation operators. For 3-point CFT correlators, these are indeed the inverse operators of the shift operators involved in the reduction scheme, acting on a 3-point function to lower d by two. We derive these shift operators for various Feynman integrals and for special classes of 4-point (and n-point) conformal correlators, such as contact Witten diagrams, consisting of integral over multiple Bessel functions. Using this formalism, we also derive novel weight-shifting operators connecting contact and exchange Witten diagrams with different operators [84, 36], these novel shift operators generate shifted exchange diagrams with purely non-derivative vertices and can be applied for any values of the parameters. Finally, in Chapter 6, we conclude with a summary and open questions.

Part II Conformal field theory

Chapter 2

Conformal field theory in position space

In this chapter we briefly introduce the main features of conformal symmetry. We define conformal transformations (translations, rotations, dilatations and special conformal transformations) and describe the conformal group by finding the generators of these transformations and their commutation relations. We then analyse the consequences of conformal symmetry on correlation functions of scalar operators: symmetry imposes strong constraints on their form. This material is discussed in many books and reviews from where the content of this chapter is inspired [87–89].

2.1 Conformal transformations

In this section we present conformal transformations and explain their geometrical meaning.

Let us consider a spacetime with metric $g_{\mu\nu}$. A Weyl transformation is a spacetimedependent rescaling of the metric sending the initial $g_{\mu\nu}$ to the rescaled $g'_{\mu\nu}$ [90],

$$g_{\mu\nu} \to g'_{\mu\nu} = e^{2\sigma(x)} g_{\mu\nu}, \qquad (2.1)$$

where $\sigma(x)$ is a generic function of the coordinates defining the rescaling factor $e^{2\sigma(x)}$. The infinitesimal version of this transformation is

$$g_{\mu\nu} \rightarrow g'_{\mu\nu} = g_{\mu\nu} + \delta g_{\mu\nu}, \quad \text{with} \quad \delta g_{\mu\nu} = 2\sigma(x)g_{\mu\nu}.$$
 (2.2)

In general, the effect of a Weyl transformation is a change of the spacetime geometry. If we consider the initial spacetime to be flat, with metric $\eta_{\mu\nu}$, a general Weyl transformation sends this flat metric to $g'_{\mu\nu}$, a curved spacetime metric. We look for special Weyl transformations ($\sigma(x)$) that can be undone by a special diffeomorphism so that the metric remains flat. Let us consider the diffeomorphism

$$x^{\mu} \rightarrow x^{\mu} - \xi^{\mu}, \qquad (2.3)$$

then, the corresponding infinitesimal transformation of the metric reads

$$g_{\mu\nu} \rightarrow g'_{\mu\nu} = g_{\mu\nu} + \delta_{\xi} g_{\mu\nu}, \quad \text{with} \quad \delta_{\xi} g_{\mu\nu} = 2\nabla_{(\mu} \xi_{\nu)}, \tag{2.4}$$

where ∇ denotes the covariant derivative. In order to leave the spacetime metric flat, we require the overall change in the metric – due to the Weyl transformation and the diffeomorphism – to vanish. In other words, we require the following condition to be satisfied:

$$\delta g_{\mu\nu} = \delta_{\sigma} g_{\mu\nu} + \delta_{\xi} g_{\mu\nu} = 0. \tag{2.5}$$

Evaluating (2.2) and (2.4) on a flat metric, this reads

$$2\partial_{(\mu}\xi_{\nu)} = -2\sigma\eta_{\mu\nu},\tag{2.6}$$

where we now wrote the standard partial derivative and $\eta_{\mu\nu}$ denotes the (flat) Euclidean metric $\eta_{\mu\nu} = \text{diag}(1, 1, ..., 1)$. Note that in this thesis we will work in Euclidean signature. By contracting equation (2.6) we find the relation between the function $\sigma(x)$ and the vector ξ^{μ}

$$\sigma = -\frac{1}{d}\partial_{\mu}\xi^{\mu}.$$
(2.7)

Substituting (2.7) in (2.6), we obtain

$$\partial_{(\mu}\xi_{\nu)} = \frac{1}{d}\partial_{\rho}\xi^{\rho}\eta_{\mu\nu}.$$
(2.8)

This is the conformal Killing equation and defines the condition that ξ^{μ} must satisfy to generate a diffeomorphism acting on the metric to undo the Weyl transformation. In the following we show that when d > 2 the conformal Killing equation (2.8) has a finite number of solutions, while when d = 2 an infinite number of solutions exists. However, in this thesis we are interested in spacetime dimensions $d \ge 3$. To find the general solution of (2.8), we act with the partial derivative ∂_{ρ} on (2.6), giving a second-order differential equation. By taking a linear combination of this equation with Lorentz indexes cyclically permuted, we obtain

$$\partial_{\mu}\partial_{\nu}\xi_{\rho} = -\eta_{\mu\rho}\partial_{\nu}\sigma - \eta_{\nu\rho}\partial_{\mu}\sigma + \eta_{\mu\nu}\partial_{\rho}\sigma, \qquad (2.9)$$

then by contracting it

$$\partial^2 \xi_\mu = (d-2)\partial_\mu \sigma. \tag{2.10}$$

We now act with ∂_{ν} on this last equation and with ∂^2 on (2.6). Combining the resulting expressions we have

$$(d-2)\partial_{\mu}\partial_{\nu}\sigma = -(\partial^2\sigma)\eta_{\mu\nu}, \qquad (2.11)$$

whose contraction gives

$$2(d-1)\partial^2 \sigma = 0. \tag{2.12}$$

Hence, $\partial^2 \sigma = 0$ for d > 1. Consequently, from (2.11) we deduce that

$$(d-2)\partial_{\mu}\partial_{\nu}\sigma = 0. \tag{2.13}$$

Thus, for d > 2 the following condition holds

$$\partial_{\mu}\partial_{\nu}\sigma = 0, \tag{2.14}$$

which amounts to say that σ is at most linear in the spacetime coordinates x^{μ} and, according to (2.7), this means that the Killing vector ξ^{μ} is at most quadratic in x^{μ} :

$$\xi_{\mu} = A_{\mu} + B_{\mu\nu}x^{\nu} + C_{\mu\nu\rho}x^{\nu}x^{\rho}, \qquad (2.15)$$

with A_{μ} , $B_{\mu\nu}$ and $C_{\mu\nu\rho}$ some coefficients we are going to find. To this aim, we substitute (2.15) back into equation (2.6) and find that $A_{\mu} \equiv a_{\mu}$ is an arbitrary constant vector, while $B_{\mu\nu}$ is the sum of an antisymmetric term $\omega_{\mu\nu}$ and a symmetric term proportional to the metric

$$B_{\mu\nu} = \omega_{\mu\nu} + \lambda \eta_{\mu\nu}. \tag{2.16}$$

Finally, taking into account that $C_{\mu\nu\rho}$ is symmetric in the last two indexes, it must be of the form

$$C_{\mu\nu\rho} = -b_\rho \eta_{\mu\nu} + b_\mu \eta_{\nu\rho} - b_\nu \eta_{\mu\rho}, \qquad (2.17)$$

where b_{ν} is an arbitrary constant vector. Hence, the general solution of the conformal Killing equation for $d \geq 3$ is

$$\xi^{\mu} = a^{\mu} + \omega^{\mu}{}_{\nu}x^{\nu} + \lambda x^{\mu} + b^{\mu}x^2 - 2(b_{\nu}x^{\nu})x^{\mu}.$$
(2.18)

The infinitesimal change of coordinates generated by this conformal Killing vector x_i^{μ} defines four class of transformations:

- 1. translations: $\xi_T^{\mu} = a^{\mu}$,
- 2. rotations: $\xi^{\mu}_{R} = \omega^{\mu}{}_{\nu}x^{\nu}$,
- 3. scale transformations (dilatation): $\xi_D^{\mu} = \lambda x^{\mu}$,
- 4. special conformal transformations (SCT): $\xi^{\mu}_{SCT} = b^{\mu}x^2 2(b_{\nu}x^{\nu})x^{\mu}$.

This is what we anticipated at the beginning, *i.e.*, that conformal transformations define a group larger than the Poincaré one, by including dilatations and special conformal transformations. Let us note that translations and rotations (or Lorentz transformations in the case we are considering the Minkowski metric, instead of the Euclidean one) are isometries, in fact $\sigma(x) = 0$ both for ξ_T^{μ} and ξ_R^{μ} . For scale transformations, the metric is rescaled by a constant $\sigma(x) = -\lambda$, independent of the spacetime coordinates. Finally, for SCT we find $\sigma(x) = 2b \cdot x$, which corresponds to a spacetime-dependent rescaling. While the first three transformations are intuitive to visualise, the special conformal transformations defined by

$$x^{\mu} \rightarrow x^{\mu} - \xi^{\mu}_{\text{SCT}} = x^{\mu} - b^{\mu}x^{2} + 2(b_{\nu}x^{\nu})x^{\mu},$$
 (2.19)

are harder to visualise. However, we can see them as a combination of an inversion, a translation by b^{μ} and an inversion again:

$$x^{\mu} \rightarrow \frac{x'^{\mu}}{x'^2} = \frac{x^{\mu}}{x^2} - b^{\mu},$$
 (2.20)

leading to

$$x^{\prime \mu} = \frac{x^{\mu} - b^{\mu} x^2}{1 - 2\mathbf{b} \cdot \mathbf{x} + b^2 x^2}.$$
(2.21)

This corresponds to the finite version of the special conformal transformations (2.19) as one can see by expanding the last expression for an infinitesimal vector b^{μ} , recovering $x'^{\mu} = x^{\mu} - \xi^{\mu}_{SCT}$.

To conclude this section, we show that conformal transformations act locally as the composition of a rotation and a scale transformation. To see this, we compute the Jacobian associated to conformal transformations

$$\left|\frac{\partial x^{\prime \mu}}{\partial x^{\nu}}\right| = (1 - \lambda + 2b \cdot x)\delta^{\mu}{}_{\nu} - \omega^{\mu}{}_{\nu} + 2(b_{\nu}x^{\mu} - b^{\mu}x_{\nu})$$
$$\approx (1 - \lambda + 2b \cdot x)\left(\delta^{\mu}{}_{\nu} - \omega^{\mu}{}_{\nu} + 2(b_{\nu}x^{\mu} - b^{\mu}x_{\nu})\right) = e^{\sigma}R^{\mu}{}_{\nu}, \qquad (2.22)$$

where $R^{\mu}{}_{\nu} = \delta^{\mu}{}_{\nu} - \omega^{\mu}{}_{\nu} + 2(b_{\nu}x^{\mu} - b^{\mu}x_{\nu})$ is an orthogonal matrix responsible for the rotation, while the factor $(1 - \lambda + 2b \cdot x) = e^{\sigma}$ is the local scale transformation. It is now patent why the name *conformal*: conformal transformations preserve angles.

2.2 Conformal group

Conformal transformations form a group, i.e., given the infinitesimal conformal transformation

$$x^{\mu} \rightarrow x'^{\mu} = x^{\mu} - \xi^{\mu}_T - \xi^{\mu}_R - \xi^{\mu}_D - \xi^{\mu}_{SCT},$$
 (2.23)

the identity and inverse elements exist plus the composition of two conformal transformations is still a conformal transformation. Moreover, it is a continuous group since it acts on the spacetime coordinates. This means that we can describe it through its generators and the commutation relations amongst them. We start by finding the generator of translations. Let us assume f(x) to be an element of the conformal group, and act with an infinitesimal translation described by the parameter a^{μ}

$$f(x^{\mu}) \to f(x'^{\mu}) = f(x^{\mu} - a^{\mu}) \approx (1 - iP_{\mu}a^{\mu})f(x^{\mu}), \qquad P_{\mu} = -i\partial_{\mu},$$
 (2.24)

we found the expected generator for translations P_{μ} . The finite version of the transformation is then given by exponentiating the generator

$$f(x) \rightarrow e^{ia_{\mu}P^{\mu}}f(x), \qquad (2.25)$$

therefore the generators are fundamental to define the group. Knowing the four infinitesimal transformations defining the conformal group, we can find the associated generators:

$$P_{\mu} = -i\partial_{\mu} \qquad (\text{translations}), \qquad (2.26)$$

$$M_{\mu\nu} = -i(x_{\mu}\partial_{\nu} - x_{\nu}\partial_{\mu}) \qquad (\text{rotations/Lorentz transformations}), \qquad (2.27)$$

$$D = -ix^{\mu}\partial_{\mu} \tag{2.28}$$

$$K_{\mu} = -i(x^2 \partial_{\mu} - 2x_{\mu} x_{\nu} \partial^{\nu}) \qquad (\text{special conformal transformations}). \qquad (2.29)$$

The algebra of the conformal group is then defined by the following commutation relations

$$[M_{\mu\nu}, M_{\rho\sigma}] = -i(\eta_{\mu\rho}M_{\nu\sigma} - \eta_{\mu\sigma}M_{\nu\rho} - \eta_{\nu\rho}M_{\mu\sigma} + \eta_{\nu\sigma}M_{\mu\rho}),$$

$$[M_{\mu\nu}, P_{\rho}] = -i(\eta_{\mu\rho}P_{\nu} - \eta_{\nu\rho}P_{\mu}),$$

$$[D, P_{\mu}] = iP_{\mu},$$

$$[P_{\mu}, K_{\nu}] = 2i(\eta_{\mu\nu}D - M_{\mu\nu}),$$

$$[D, K_{\mu}] = -iK_{\mu},$$

$$[M_{\mu\nu}, K_{\rho}] = -i(\eta_{\mu\rho}K_{\nu} - \eta_{\nu\rho}K_{\mu}),$$

$$[P_{\mu}, P_{\nu}] = [D, D] = [M_{\mu\nu}, D] = [K_{\mu}, K_{\nu}] = 0.$$
(2.30)

We can show that the Euclidean conformal algebra in d dimensions is isomorphic to the algebra of the Lorentz group SO(d+1,1). In fact, let us define the operators $J_{ab} = -J_{ba}$, for a, b = -1, 0, 1, ..., d:

$$J_{\mu\nu} = M_{\mu\nu}, \qquad \qquad J_{-1,0} = D, \qquad (2.31)$$

$$J_{-1,\mu} = \frac{1}{2}(P_{\mu} - K_{\mu}), \qquad \qquad J_{0,\mu} = \frac{1}{2}(P_{\mu} + K_{\mu}). \qquad (2.32)$$

They satisfy the commutation relations of the SO(d+1, 1) algebra:

$$[J_{ab}, J_{cd}] = -i(\eta_{ac}M_{bd} - \eta_{ad}M_{bc} - \eta_{bc}M_{ad} + \eta_{bd}M_{ac}), \qquad (2.33)$$

where $\eta_{ab} = (-1, ..., 1, 1, 1)$. In the same way one can show that the Lorentzian conformal d-dimensional group is isomorphic to SO(d, 2). Finally, it is worth noticing that the commutation relations above show that while $M_{\mu\nu}$ and P_{μ} form a group, which is the Poincaré group, the generators $M_{\mu\nu}$, P_{μ} and D also form a group. This implies that if we enhanced the Poincaré group only by introducing the scale transformation, we would not obtain the full conformal group.

2.3 Conformal transformations for operators

In the first section we defined the action of conformal transformations on the spacetime coordinates and the metric. However, in a conformal field theory, the fields also transform. To find their transformation we combine, as before, a general Weyl transformation with the diffeomorphism found in section 1 and we find that, while we require the metric to stay flat, other fields do not transform trivially. According to Weyl transformations, if the metric transforms as $g_{\mu\nu} \rightarrow g'_{\mu\nu} = e^{2\sigma(x)}g_{\mu\nu}$, then a Weyl transformation of a scalar field \mathcal{O} is

$$\mathcal{O} \rightarrow \mathcal{O}' = \mathrm{e}^{-\Delta\sigma} \mathcal{O},$$
 (2.34)

where Δ is the Weyl weight. Note that for scalar fields the Weyl weight coincides with the scaling dimension of the operator. For instance in a Weyl-invariant free scalar field theory described by the action

$$S = -\frac{1}{2} \int \mathrm{d}^d x \sqrt{-g} \left(g_{\mu\nu} \partial^\mu \mathcal{O} \partial^\nu \mathcal{O} + \zeta R \mathcal{O}^2 \right), \qquad (2.35)$$

one finds that the Weyl weight is the same as the canonical scaling dimension Δ , *i.e.*, $\Delta = \frac{d}{2} - 1$. Moreover, conformal invariance also requires $\zeta = (d-2)/(4(d-1))$. In the following discussion, however, we will not need any Lagrangian to study conformal field theory.

To determine the infinitesimal conformal transformation of scalar fields we compose the infinitesimal Weyl transformation

$$\delta_{\sigma}\mathcal{O} = -\Delta\sigma\mathcal{O} \tag{2.36}$$

with the transformation of the scalar field due to the diffeomorphism

$$\delta_{\xi} \mathcal{O} = \xi^{\mu} \partial_{\mu} \mathcal{O}, \qquad (2.37)$$

giving

$$\delta \mathcal{O} = \delta_{\sigma} \mathcal{O} + \delta_{\xi} \mathcal{O} = \left[\xi^{\mu} \partial_{\mu} + \frac{\Delta}{d} (\partial_{\mu} \xi^{\mu}) \right] \mathcal{O}, \qquad (2.38)$$

where we used (2.7) to express the Weyl parameter σ in terms of the vector ξ^{μ} . To find how the field \mathcal{O} transforms under purely dilatations or special conformal transformations, we consider $\xi^{\mu} = \xi^{\mu}_{D} = \lambda x^{\mu}$ and $\xi^{\mu} = \xi^{\mu}_{\text{SCT}} = b^{\mu} x^{2} - 2(b_{\nu} x^{\nu}) x^{\mu}$ in the general relation (2.38):

$$\delta_D \mathcal{O}(x) = \lambda(x^\mu \partial_\mu + \Delta) \mathcal{O}(x), \qquad (2.39)$$

$$\delta_{\text{SCT}}\mathcal{O}(x) = b_{\nu} \left[(x^2 \eta^{\mu\nu} - 2x^{\mu} x^{\nu}) \partial_{\mu} - 2\Delta x^{\nu} \right] \mathcal{O}(x).$$
(2.40)

The finite conformal transformation for scalar operators is

$$\mathcal{O}(x) \rightarrow \mathcal{O}'(x') = \left| \frac{\partial x'}{\partial x} \right|^{-\frac{\Delta}{d}} \mathcal{O}(x),$$
 (2.41)

which we can see as a rescaling of the field \mathcal{O} by a power $-\Delta$ of the length rescaling factor $|\partial x'/\partial x|^{1/d} = e^{\sigma}$. In fact $|\partial x'/\partial x|$ is a local hypervolume in d dimensions, consequently $|\partial x'/\partial x|^{1/d}$ is a length rescale. This transformation rule defines a so-called *primary* operator.

2.4 Conformal Ward Identities

In this section we analyse the consequences of conformal symmetry. We will show that conformal symmetry imposes constraints on the observables allowing a non-perturbative approach to CFT.

In field theory, invariance implies the following equivalence between correlators [55]

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\cdots\mathcal{O}_n(\boldsymbol{x}_n)\rangle = \langle \mathcal{O}_1'(\boldsymbol{x}_1)\mathcal{O}_1'(\boldsymbol{x}_2)\cdots\mathcal{O}_n'(\boldsymbol{x}_n)\rangle, \qquad (2.42)$$

where $\mathcal{O}_i(\boldsymbol{x}_i)$ is a scalar operator with scaling dimension Δ_i and $\mathcal{O}'_i(\boldsymbol{x}_i)$ is the transformed operator. One can show that the above equation holds by considering the path-integral

formalism for correlators. Let $S_{\rm CI}$ be a conformally invariant action, then

$$\langle \mathcal{O}_{1}(\boldsymbol{x}_{1})\mathcal{O}_{2}(\boldsymbol{x}_{2})\cdots\mathcal{O}_{n}(\boldsymbol{x}_{n})\rangle = \int [\mathcal{D}\mathcal{O}] \mathcal{O}_{1}(\boldsymbol{x}_{1})\mathcal{O}_{2}(\boldsymbol{x}_{2})\cdots\mathcal{O}_{n}(\boldsymbol{x}_{n})\mathrm{e}^{-S_{\mathrm{CI}}[\mathcal{O}]}$$

$$= \int [\mathcal{D}\mathcal{O}'] \mathcal{O}'_{1}(\boldsymbol{x}_{1})\mathcal{O}'_{2}(\boldsymbol{x}_{2})\cdots\mathcal{O}'_{n}(\boldsymbol{x}_{n})\mathrm{e}^{-S_{\mathrm{CI}}[\mathcal{O}']}$$

$$= \int [\mathcal{D}\mathcal{O}] \mathcal{O}'_{1}(\boldsymbol{x}_{1})\mathcal{O}'_{2}(\boldsymbol{x}_{2})\cdots\mathcal{O}'_{n}(\boldsymbol{x}_{n})\mathrm{e}^{-S_{\mathrm{CI}}[\mathcal{O}]}$$

$$= \langle \mathcal{O}'_{1}(\boldsymbol{x}_{1})\mathcal{O}'_{2}(\boldsymbol{x}_{2})\cdots\mathcal{O}'_{n}(\boldsymbol{x}_{n})\rangle, \qquad (2.43)$$

where in the second line we renamed $\mathcal{O} \to \mathcal{O}'$, while in the third line we considered conformal invariance both of the action and the functional measure $[\mathcal{D}\mathcal{O}]$. When the latter is not invariant, however, there will be an anomaly corresponding to the symmetry breaking and equation (2.42) will be modified with an additional term. This is related to renormalisation which we will discuss later in Chapter 3. In the following we assume the symmetry is not broken. At the infinitesimal level, equation (2.42) reads

$$0 = \delta \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle$$

= $\sum_{i=1}^n \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \delta \mathcal{O}_i(\boldsymbol{x}_i) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle.$ (2.44)

This equation amounts to a set of differential equations, known as conformal Ward identities (CWIs). To obtain their expressions, we express $\delta \mathcal{O}_i(\boldsymbol{x}_i) = \mathcal{O}(\boldsymbol{x}'_i) - \mathcal{O}(\boldsymbol{x})$ in (2.44) using (2.38). For translations, $\delta \mathcal{O}(\boldsymbol{x}) = a^{\mu} \partial_{\mu} \mathcal{O}(\boldsymbol{x})$ and the Ward identity reads

$$\sum_{i=1}^{n} \frac{\partial}{\partial x_{i}^{\mu}} \langle \mathcal{O}_{1}(\boldsymbol{x}_{1}) \mathcal{O}_{2}(\boldsymbol{x}_{2}) \cdots \mathcal{O}_{n}(\boldsymbol{x}_{n}) \rangle = 0, \qquad (2.45)$$

while rotation Ward identity is

$$\sum_{i=1}^{n} \left(x_i^{\mu} \partial_i^{\nu} - x_i^{\nu} \partial_i^{\mu} \right) \left\langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \right\rangle = 0.$$
(2.46)

These two transformations require the scalar n-point correlator to be a function of the coordinate separations

$$x_{ij} = |\mathbf{x}_i - \mathbf{x}_j|, \quad i, j = 1, ..., n.$$
 (2.47)

Less trivial are the constraints imposed by dilatation and special conformal Ward identities which read respectively

$$\sum_{i=1}^{n} \left(x_i^{\mu} \frac{\partial}{\partial x_i^{\mu}} + \Delta_i \right) \left\langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \right\rangle = 0,$$
(2.48)

$$\sum_{i=1}^{n} \left\{ \left(x_{i}^{2} \eta^{\mu\nu} - 2x_{i}^{\mu} x_{i}^{\nu} \right) \frac{\partial}{\partial x_{i}^{\mu}} - 2\Delta_{i} x_{i}^{\nu} \right\} \left\langle \mathcal{O}_{1}(\boldsymbol{x}_{1}) \mathcal{O}_{2}(\boldsymbol{x}_{2}) \cdots \mathcal{O}_{n}(\boldsymbol{x}_{n}) \right\rangle = 0, \qquad (2.49)$$

where we used equations (2.39) and (2.40). The dilatation Ward identity (DWI) implies

that the correlator has to be a homogeneous function of the positions of degree $-\Delta_t$, with $\Delta_t = \sum_{i=1}^n \Delta_i$. As it is now evident, each Ward identity constrains the shape of correlators. Finally, let us mention that conformal symmetry also implies the following equivalence between correlators

$$\langle \mathcal{O}_1(\boldsymbol{x}_1')\cdots\mathcal{O}_n(\boldsymbol{x}_n')\rangle = \left|\frac{\partial \boldsymbol{x}_1'}{\partial \boldsymbol{x}_1}\right|^{-\Delta_1/d}\cdots \left|\frac{\partial \boldsymbol{x}_n'}{\partial \boldsymbol{x}_n}\right|^{-\Delta_n/d} \langle \mathcal{O}_1(\boldsymbol{x}_1)\cdots\mathcal{O}_n(\boldsymbol{x}_n)\rangle, \quad (2.50)$$

which directly stems from equation (2.41). This is how the correlator transforms under a finite conformal transformation.

In the following we will list the solutions for 2-, 3-, 4- and n-point scalar correlators obtained by solving these constraints.

2.5 Position-space conformal correlators

In this section we present the solutions of position-space conformal Ward identities. We will show that up to 3-point functions, the solution is unique.

First, let us note that the 1-point function vanishes

$$\langle \mathcal{O}_1(\boldsymbol{x}_1) \rangle = 0. \tag{2.51}$$

In fact, translations and rotations require it to be a constant, $\langle \mathcal{O}_1(\boldsymbol{x}_1) \rangle = \text{const.}$ A nonvanishing constant, however, would violate scaling invariance. To find solutions for $n \geq 2$ we need instead the full set of CWIs as we show in the following sections.

2.5.1 2-point function

Translation and rotation symmetries imply that the 2-point correlator is a function of x_{12} , while dilatations require the correlator to be a homogeneous function in the positions with degree $-\Delta_t$. Hence the 2-point correlator must be of the form

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\rangle \propto x_{12}^{-\Delta_1-\Delta_2}.$$
 (2.52)

Finally, the SCWI imposes a further constraint on the scaling dimension. By acting with the SCWI on the 2-point function above, one finds that a non-vanishing solution exists if and only if $\Delta_1 = \Delta_2 = \Delta$. Therefore the general 2-point scalar function is

$$\left\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\right\rangle = \begin{cases} C_{12}x_{12}^{-2\Delta}, & \Delta_1 = \Delta_2 = \Delta\\ 0, & \Delta_1 \neq \Delta_2 \end{cases},$$
(2.53)

where C_{12} is a normalisation constant, it can be set to one by normalising the scalar operators.

2.5.2 3-point function

The solution of 3-point function CWIs is also unique. As above, translation and rotation invariance requires the 3-point scalar correlator to be a function of x_{ij} , with $i \neq j = 1, 2, 3$

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\mathcal{O}_3(\boldsymbol{x}_3)\rangle = f(x_{12}, x_{13}, x_{23}).$$
 (2.54)

The dilatation Ward identity specifies the function to be

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\mathcal{O}_3(\boldsymbol{x}_3)\rangle \propto x_{12}^{2\alpha_{12}}x_{13}^{2\alpha_{13}}x_{23}^{2\alpha_{23}},$$
 (2.55)

where α_{ij} are some constants that satisfy

$$\sum_{1 \le i < j \le 3} \alpha_{ij} = -\Delta_t. \tag{2.56}$$

Finally, the SCWI fixes uniquely the values of the parameters α_{ij} , since the terms of the sum in equation (2.49) are three, as the number of the unknown parameters. Instead of substituting the ansatz in the SCWI, we use equation (2.50) to find their values. Squaring equation (2.21), one can show that

$$x_{ij}^{\prime 2} = \frac{x_{ij}^2}{\gamma_i \gamma_j},\tag{2.57}$$

where we defined $\gamma_i = 1 - 2\boldsymbol{b} \cdot \boldsymbol{x}_i + b^2 x_i^2$. And taking into account that for SCTs

$$\left|\frac{\partial \boldsymbol{x}_{i}'}{\partial \boldsymbol{x}_{i}}\right|^{1/d} = \gamma_{i}^{-1}, \qquad (2.58)$$

using equation (2.50) we find

$$\Delta_i = -\sum_{j=1}^3 \alpha_{ij}, \quad i = 1, 2, 3, \tag{2.59}$$

which fixes α_{ij} to

$$2\alpha_{12} = 2\Delta_3 - \Delta_t, \tag{2.60}$$

along with cyclic permutations. Hence the general solution of 3-point scalar CWIs is unique and it reads

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\mathcal{O}_3(\boldsymbol{x}_3)\rangle = \frac{C_{123}}{x_{12}^{\Delta_1+\Delta_2-\Delta_3}x_{13}^{\Delta_1-\Delta_2+\Delta_3}x_{23}^{-\Delta_1+\Delta_2+\Delta_3}}.$$
 (2.61)

Note that while the constant C_{12} could be set to one, this is not possible with the constant C_{123} . In fact, the latter is related to physical properties and is called "OPE constant" or "structure constant".

2.5.3 4-point function

While symmetry fixes 2- and 3-point functions completely up to constants, 4-point and higher-point functions are not uniquely fixed. However, conformal symmetry imposes strong constraints on their form. Here we will analyse the solution of 4-point CWIs and in the next section we will generalise the result to n-point functions.

Translation, rotation and dilatation invariance constrain the form of the solution to be

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\mathcal{O}_3(\boldsymbol{x}_3)\mathcal{O}_4(\boldsymbol{x}_4)\rangle \propto \prod_{1\leq i< j\leq 4} x_{ij}^{2\alpha_{ij}},$$
 (2.62)

where

$$\sum_{1 \le i < j \le 4} 2\alpha_{ij} = -\Delta_t, \tag{2.63}$$

and without loss of generality we assume $\alpha_{ij} = \alpha_{ji}$ and $\alpha_{ii} = 0$.

Note that in this case the number of coordinate separations x_{ij} is larger than the number n of constraints following from the SCWI. To be more specific, there are n(n-1)/2 coordinate separations and n constraints from the SCWI. This implies that there are n(n-3)/2 degree of freedom in the general solution. As showed earlier, under special conformal transformations, equation (2.57) holds, therefore

$$\Delta_i = -\sum_{j=1}^4 \alpha_{ij}, \quad i = 1, .., 4.$$
(2.64)

Note that, unlike for n = 3, this condition does not fully fix the parameters α_{ij} . Moreover, due to equation (2.57), 4-point functions admit two simple conformal invariants, the so called *conformal cross ratios*:

$$u = \frac{x_{12}^2 x_{34}^3}{x_{13}^2 x_{24}^2}, \qquad v = \frac{x_{13}^2 x_{24}^2}{x_{14}^2 x_{23}^2}.$$
 (2.65)

Therefore, the general 4-point function also depends on an arbitrary function f of cross ratios:

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\mathcal{O}_2(\boldsymbol{x}_2)\mathcal{O}_3(\boldsymbol{x}_3)\mathcal{O}_4(\boldsymbol{x}_4)\rangle = f(u,v)\prod_{1\le i< j\le 4} x_{ij}^{2\alpha_{ij}},$$
(2.66)

where the parameters α_{ij} are related to the scaling dimension as in equation (2.64). In the following section we generalise this result to *n*-point and discuss the dependence of the number of independent cross ratios on the spacetime dimension *d* and the number of points *n*.

2.5.4 *n*-point function

The result given in the previous section can be generalised to n points. The general solution is the following conformally invariant n-point function of scalar operators $\mathcal{O}_1, ..., \mathcal{O}_n$ with scaling dimensions $\Delta_1, ..., \Delta_n$:

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\cdots\mathcal{O}_n(\boldsymbol{x}_n)\rangle = \prod_{1\leq i< j\leq n} x_{ij}^{2\alpha_{ij}} f(\boldsymbol{u}),$$
 (2.67)

where the parameters α_{ij} are related to the scaling dimensions by the formula implied by special conformal invariance

$$\Delta_i = -\sum_{j=1}^n \alpha_{ij}, \quad i = 1, 2, ..., n,$$
(2.68)

and f is an arbitrary function of $\mathcal{N}_{d,n}$ independent cross ratios

$$u_{pqrs} = \frac{x_{pr}^2 x_{qs}^2}{x_{pq}^2 x_{rs}^2},$$
(2.69)

where p, q, r, s = 1, 2, ..., n. We denote the set of all independent cross ratios with the symbol u. As discussed in [91–93], the number of independent cross ratios $N_{d,n}$ depends on the number of points n and the spacetime dimension d

$$N_{d,n} = n(n-3)/2, n \le d+2, N_{d,n} = nd - (d+2)(d+1)/2, n > d+2. (2.70)$$

To understand this counting, let us consider n points $x_1, ..., x_n$ in a d-dimensional spacetime. Using conformal transformations some of these n points can be fixed in the spacetime. For example, x_1 can be sent to infinity by using a special conformal transformation, while using translations x_2 can be fixed at the origin. Then x_3 can be set at (1, 0, ..., 0)by performing a rotation together with a dilatation (which fixes the non-zero coordinate to be equal to one):

$$\boldsymbol{x}_3 = (1, 0, ..., 0), \tag{2.71}$$

where the vector contains d-1 zero components. For the other points, we can use the remaining rotations if available, depending on the spacetime dimension d. By doing a rotation in a (d-1)-dimensional spacetime (so that x_3 remains fixed) we can move x_4 to lie in the plane spanned by the first two axes, *i.e.*,

$$\boldsymbol{x}_4 = (X_1^{(4)}, X_2^{(4)}, 0, ..., 0), \qquad (2.72)$$

which has m = 2 degrees of freedom. We iterate this procedure by performing rotations in successively lower-dimensional spaces, giving for instance

$$\boldsymbol{x}_5 = (X_1^{(5)}, X_2^{(5)}, X_3^{(5)}, 0, ..., 0), \qquad (2.73)$$

that has m = 3 degrees of freedom. The *n*th point will be

$$\boldsymbol{x}_n = (X_1^{(n)}, X_2^{(n)}, ..., X_{n-2}^{(n)}, 0, ..., 0).$$
(2.74)

Summing all the degrees of freedom m for each point, we find

$$N_{d,n} = \sum_{m=2}^{n-2} m = \frac{1}{2}n(n-3).$$
(2.75)

Note that we implicitly assumed $d \ge n-2$ so far. When n = 4, for $d \ge 2$ (and in this thesis we are considering $d \ge 3$), this assumption holds. Hence this counting of cross ratios holds and gives, as expected, $N_{d,n} = 2$. On the other hand, if d < n-2 (or equivalently n > d + 2), then all \boldsymbol{x}_k , with $k \ge d + 2$, will have d free parameters since there are no rotational degrees of freedom left to fix them. So we have

$$\boldsymbol{x}_k = (X_1, X_2, ..., X_d), \quad k = d + 3, ..., n.$$
 (2.76)

Hence, the counting of the degrees of freedom becomes

$$N_{d,n} = \left(\sum_{m=2}^{d} m\right) + d\left(n - (d+2)\right) = nd - \frac{1}{2}(d+2)(d+1).$$
(2.77)

Note that the two values for $N_{d,n}$ in equations (2.75) and (2.77) coincide when n = d + 1 or n = d + 2.

Chapter 3

Conformal field theory in momentum space

3.1 Introduction

In this chapter we give an overview of the main results in momentum-space CFT. As a counterpart to the previous chapter, we first derive CWIs in momentum space. Then, we solve them from 2- to general *n*-point. We show that up to three points the symmetry fixes uniquely the form of correlators and discuss their singularities and renormalisation. In particular, we derive the 3-point function in the form of the *triple-K integral*, an integral of three Bessel functions of the second kind and find equivalent representations. Then, we introduce shift operators acting on the 3-point function to shift the parameters. This helps the evaluation of special 3-point functions which would be otherwise difficult. While 3-point functions are well understood, the knowledge of higher-point functions is less complete. We then present the general *n*-point function recently found as a Feynman integral over a (n - 1)-simplex, featuring an arbitrary function of momentum-space cross ratios and conclude with a summary and open questions which will be addressed in the second part of the thesis.

3.2 Conformal Ward identities

In Chapter 2 we presented the scalar conformal Ward identities in position space. To explore the implications of conformal symmetry in momentum space we start with deriving the corresponding momentum-space conformal Ward identities. We obtain a set of differential equations that must be satisfied by conformal correlators. This means that, as in position space, CWIs constrain the form of conformal correlators. Moreover, the theory of differential equations and multivariable hypergeometric functions reveals a description of 3-point functions and certain special 4-point functions [94, 95] in terms of known hypergeometric functions such as Appell F_4 or Lauricella F_C . In Chapter 5 we will also describe some of these solutions as generalised hypergeometric functions introduced by Gelfand, Kapranov, and Zelevinsky (GKZ systems).

The momentum-space dilatation and special conformal Ward identities read

$$0 = D \langle\!\langle \mathcal{O}(\boldsymbol{p}_1) ... \mathcal{O}(\boldsymbol{p}_n) \rangle\!\rangle, \tag{3.1}$$

$$0 = \mathcal{K}^{\mu} \langle\!\langle \mathcal{O}(\boldsymbol{p}_1) ... \mathcal{O}(\boldsymbol{p}_n) \rangle\!\rangle, \tag{3.2}$$

where double brackets denote the *reduced* correlators related to the standard correlator by pulling out the delta function of momentum-conservation:

$$\langle \mathcal{O}_1(\boldsymbol{p}_1)\cdots\mathcal{O}_n(\boldsymbol{p}_n)\rangle = (2\pi)^d \delta(\boldsymbol{p}_1+\cdots+\boldsymbol{p}_n) \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\cdots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle, \tag{3.3}$$

while D and \mathcal{K}^{μ} respectively are the dilatation and the special conformal operators:

$$D = -(n-1)d + \Delta_t - \sum_{j=1}^{n-1} p_j^{\mu} \frac{\partial}{\partial p_j^{\mu}},$$

$$\mathcal{K}^{\mu} = \sum_{j=1}^{n-1} \mathcal{K}_j^{\mu},$$
 (3.4)

with

$$\mathcal{K}_{j}^{\mu} = 2(\Delta_{j} - d)\frac{\partial}{\partial p_{j\mu}} - 2p_{j}^{\nu}\frac{\partial}{\partial p_{j}^{\nu}}\frac{\partial}{\partial p_{j\mu}} + p_{j}^{\mu}\frac{\partial}{\partial p_{j}^{\nu}}\frac{\partial}{\partial p_{j\nu}},\tag{3.5}$$

and $\Delta_t = \sum_{j=1}^n \Delta_j$. To derive the momentum-space CWIs above, we consider the inverse Fourier transform of position-space correlators

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\cdots\mathcal{O}_n(\boldsymbol{x}_n)\rangle = \left(\prod_{j=1}^n \int \frac{\mathrm{d}^d \boldsymbol{p}_j}{(2\pi)^d} \mathrm{e}^{i\boldsymbol{p}_j\cdot\boldsymbol{x}_j}\right) \langle \mathcal{O}_1(\boldsymbol{p}_1)\cdots\mathcal{O}_n(\boldsymbol{p}_n)\rangle.$$
(3.6)

Translational invariance corresponds to pulling out a momentum-conserving delta function. To see this, we take the inverse Fourier transform of

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\cdots\mathcal{O}_n(\boldsymbol{x}_n)\rangle = \langle \mathcal{O}_1(\boldsymbol{x}_1-\boldsymbol{x}_n)\cdots\mathcal{O}_n(\boldsymbol{0})\rangle,$$
 (3.7)

leading to

$$\langle \mathcal{O}_1(\boldsymbol{x}_1)\cdots\mathcal{O}_n(\boldsymbol{x}_n)\rangle = \left(\prod_{j=1}^{n-1}\int \frac{\mathrm{d}^d \boldsymbol{p}_j}{(2\pi)^d} \mathrm{e}^{i\boldsymbol{p}_j\cdot\boldsymbol{x}_{jn}}\right) \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\cdots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle,\tag{3.8}$$

with $\mathbf{x}_{jn} = \mathbf{x}_j - \mathbf{x}_n$. We act with the position-space dilatation and special conformal operators, (2.48) and (2.49), on the right-hand side of (3.8). This amounts to considering the action only on the exponential factor $\exp\left(\sum_{j=1}^{n-1} \mathbf{p}_j \cdot \mathbf{x}_{jn}\right)$. For this purpose, we rewrite the position-space CWIs by eliminating the derivative with respect to x_n^{μ} via the translational Ward identity (2.45),

$$\frac{\partial}{\partial x_n^{\mu}} \rightarrow -\sum_{j=1}^{n-1} \frac{\partial}{\partial x_j^{\mu}}.$$
(3.9)

Thus, the rotation and dilatation Ward identities in position space read

$$0 = \sum_{j=1}^{n-1} \left(x_{jn}^{\mu} \frac{\partial}{\partial_j^{\nu}} - x_{jn}^{\nu} \frac{\partial}{\partial_j^{\mu}} \right) \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle, \qquad (3.10)$$

$$0 = \left(\Delta_t + \sum_{j=1}^{n-1} x_{jn}^{\mu} \frac{\partial}{\partial x_j^{\mu}}\right) \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle.$$
(3.11)

Rearranging the SCWI we have

$$0 = \sum_{j=1}^{n-1} \left\{ \left(x_{jn}^2 \eta^{\mu\nu} - 2x_{jn}^{\mu} x_{jn}^{\nu} \right) \frac{\partial}{\partial x_j^{\mu}} - 2\Delta_j x_{jn}^{\nu} \right\} \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle + 2x_n^{\nu} \sum_{j=1}^{n-1} \left(x_{jn}^{\mu} \frac{\partial}{\partial y_j^{\nu}} - x_{jn}^{\nu} \frac{\partial}{\partial y_j^{\mu}} \right) \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle - 2x_n^{\mu} \left(\Delta_t + \sum_{j=1}^{n-1} x_{jn}^{\mu} \frac{\partial}{\partial x_j^{\mu}} \right) \langle \mathcal{O}_1(\boldsymbol{x}_1) \mathcal{O}_2(\boldsymbol{x}_2) \cdots \mathcal{O}_n(\boldsymbol{x}_n) \rangle,$$
(3.12)

where the last two lines vanish upon (3.10) and (3.11) respectively. After having expressed the position-space CWIs only in terms of \boldsymbol{x}_{jn} and $\partial/\partial x_j^{\mu}$, with j = 1, ..., n-1, the standard Fourier correspondence holds

$$x_{jn}^{\mu} \to -i\frac{\partial}{\partial p_j^{\mu}}, \qquad \frac{\partial}{\partial x_j^{\mu}} \to ip_j^{\mu}, \qquad j = 1, .., n-1.$$
 (3.13)

To find the dilatation and special conformal Ward identities in momentum space, we then act with the operators in (3.11) and (3.12) on the right-hand side of (3.8) and integrate by parts with respect to the momenta. This leads to the CWIs (3.1) and (3.2).

In the following section we will solve 2- and 3-point CWIs directly and find a representation for the general *n*-point function that solves the CWIs. Before moving to the next sections, let us note that we can obtain a set of scalar SCWIs. In fact, in momentum space we have the advantage of decomposing the operator \mathcal{K}^{μ} into a basis of n-1 independent vectors p_{j}^{μ} , with j = 1, ..., n-1:

$$\mathcal{K}^{\mu} = p_1^{\mu} \mathcal{K}_1 + \dots + p_{n-1}^{\mu} \mathcal{K}_{n-1}.$$
(3.14)

Hence, the SCWI (3.2) is equivalent to n-1 scalar equations

$$\mathcal{K}_{j}\langle\!\langle \mathcal{O}(\boldsymbol{p}_{1})...\mathcal{O}(\boldsymbol{p}_{n})\rangle\!\rangle = 0, \qquad j = 1, ..., n - 1.$$
(3.15)

3.3 2-point function

Conformal invariance implies that the 1-point function vanishes, hence the first non trivial correlator is the 2-point function. In this section we solve momentum-space CWIs at two points. The complexity of the set of differential equations to be solved and their solutions

increases with the number of points n.

Translational invariance corresponds to momentum conservation, hence the 2-point correlator depends only on $p \equiv p_1 = -p_2$. Therefore the expansion in (3.14) contains only one term. Moreover, rotational invariance implies that the scalar correlator only depends on scalar quantities, which in this case is the magnitude p of the momentum p. Using the chain rule

$$\frac{\mathrm{d}}{\mathrm{d}p_{\mu}} = \frac{p^{\mu}}{p} \frac{\mathrm{d}}{\mathrm{d}p},\tag{3.16}$$

from (3.1) and (3.15), we obtain the 2-point DWI and SCWI

$$0 = \left(d - \Delta_1 - \Delta_2 + p \frac{\mathrm{d}}{\mathrm{d}p}\right) \langle\!\langle \mathcal{O}(\boldsymbol{p}) \mathcal{O}(-\boldsymbol{p}) \rangle\!\rangle, \qquad (3.17)$$

$$0 = \left(\frac{\mathrm{d}^2}{\mathrm{d}p^2} + \frac{d - 2\Delta_1 + 1}{p} \frac{\mathrm{d}}{\mathrm{d}p}\right) \langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle.$$
(3.18)

We start with solving the SCWI. The general solution of (3.18) reads

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle = c_0 p^{2\Delta_1 - d} + c_1,$$
(3.19)

where c_0 and c_1 are integration constants. We then plug this expression into the dilatation Ward identity (3.17) and find

$$\Delta_1 = \Delta_2 \equiv \Delta, \quad c_1 = 0, \tag{3.20}$$

as expected from the 2-point position-space solution. So the general 2-point conformal correlator is

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle = c_0 p^{2\Delta-d}.$$
 (3.21)

We could have found this solution by Fourier transforming the known position-space solution. Setting $C_{12} = 1$ in (2.53), we find the correspondent solution in momentum space

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle = \int \frac{\mathrm{d}^d \boldsymbol{x}}{(2\pi)^d} \mathrm{e}^{-i\boldsymbol{p}\cdot\boldsymbol{x}} x^{-2\Delta} = \frac{2^{d-2\Delta}\pi^{d/2}\Gamma\left(\frac{d}{2}-\Delta\right)}{\Gamma(\Delta)} p^{2\Delta-d},\tag{3.22}$$

where we wrote $x^{-2\Delta}$ using the Schwinger parametrisation [65]

$$\frac{1}{A^{\nu}} = \frac{1}{\Gamma(\nu)} \int_0^\infty d\lambda \lambda^{\nu-1} e^{-\lambda A}, \qquad (3.23)$$

and performed the resulting Gaussian integral.

3.4 3-point function

In this section we present the 3-point function. First we solve the 3-point conformal Ward identities by separation of variables, giving a scalar representation of the solution. Then, we discuss the uniqueness of this solution by looking at its asymptotic behaviour and the singularities arising from collinear configurations of the momenta. This leads to a unique 3-point function known as the *triple-K integral*. We then present equivalent representations in terms of the generalised hypergeometric function Appell F_4 and of a 1-loop triangle Feynman diagram.

Let us first write the 3-point CWIs in terms of scalar invariants. Poincaré invariance implies that 3-point correlators depend on the scalars formed with p_1 and p_2 : we choose the momenta magnitudes p_j , with j = 1, 2, 3. Let us start with the dilatation Ward identity. Taking into account that $p_3 = -p_1 - p_2$, we find the chain rules

$$\frac{\partial}{\partial p_1^{\mu}} = \frac{p_1^{\mu}}{p_1} \frac{\partial}{\partial p_1} + \frac{p_1^{\mu} + p_2^{\mu}}{p_3} \frac{\partial}{\partial p_3},$$

$$\frac{\partial}{\partial p_2^{\mu}} = \frac{p_2^{\mu}}{p_2} \frac{\partial}{\partial p_2} + \frac{p_1^{\mu} + p_2^{\mu}}{p_3} \frac{\partial}{\partial p_3}.$$
(3.24)

Hence, we can write the dilatation Ward identity in terms of scalar variables:

$$0 = D \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2 \boldsymbol{p}_2 \mathcal{O}_3 \boldsymbol{p}_3) \rangle\!\rangle = \left(2d - \Delta_t + \sum_{j=1}^3 p_j \frac{\partial}{\partial p_j} \right) \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2 \boldsymbol{p}_2 \mathcal{O}_3 \boldsymbol{p}_3) \rangle\!\rangle.$$
(3.25)

As noted earlier, this equation tells us that its solution is a homogeneous function of degree $\Delta_t - 2d$. This means that we can write the solution in the following way

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2\boldsymbol{p}_2\mathcal{O}_3\boldsymbol{p}_3)\rangle\!\rangle = p_3^{\Delta_t - 2d} F\left(\frac{p_1}{p_3}, \frac{p_2}{p_3}\right),\tag{3.26}$$

where here F is a general function. To determine its explicit expression we need the special conformal Ward identity

$$\mathcal{K}^{\mu} \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2 \boldsymbol{p}_2 \mathcal{O}_3 \boldsymbol{p}_3) \rangle\!\rangle = 0, \qquad (3.27)$$

where

$$\mathcal{K}^{\mu} = p_1^{\mu} \mathcal{K}_1 + p_2^{\mu} \mathcal{K}_2. \tag{3.28}$$

Therefore the SCWI (3.27) is equivalent to the system formed by the following two scalar equations

$$\mathcal{K}_j \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2(\boldsymbol{p}_2) \mathcal{O}_3(\boldsymbol{p}_3) \rangle\!\rangle = 0, \quad j = 1, 2.$$
(3.29)

Using the chain rule (3.24) we find the operators \mathcal{K}_j . The explicit expression of \mathcal{K}_1 is

$$\mathcal{K}_{1} = \frac{\partial^{2}}{\partial p_{1}^{2}} + \frac{\partial^{2}}{\partial p_{3}^{2}} + \frac{2p_{1}}{p_{3}} \frac{\partial^{2}}{\partial p_{1} \partial p_{3}} + \frac{2p_{2}}{p_{3}} \frac{\partial^{2}}{\partial p_{2} \partial p_{3}} - \frac{2\Delta_{1} - d - 1}{p_{1}} \frac{\partial}{\partial p_{1}} - \frac{2\Delta_{1} + 2\Delta_{2} - 3d - 1}{p_{3}} \frac{\partial}{\partial p_{3}}, \qquad (3.30)$$

and \mathcal{K}_2 can be obtained by permuting $1 \leftrightarrow 2$ indices. Since these operators contain mixed derivatives, we eliminate the mixed terms by invoking the DWI (3.25). We define

$$K_{13} = \mathcal{K}_1 - \frac{2}{p_3} \frac{\partial}{\partial p_3} D, \qquad K_{23} = \mathcal{K}_2 - \frac{2}{p_3} \frac{\partial}{\partial p_3} D, \qquad (3.31)$$
hence the set of equations (3.25) amounts to the following simpler version

$$K_{j3} \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2(\boldsymbol{p}_2) \mathcal{O}_3(\boldsymbol{p}_3) \rangle\!\rangle = 0, \quad j = 1, 2,$$

$$(3.32)$$

where

$$K_{ij} = K_i - K_j,$$

$$K_i \equiv \frac{\partial^2}{\partial p_i^2} - \frac{2\Delta_i - d - 1}{p_i} \frac{\partial}{\partial_i}, \qquad i \neq j = 1, 2, 3.$$
(3.33)

Note that the operator K_i appeared already in the 2-point special conformal Ward identity in equation (3.18).

In the following we find the general solution of the CWIs that – to summarise – we wrote as the following set of equations

$$D\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = 0,$$

$$K_{13}\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = 0,$$

$$K_{23}\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = 0.$$
(3.34)

Before showing how to solve these, we give the general solution known as the triple-K integral:

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = c_{123} \int_0^\infty \mathrm{d}x \, x^\alpha \prod_{j=1}^3 p_j^{\beta_j} K_{\beta_j}(p_j x),$$
 (3.35)

where c_{123} is an integration constant and K_{β_j} is the modified Bessel function of the second kind, also known as Bessel-K function [96]. The parameters α and β_j are related to the physical spacetime dimension and the scaling dimensions as follows:

$$\alpha = \frac{d}{2} - 1,$$
 $\beta_j = \Delta_j - \frac{d}{2},$ $j = 1, 2, 3.$ (3.36)

In the following we will denote the triple-K integral (3.35) with $I_{\alpha\{\beta_1\beta_2\beta_3\}}$.

Let us present the derivation of the general solution. We denote the general solution as f. Taking into account the definitions in (3.33), we can write the last two equations of (3.34) as

$$K_1 f = K_2 f = K_3 f. (3.37)$$

Therefore we use separation of variables

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = f(p_1)f(p_2)f(p_3), \qquad (3.38)$$

leading to

$$\frac{K_1 f_1}{f_1} = \frac{K_2 f_2}{f_2} = \frac{K_3 f_3}{f_3} = x^2,$$
(3.39)

where x^2 is a constant. Then, we have to solve the equation

$$\mathbf{K}_{i} f_{i} = \left(\frac{\partial^{2}}{\partial p_{i}} + \frac{1 - 2\beta_{i}}{p_{i}}\frac{\partial}{\partial p_{i}}\right) f_{i} = x^{2} f_{i}, \qquad (3.40)$$

which is equivalent to the modified Bessel's equation,

$$\left[p_i^2 \partial_i^2 + p_i \partial_i - (p_i^2 x^2 + \beta^2)\right] u_i = 0, \qquad (3.41)$$

with $f_i = p_i^{\beta_i} u_i$ and $\partial_i = \partial/\partial p_i$. Hence, the general solution of (3.40) is a linear combination of Bessel-K and Bessel-I functions:

$$f_i(p_i) = p_i^{\beta_i} \left[a_K K_{\beta_i}(p_i x) + a_I I_{\beta_i}(p_i x) \right], \qquad (3.42)$$

and $f = \prod_i f_i(p_i)$ solves the SCWIs. By linearity, indeed, any integral $\int dx \rho(x) f$ over a spectral function $\rho(x)$ solves the CWIs. The dilatation Ward identity fixes the form of this function. Noting that

$$\sum_{i=1}^{3} p_i \partial_i f(p_1 x, p_2 x, p_3 x) = x \partial_x f(p_1 x, p_2 x, p_3 x), \qquad (3.43)$$

we have

$$\int_{0}^{\infty} \mathrm{d}x\rho(x)x\partial_{x}f(p_{1}x, p_{2}x, p_{3}x) = (\beta_{t} - \alpha - 1)\int_{0}^{\infty} \mathrm{d}x\rho(x)f(p_{1}x, p_{2}x, p_{3}x), \qquad (3.44)$$

where we defined $\beta_t = \beta_1 + \beta_2 + \beta_3$. Integrating by parts we find $\rho(x) = x^{\alpha - \beta_t}$. Therefore we conclude that if $f(p_1, p_2, p_3)$ is the solution of the SCWIs, then

$$I = \int_0^\infty \mathrm{d}x x^{\alpha - \beta_t} f(p_1 x, p_2 x, p_3 x)$$
(3.45)

solves the DWI. The solution we found is not exactly the triple-K integral in (3.35), since the function f involves also Bessel-I functions. A discussion based on physical properties of the solution restricts I to the final form of a triple-K integral, as we explain in the following section.

3.4.1 Collinear singularities

For (3.45) to converge, at least one of the $f_i(p_i)$ needs to be a Bessel-K function. Let us assume $f_3 \sim K_{\beta_3}$, then the conformal Ward identities admit four independent solutions that schematically we can refer to as *IIK*, *KIK*, *IKK*, *KKK*. In this section we show that only one of them is physically acceptable, leading to the triple-K integral (3.35). This conclusion agrees with the result in position-space, where the 3-point function is unique.

To see this, consider the behaviour of modified Bessel functions when their argument is large,

$$I_{\nu}(x) = \frac{1}{\sqrt{2\pi}} \frac{e^x}{\sqrt{x}} + \dots, \qquad K_{\nu}(x) = \sqrt{\frac{\pi}{2}} \frac{e^{-x}}{\sqrt{x}} + \dots, \qquad x \to \infty.$$
(3.46)

Since the Bessel-I function diverges for large x, as we anticipated, the solution must have at least one Bessel-K function in the integrand. Now, let us consider the case where the integrand is of the form *IIK*. According to the asymptotic behaviours (3.46), in order for the integral to converge, the following condition must hold

$$p_3 \ge p_1 + p_2. \tag{3.47}$$

This violates the triangle inequality $(p_3 \leq p_1 + p_2)$, therefore we exclude solutions of the form *IIK*. If we consider the integrand with only one Bessel-I (*IKK*), we find that this is also forbidden since it is singular for collinear momentum configurations. In fact, at large x the integral would approximately be

$$I \sim \int \mathrm{d}x x^{\alpha - \frac{3}{2}} \mathrm{e}^{(p_1 - p_2 - p_3)x}.$$
 (3.48)

While the triangle inequality is respected, if the momenta are collinear, *i.e.*, $p_1 = p_2 + p_3$, the integral diverges when $\alpha \ge 1/2$, hence $d \ge 3$. Therefore we find that only a unique 3-point function exists, given by the triple-K integral

$$I_{\alpha\{\beta_1\beta_2\beta_3\}} = c_{123} \int_0^\infty \mathrm{d}x x^\alpha \prod_{j=1}^3 p_j^{\beta_j} K_{\beta_j}(p_j x).$$
(3.49)

In the following, we present some explicit examples of triple-K integrals, for special values of the parameters.

3.4.2 Examples

In this section we consider two examples. When the β_i are half-integer, the integral is given by elementary functions, while for integer β_i , the triple-K can be expressed in terms of the Bloch-Wigner function.

1. d = 4, $\Delta_1 = \Delta_2 = \Delta_3 = 5/2$. In this case all β_i are half integers:

 $\alpha = \frac{d}{2} - 1 = 1, \qquad \beta_i = \Delta_i - \frac{d}{2} = \frac{1}{2},$ (3.50)

and the 3-point correlator reads

$$\langle\!\langle O_1(\boldsymbol{p}_1)O_2(\boldsymbol{p}_2)O_3(\boldsymbol{p}_3)\rangle\!\rangle = c(p_1p_2p_3)^{\frac{1}{2}} \int_0^\infty \mathrm{d}x \ xK_{\frac{1}{2}}(p_1x)K_{\frac{1}{2}}(p_2x)K_{\frac{1}{2}}(p_3x).$$
(3.51)

The Bessel-K function in the integrand is

$$K_{\frac{1}{2}}(x) = \sqrt{\frac{\pi}{2}} \frac{\mathrm{e}^{-x}}{x^{1/2}},\tag{3.52}$$

and the integral is convergent giving

$$\langle\!\langle O_1(\boldsymbol{p}_1)O_2(\boldsymbol{p}_2)O_3(\boldsymbol{p}_3)\rangle\!\rangle = \frac{c\pi^2}{2^{\frac{3}{2}}} \frac{1}{\sqrt{p_1 + p_2 + p_3}}.$$
 (3.53)

2. d = 4, $\Delta_1 = \Delta_2 = \Delta_3 = 2$.

In this case all β_i are integers

$$\alpha = \frac{d}{2} - 1 = 1, \qquad \beta_i = \Delta_i - \frac{d}{2} = 0,$$
(3.54)

and we denote the 3-point function as

$$\langle\!\langle O_1(\boldsymbol{p}_1)O_2(\boldsymbol{p}_2)O_3(\boldsymbol{p}_3)\rangle\!\rangle = cI_{1\{000\}},$$
(3.55)

where the integral $I_{1\{000\}}$ is a known integral in literature [97, 98] and is expressed in terms of the dilogarithm function Li₂ [99]. We show the explicit computation of $I_{1\{000\}}$ in appendix A and quote here the final result:

$$I_{1\{000\}} = \frac{1}{2\sqrt{-J^2}} \left[\text{Li}_2(\bar{z}) - \text{Li}_2(z) - \frac{1}{2}\ln(z\bar{z})\ln\left(\frac{1-z}{1-\bar{z}}\right) \right], \quad (3.56)$$

where the z-variables are related to the momenta magnitudes by

$$z\bar{z} = \frac{p_1^2}{p_3^2}, \qquad (1-z)(1-\bar{z}) = \frac{p_2^2}{p_3^2},$$
 (3.57)

or equivalently

$$z = \frac{1}{2p_3^2} \left(p_1^2 - p_2^2 + p_3^2 + \sqrt{-J^2} \right), \qquad \bar{z} = \frac{1}{2p_3^2} \left(p_1^2 - p_2^2 + p_3^2 - \sqrt{-J^2} \right), \qquad (3.58)$$

where we defined

$$J^{2} = (p_{1} + p_{2} - p_{3})(p_{1} - p_{2} + p_{3})(-p_{1} + p_{2} + p_{3})(p_{1} + p_{2} + p_{3})$$

= $-(z - \bar{z})^{2}p_{3}^{4}$. (3.59)

Note that J^2 is related to the Gram determinant of p_1 and p_2 :

$$J^2 = 4G(p_1, p_2), (3.60)$$

and hence $\sqrt{J^2}$ is proportional to the area of the triangle $A_T(p_1, p_2, p_3)$ whose sides are p_1, p_2, p_3 . The solution $I_{1\{000\}}$ is related to the Bloch-Wigner function D(z) in the following way

$$I_{1\{000\}} = \frac{D(z)}{\sqrt{-J^2}} = \frac{\text{Vol}(\Delta)}{4A_{\text{T}}(p_1, p_2, p_3)}.$$
(3.61)

Finally, let us note that the integral $I_{1\{000\}}$ is an example of a master integral: we can obtain all triple-K integrals with integral β_i from it by acting with shift operators, as we will explain later in the thesis.

Below we summarise the dependence of the triple-K integral on β_i :

- all β_i half-integral $\Rightarrow I_{\alpha\{\beta_1\beta_2\beta_3\}}$ in terms of elementary functions of momentum magnitudes
- all β_i integral $\Rightarrow I_{\alpha\{\beta_1\beta_2\beta_3\}}$ in terms of dilogarithms.

3.4.3 The 3-point function as a hypergeometric system

In this section we show an alternative way of solving the conformal Ward identities with the aim of stressing the connection between CWIs and hypergeometric systems. It has been shown independently in [43] and [100] that the solution of 3-point CWIs can be expressed in terms of the generalised hypergeometric function of two variables Appell F_4 [101, 102]. In fact, the system of dilatation and special conformal Ward identities is equivalent to the set of differential equations defining Appell F_4 function:

$$0 = \left[\xi(1-\xi)\frac{\partial^2}{\partial\xi^2} - \eta^2\frac{\partial^2}{\partial\eta^2} - 2\xi\eta\frac{\partial^2}{\partial\xi\partial\eta} + \left(\gamma - (\alpha+\beta+1)\xi\right)\frac{\partial}{\partial\xi} - (\alpha+\beta+1)\eta\frac{\partial}{\partial\eta} - \alpha\beta\right]F(\xi,\eta), \quad (3.62)$$

$$0 = \left[\eta(1-\eta)\frac{\partial^2}{\partial\eta^2} - \xi^2 \frac{\partial^2}{\partial\xi^2} - 2\xi\eta \frac{\partial^2}{\partial\xi\partial\eta} + \left(\gamma' - (\alpha+\beta+1)\eta\right)\frac{\partial}{\partial\eta} - (\alpha+\beta+1)\xi\frac{\partial}{\partial\xi} - \alpha\beta\right]F(\xi,\eta).$$
(3.63)

where

$$\xi = \frac{p_1^2}{p_3^2}, \qquad \eta = \frac{p_2^2}{p_3^2}, \tag{3.64}$$

and $\alpha, \beta, \gamma, \gamma'$ are some linear combinations of the spacetime d and the scaling dimensions Δ_i . Amongst the well known properties of the Appell F_4 system, let us note that it admits four linearly independent solutions:

$$F_{4}(\alpha,\beta;\gamma,\gamma';\xi,\eta), \xi^{1-\gamma}F_{4}(\alpha+1-\gamma,\beta+1-\gamma;2-\gamma,\gamma';\xi,\eta), \eta^{1-\gamma'}F_{4}(\alpha+1-\gamma',\beta+1-\gamma';\gamma,2-\gamma';\xi,\eta), \xi^{1-\gamma}\eta^{1-\gamma'}F_{4}(\alpha+2-\gamma-\gamma',\beta+2-\gamma-\gamma';2-\gamma,2-\gamma';\xi\eta).$$
(3.65)

In the following we briefly explain how to show the equivalence between the Appell F_4 system and the conformal Ward identities. First, we recall that the dilatation Ward identity allows us to write the following ansatz for the general solution:

$$\langle\!\langle O(p_1)O(p_2)O(p_3)\rangle\!\rangle = p_3^{\Delta_t - 2d} \left(\frac{p_1^2}{p_3^2}\right)^\mu \left(\frac{p_2^2}{p_3^2}\right)^\lambda F\left(\frac{p_1^2}{p_3^2}, \frac{p_2^2}{p_3^2}\right), \tag{3.66}$$

with λ and μ arbitrary parameters. We also use the dilatation Ward identity (3.25) to eliminate the derivatives with respect to p_3 appearing in the special conformal Ward identities:

$$\frac{\partial}{\partial p_3} \to \frac{1}{p_3} \left(\Delta_t - 2d - p_1 \frac{\partial}{\partial p_1} - p_2 \frac{\partial}{\partial p_2} \right). \tag{3.67}$$

Then, by using the chain rule to rewrite the special conformal Ward identities in terms of ξ , η and their derivatives, we find that these equations coincide with those defining the

Appell F_4 system, when

$$\mu = \frac{1}{2} \left(\Delta_1 - \frac{d}{2} \right) (\epsilon_1 + 1), \qquad \lambda = \frac{1}{2} \left(\Delta_2 - \frac{d}{2} \right) (\epsilon_2 + 1), \tag{3.68}$$

with $\epsilon_1, \epsilon_2 \in \{-1, +1\}$. And

$$\alpha = \frac{1}{2} \left[\epsilon_1 \left(\Delta_1 - \frac{d}{2} \right) + \epsilon_2 \left(\Delta_2 - \frac{d}{2} \right) + \Delta_3 \right],$$

$$\beta = \alpha - \left(\Delta_3 - \frac{d}{2} \right),$$

$$\gamma = 1 + \epsilon_1 \left(\Delta_1 - \frac{d}{2} \right),$$

$$\gamma' = 1 + \epsilon_2 \left(\Delta_2 - \frac{d}{2} \right).$$
(3.69)

As showed in the previous section, the 3-point function is unique. Here, again, to avoid collinear singularities, only a particular linear combination of the four linearly independent solutions (3.65) is acceptable. In fact, Appell F_4 has the integral representation [103]

$$F_4\left(\alpha,\beta;\gamma,\gamma';\frac{p_1^2}{p_3^2},\frac{p_2^2}{p_3^2}\right) = \frac{\Gamma(\gamma)\Gamma(\gamma')}{2^{\alpha+\beta-\gamma-\gamma'}\Gamma(\alpha)\Gamma(\beta)} \cdot \frac{p_3^{\alpha+\beta}}{p_1^{\gamma-1}p_2^{\gamma'-1}} \times \int_0^\infty \mathrm{d}x x^{\alpha+\beta-\gamma-\gamma'+1} I_{\gamma-1}(p_1x) I_{\gamma'-1}(p_2x) K_{\beta-\alpha}(p_3x), \quad (3.70)$$

and taking into account (3.69), the four solutions read

$$p_1^{\Delta_1 - \frac{d}{2}} p_2^{\Delta_2 - \frac{d}{2}} p_3^{\Delta_3 - \frac{d}{2}} \int_0^\infty \mathrm{d}x \, x^{\frac{d}{2} - 1} I_{\pm(\Delta_1 - \frac{d}{2})}(p_1 x) I_{\pm(\Delta_2 - \frac{d}{2})}(p_2 x) K_{\Delta_3 - \frac{d}{2}}(p_3 x). \tag{3.71}$$

We explained earlier that `IIK' integrals have singularities for collinear configurations of the momenta. However, since

$$K_{\nu}(x) = \frac{\pi}{2\sin(\nu\pi)} \left[I_{\nu}(x) - I_{-\nu}(x) \right], \qquad (3.72)$$

we recover the triple-K solution by taking the following linear combination of the integrals in (3.71):

$$\int_{0}^{\infty} \mathrm{d}x \, x^{\alpha - 1} K_{\beta_{1}}(p_{1}x) K_{\beta_{2}}(p_{2}x) K_{\beta_{3}}(p_{3}x) = \frac{2^{\alpha - 4}}{c^{\alpha}} \left[A(\beta_{1}, \beta_{2}) + A(\beta_{1}, -\beta_{2}) + A(-\beta_{1}, \beta_{2}) + A(-\beta_{1}, -\beta_{2}) \right],$$
(3.73)

where

$$A(\beta_1,\beta_2) = \left(\frac{p_1}{p_3}\right)^{\beta_1} \left(\frac{p_2}{p_3}\right)^{\beta_2} \Gamma\left(\frac{\alpha+\beta_1+\beta_2-\beta_3}{2}\right) \Gamma\left(\frac{\alpha+\beta_1+\beta_2+\beta_3}{2}\right) \Gamma(-\beta_1)\Gamma(-\beta_2)$$



Figure 3.1: The 1-loop massless triangle graph (3.75).



Figure 3.2: Equivalent electrical networks of resistors under star-mesh duality, where the resistances are related as given in (3.84). The external currents flowing into the corresponding dotted nodes and the overall power dissipation are equal.

$$\times F_4\left(\frac{\alpha+\beta_1+\beta_2-\beta_3}{2},\frac{\alpha+\beta_1+\beta_2+\beta_3}{2};\beta_1+1,\beta_2+1;\frac{p_1^2}{p_3^2},\frac{p_2^2}{p_3^2}\right). (3.74)$$

3.4.4 The triple-K integral as a triangle Feynman integral

In this section we introduce a further representation of the 3-point function: we show the equivalence between the triple-K integral and the 1-loop triangle Feynman integral (see fig. 3.1

$$I_{d\{\alpha_{12},\alpha_{13},\alpha_{23}\}} = \int \frac{\mathrm{d}^{d}\boldsymbol{q}}{(2\pi)^{d}} \frac{1}{|\boldsymbol{q}|^{2\alpha_{12}+d}|\boldsymbol{q}-\boldsymbol{p}_{1}|^{2\alpha_{13}+d}|\boldsymbol{q}+\boldsymbol{p}_{2}|^{2\alpha_{23}+d}}.$$
(3.75)

The relation between these two representations is

$$I_{\alpha\{\beta_1\beta_2\beta_3\}} = C_T I_{d\{\alpha_{12},\alpha_{13},\alpha_{23}\}}$$
(3.76)

with

$$C_T = c_{123} 2^{\frac{3}{2}d-4} \pi^{d/2} \Gamma\left(\frac{\Delta_t - d}{2}\right) \prod_{1 \le j < k \le 3} \Gamma\left(\alpha_{jk} + \frac{d}{2}\right).$$
(3.77)

To show that equation (3.76) holds, here we will manipulate both integrals. The reader can find how to derive the triangle integral starting from the triple-K (and vice versa) in [46] and [42]. The equivalence between these representations is the star-mesh equivalence of electrical circuits in disguise. To see this, we will Schwinger parametrise both representations and perform a star-mesh transformation between the respective Schwinger parameters u_{ij} and Z_j . The triangle integral can be thought as the 'mesh' circuit with resistances u_{ij} , while the triple-K integral corresponds to the 'star' circuit with resistances $1/Z_j$ (see fig. 3.2) with the external momenta p_j the ingoing currents. Later in this thesis we will see how the connection between electrical networks and integral representations of conformal correlation functions can be used to find new scalar representations of the general *n*-point function.

Let us consider the triple-K integral (3.35). We re-express the Bessel-K functions using the integral formula [96]

$$K_{\nu}(z) = \frac{1}{2} \left(\frac{1}{2}z\right)^{\nu} \int_{0}^{\infty} \mathrm{d}t \, t^{-\nu-1} \exp\left(-t - \frac{z^{2}}{4t}\right), \tag{3.78}$$

giving

$$p_j^{\beta_j} K_{\beta_j}(p_j x) = \frac{1}{2} \left(\frac{x}{2}\right)^{\beta_j} \int_0^\infty \mathrm{d}Z_j \, Z_j^{\beta_j - 1} \exp\left(-\frac{p_j^2}{Z_j} - x^2 \frac{Z_j}{4}\right),\tag{3.79}$$

where $Z_j = p_j^2/t$. We set $z = x^2/4$ and perform the integral over z using (3.23), leading to

$$I_{\alpha\{\beta_{1}\beta_{2}\beta_{3}\}} = c_{123}2^{\alpha-1}\Gamma\left(\frac{\Delta_{t}-d}{2}\right)\left(\prod_{j=1}^{3}\int_{0}^{\infty} \mathrm{d}Z_{j} Z_{j}^{\beta_{j}-1}\right)Z_{t}^{\frac{d-\Delta_{t}}{2}}\exp\left(-\sum_{j}\frac{p_{j}^{2}}{Z_{j}}\right), \quad (3.80)$$

where we defined $Z_t = Z_1 + Z_2 + Z_3$. The Z_j variables can be interpreted as the conductivities of the 'star' network and the argument of the exponential can be seen as the power dissipated in the same circuit.

Let us now work on the triangle Feynman integral (3.75). We Schwinger-parametrise the factors in the denominator by using equation (3.23)

$$I_{d\{\alpha_{12}\alpha_{13}\alpha_{23}\}} = C\left(\prod_{1 \le j < k \le 3} \int_0^\infty \mathrm{d}u_{jk} \, u_{jk}^{\alpha_{jk} + \frac{d}{2} - 1}\right) U^{-\frac{d}{2}} \exp\left(-\frac{F}{U}\right),\tag{3.81}$$

with

$$C = (4\pi)^{-\frac{d}{2}} \prod_{1 \le j < k \le 3} \Gamma\left(\alpha_{jk} + \frac{d}{2}\right),$$
(3.82)

and U and F the Symanzik polynomials

$$U = u_{12} + u_{13} + u_{23}, \qquad F = p_1^2 u_{12} u_{13} + p_2^2 u_{12} u_{23} + p_3^2 u_{13} u_{23}. \tag{3.83}$$

Finally, we diagonalise the exponential by performing the 'star-mesh' change of variables. In fact, the Schwinger parameters u_{jk} can be interpreted as the resistances of the triangle (or mesh) circuit which are related to the conductivities Z_i of the star circuit via

$$\frac{u_{ik}u_{jk}}{u_t} = \frac{1}{Z_k}, \qquad i, j, k = 1, 2, 3, \tag{3.84}$$

giving

$$I_{d\{\alpha_{12}\alpha_{13}\alpha_{23}\}} = C\left(\prod_{j=1}^{3} \int_{0}^{\infty} \mathrm{d}Z_{j} Z_{j}^{(\alpha_{ik}-\alpha_{t}-\frac{d}{2})-1}\right) Z_{t}^{\alpha_{t}+\frac{d}{2}} \exp\left(-\sum_{j} \frac{p_{j}^{2}}{Z_{j}}\right), \qquad (3.85)$$

with $\alpha_t = \alpha_{12} + \alpha_{13} + \alpha_{23}$. It is evident that this integral is the same as in equation (3.80) when

$$\alpha_{jk} = -\beta_j - \beta_k + \frac{1}{2} \left(\beta_t - \frac{d}{2} \right), \qquad (3.86)$$

or equivalently

$$\beta_i = -\frac{d}{2} - \sum_{j \neq i} \alpha_{ij}.$$
(3.87)

Hence, we showed the equivalence in (3.76).

3.5 Singularities and renormalisation

So far we have presented 2- and 3-point functions. Before moving to higher-point functions we discuss when divergences arise and briefly explain how to renormalise the solutions in these cases. We will see that renormalisation is necessary when the solutions are analytic functions of the squared momenta p_i^2 , since this corresponds to local solutions in position-space.

3.5.1 2-point function

In section 3.3 we found the solution of the 2-point CWIs, but we have not discussed its domain of existence. Let us recall the Fourier transformed 2-point solution (3.22):

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle = c_0 p^{2\left(\Delta - \frac{d}{2}\right)}, \qquad c_0 = c \frac{2^{d-2\Delta} \pi^{d/2} \Gamma\left(\frac{d}{2} - \Delta\right)}{\Gamma(\Delta)}.$$
 (3.88)

For finite c_0 and generic Δ this solves the CWIs. However, for values of d and Δ such that

$$\frac{d}{2} - \Delta = -n, \qquad n \in \mathbb{Z}^+, \tag{3.89}$$

the 2-point function is divergent and needs to be regulated. One way is to dimensionally regulate the correlator by shifting d and Δ as follows

$$\frac{d}{2} - \Delta = -n - \epsilon, \qquad \epsilon \ll 1. \tag{3.90}$$

Then, expanding in ϵ , the regulated correlator reads

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle_{\text{reg}} = \frac{c_0^{(-1)}}{\epsilon} p^{2n} + c_0^{(-1)} p^{2n} \ln p^2 + c_0^{(0)} p^{2n} + O(\epsilon), \qquad (3.91)$$

which has a pole in ϵ in the limit $\epsilon \to 0$. Can we cancel this singularity? We cannot rescale $c_0 \to \epsilon c_0$ since when the condition (3.89) holds, the 2-point function is an analytic

function of p^2 :

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle = c_0 p^{2n}.$$
 (3.92)

This corresponds to a *local* solution, which means it has support only at $x^2 = 0$ as one can see from its Fourier transform: this is given by 2n derivatives acting on a delta function

$$\langle \mathcal{O}(\boldsymbol{x})\mathcal{O}(0)\rangle = c_0(-\Box)^n \delta(\boldsymbol{x}). \tag{3.93}$$

This is not a physically acceptable solution and, as we showed in Chapter 2, all positionspace correlators are non-local. Moreover, when condition (3.89) holds, there is a new local term in the action of the form

$$\phi \Box^n \phi, \tag{3.94}$$

where ϕ is the source of the operator \mathcal{O} . With an appropriate choice of the coefficient, this can be treated as a counterterm that cancels the divergence of the correlator.

The contribution to the momentum-space correlator of this counterterm is

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle_{\rm ct} = \tilde{c}(\epsilon)p^{2n}\mu^{2\epsilon} = \frac{\tilde{c}^{(-1)}}{\epsilon}p^{2n} + \tilde{c}^{(-1)}p^{2n}\ln\mu^2 + \tilde{c}^{(0)}p^{2n} + O(\epsilon), \qquad (3.95)$$

where the renormalisation group (RG) scale μ appeared on dimensional grounds. If we then sum the two contribution to the correlator, (3.91) and (3.95), we can cancel the divergence by setting $\tilde{c}^{(-1)} = -c_0^{(-1)}$ and take the limit $\epsilon \to 0$, leading to the renormalised 2-point function:

$$\langle\!\langle \mathcal{O}(\boldsymbol{p})\mathcal{O}(-\boldsymbol{p})\rangle\!\rangle_{\mathrm{ren}} = p^{2n} \left(c_0^{(-1)} \ln \frac{p^2}{\mu^2} + c_0' \right).$$
 (3.96)

This solution depends on the renormalisation scale μ which breaks conformal symmetry resulting in a conformal anomaly. This has been studied in more detail in [45] and we will not need it in this thesis.

3.5.2 3-point function

In section 3.4 we found the general solution of the 3-point CWIs and discussed its uniqueness due to the absence of collinear singularities, leading to the triple-K integral (3.35). This converges at large x, according to the expansion in (3.46). However, there could be divergences in the limit of $x \to 0$. The condition for this to happen is

$$\alpha + 1 \pm \beta_1 \pm \beta_2 \pm \beta_3 = -2n, \quad n \in \mathbb{Z}^+.$$
(3.97)

To see this, we consider the series expansion of the Bessel-K function around x = 0. Let β_j be a non-integer number. Then, taking into account the series expansion of the Bessel-I, $I_{\beta} = \sum_{n=0}^{\infty} a_I x^{\beta+2n}$, and equation (3.72), the expansion of Bessel-K is

$$K_{\beta_j}(x) = x^{\beta_j} \sum_{n_j=0}^{\infty} a_+ x^{2n_j} + x^{-\beta_j} \sum_{n_j=0}^{\infty} a_- x^{2n_j}.$$
(3.98)

We use this equation to expand the integrand of the triple-K (3.35), leading to integrals of the form

$$\int_{0}^{\infty} dx \, x^{\eta} = \frac{x^{\eta+1}}{\eta+1}, \qquad \eta = \alpha \pm \beta_1 \pm \beta_2 \pm \beta_3 + 2n_t, \tag{3.99}$$

with $n_t = n_1 + n_2 + n_3$. This integral converges at x = 0 when

$$\alpha \pm \beta_1 \pm \beta_2 \pm \beta_3 > -1. \tag{3.100}$$

However, if we consider the integral as a function of its parameters while the momenta are fixed, we can perform an analytic continuation by considering the complex η -plane. Here the integral in (3.99) is well defined unless $\eta = -1$, where it has a pole. This gives the singularity condition in (3.97). For integer β_j , the series expansion of Bessel-K functions around x = 0 is

$$K_{\beta_j}(x) = x^{\beta_j} \sum_{n_j=0}^{\infty} a_+ x^{2n_j} + x^{-\beta_j} \sum_{n_j=0}^{\infty} a_- x^{2n_j} + c_I \log(x) I_{\beta_j}(x).$$
(3.101)

Then the corresponding expansion of the integrand in the triple-K contains also logarithms,

$$\int_0^\infty \mathrm{d}x \, x^\eta \log(x) = -\frac{\mathrm{const}_1}{(\eta+1)^2} + \frac{\mathrm{const}_2}{(\eta+1)}.$$
(3.102)

This, however, does not change the loci of singularities $\eta = -1$, it may only increase the order of the pole.

When equation (3.97) holds, we need to regulate the triple-K integral. We dimensional regulate it by shifting the parameters as follows

$$\alpha \to \tilde{\alpha} = \alpha + u\epsilon, \qquad \beta_i \to \tilde{\beta}_i = \beta_i + v\epsilon, \quad i = 1, 2, 3.$$
 (3.103)

Note that this leaves the form of the triple-K unchanged, but the parameters are now $\tilde{\alpha}$ and $\tilde{\beta}_i$ and the constant c_{123} depends on the regularisation parameters ϵ, u and v. When the singularity condition (3.97) holds, we find that the regulated solution is singular when we take the limit $\epsilon \to 0$. Depending on the type of singularity, this is canceled either by renormalisation or by an appropriate choice of the constant c_{123} .

Four different singularity conditions arise from (3.97), depending on the relative signs of the β_i . These are the (---), (--+), (-++), (+++) conditions. The first two singularities, *i.e.*, (---) and (--+), correspond respectively to ultralocal and semilocal solutions. By ultralocal we mean that the solution only has support on configurations where three positions collapse to one single point. By semilocal, we refer to solutions that have support where two of the three positions coincide at one point. In these cases, counterterms exist and divergences are canceled via renormalisation. The latter leads to a conformal anomaly for the condition (--), while the condition (--+) corresponds to beta functions¹.

On the other hand, the other two conditions with mostly '+' signs, correspond to non-local solutions. Consequently, counterterms do not exist and the singularities are just

¹Note, this is not in contradiction with conformal symmetry, since this type of beta functions are associated to couplings of composite operators which are not couplings appearing in the Lagrangian of fundamental fields.

singularities of the triple-K integral that can be cured by choosing the constant c_{123} to be proportional to ϵ .

Finally, note that more than one condition can be satisfied simultaneously, resulting in higher-order singularities.

In the next section we illustrate some examples where the condition (3.97) holds.

3.5.3 Examples

1. (+++) condition: d = 3, $\Delta_1 = \Delta_2 = \Delta_3 = 1$ ($\alpha = 1/2$ and $\beta_i = -1/2$, with i = 1, 2, 3).

Let us consider the following regularisation scheme

$$d \to d + 2\epsilon, \qquad \Delta_i \to \Delta_i + \epsilon,$$
 (3.104)

so that the indices β_i (3.36) of Bessel-K functions don't change. Then, the triple-K integral reads

$$\langle\!\langle O_1(\boldsymbol{p}_1)O_2(\boldsymbol{p}_2)O_3(\boldsymbol{p}_3)\rangle\!\rangle = c(p_1p_2p_3)^{-\frac{1}{2}} \int_0^\infty \mathrm{d}x \, x^{\frac{1}{2}+\epsilon} K_{\frac{1}{2}}(p_1x) K_{\frac{1}{2}}(p_2x) K_{\frac{1}{2}}(p_3x),$$
(3.105)

where the Bessel-K are elementary functions,

$$K_{\frac{1}{2}}(x) = \sqrt{\frac{\pi}{2}} \frac{\mathrm{e}^{-x}}{x^{1/2}}.$$
(3.106)

Then, the integral evaluates to

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = \frac{c_{123}}{p_1p_2p_3} \left(\frac{\pi}{2}\right)^{3/2} \int_0^\infty \mathrm{d}x \, x^{-1+\epsilon} \mathrm{e}^{-(p_1+p_2+p_3)x} = \frac{c_{123}}{p_1p_2p_3} \left(\frac{\pi}{2}\right)^{3/2} (p_1+p_2+p_3)^{-\epsilon} \Gamma(\epsilon) = \frac{c_{123}}{p_1p_2p_3} \left(\frac{\pi}{2}\right)^{3/2} \left[\frac{1}{\epsilon} -\log(p_1+p_2+p_3) - \gamma_E + O(\epsilon)\right],$$

$$(3.107)$$

where in the last equality we have expanded in ϵ , using

$$\Gamma(\epsilon) = \frac{1}{\epsilon} - \gamma_E + O(\epsilon), \qquad (3.108)$$

with γ_E the Euler-Mascheroni constant. This 3-point function is divergent for $\epsilon \to 0$. However, the (+ + +) condition does not admit any counterterm. This is because the leading term of (3.107) is a non-analytic function of the squared momentum magnitudes and hence this is a non-local, physically acceptable, solution. We then eliminate the divergence by choosing $c_{123} = C_{123}\epsilon$. Thus,

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = \left(\frac{\pi}{2}\right)^{3/2} \frac{C_{123}}{p_1 p_2 p_3}.$$
 (3.109)

2. (--) condition: d = 3, $\Delta_1 = \Delta_2 = \Delta_3 = 2$ ($\alpha = 1/2$ and $\beta_i = 1/2$, with i = 1, 2, 3).

Using the regularisation scheme defined in (3.104), this 3-point function reads

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = -c_{123}(p_1p_2p_3)^{\frac{1}{2}} \int_0^\infty \mathrm{d}x \, x^{\frac{1}{2}+\epsilon} K_{\frac{1}{2}}(p_1x)K_{\frac{1}{2}}(p_2x)K_{\frac{1}{2}}(p_3x),$$
(3.110)

where the overall minus sign is for convenience. This integral is the same as the one in (3.105), so here we have

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle = c_{123} \left(\frac{\pi}{2}\right)^{3/2} \left[-\frac{1}{\epsilon} + \log(p_1 + p_2 + p_3) + \gamma_E + O(\epsilon)\right].$$
(3.111)

However, the leading term in (3.111) is a constant and the divergence for $\epsilon \to 0$ is ultralocal. This divergence is canceled by the counterterm

$$S_{ct} = a(\epsilon) \int d^{3+2\epsilon} \boldsymbol{x} \phi^3 \mu^{-\epsilon}, \qquad (3.112)$$

where μ is the renormalisation scale and ϕ is the source field of the operators \mathcal{O} . To cancel the divergence, we choose

$$a(\epsilon) = \frac{1}{6} c_{123} \left(\frac{\pi}{2}\right)^{3/2} \left(\frac{1}{\epsilon} + a_0\right), \qquad (3.113)$$

where a_0 is an arbitrary constant dependent on the regularisation scheme. Then, the renormalised 3-point function is

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\rangle\!\rangle_{\mathrm{ren}} = c_{123} \left(\frac{\pi}{2}\right)^{3/2} \left[\log\left(\frac{p_1+p_2+p_3}{\mu}\right) + c_1\right],$$
 (3.114)

where $c_1 = c_0 + \gamma_E$. Therefore we find, as expected for the (---) condition, that the correlator depends on the RG scale μ . Hence the conformal symmetry is broken and a conformal anomaly A exists:

$$A = \int \mathrm{d}^d \boldsymbol{x} \mathcal{A}_{222} (\Box^2 \phi)^3, \qquad (3.115)$$

where we defined

$$\mathcal{A}_{222} = \mu \frac{\mu}{\partial \mu} \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2(\boldsymbol{p}_2) \mathcal{O}_3(\boldsymbol{p}_3) \rangle\!\rangle_{\text{ren}} = -c_{123} \left(\frac{\pi}{2}\right)^{3/2}.$$
 (3.116)

3.6 Shift operators

We conclude the analysis on 3-point functions by presenting their shift operators. These are operators that act on a 3-point function to shift the parameters d and Δ_i (i = 1, 2, 3), or equivalently α and β_i . In other words, they connect two solutions of the CWIs with shifted parameters. Earlier in this chapter we discussed the form of 3-point functions depending on the values of β_i . For half-integer β_i the triple-K integral can easily be computed in terms of elementary functions. For integer β_i the 3-point function is expressible in terms of the dilogarithm function, however, the computation is cumbersome. We showed the evaluation of the master integral $I_{1\{000\}}$ ($\alpha = 1$ and $\beta_i = 0$) in appendix A. For larger integer values of β_i the computation is more complicated. One strategy to obtain this class of triple-K is to generate them by acting on the master integral with a shift operator. This leads to the reduction scheme for the evaluation of 3-point functions discussed in [44].

In this section we show that two types of such operators exist. The first family of shift operators acts to shift both the spacetime dimension d up by two (hence α up by one) and one β_i up or down by one. We derive their form using the properties of Bessel-K functions, as shown in [42, 85]. In Chapter 4, we will find the general form of these operators acting on n-point functions, using a new scalar representation of the n-point functions. The second set of operators act to shift two of the β_i up or down while leaving the spacetime dimension invariant and will be further discussed in Chapter 5. We show that their expression can be understood easily in position space. Finally, we derive some recursion relations for triple-K integrals.

3.6.1 Operators shifting d

The 3-point d-shifting operators are

$$\mathcal{L}_{i} = -\frac{1}{p_{i}}\frac{\partial}{\partial p_{i}}, \qquad \mathcal{R}_{i} = 2\beta_{i} - p_{i}\frac{\partial}{\partial p_{i}}, \qquad \beta_{i} = \Delta_{i} - \frac{d}{2}.$$
(3.117)

They act on the 3-point function by sending

 $\mathcal{L}_i: \quad \beta_i \to \beta_i - 1, \quad \alpha \to \alpha + 1, \qquad \mathcal{R}_i: \quad \beta_i \to \beta_i + 1, \quad \alpha \to \alpha + 1, \qquad (3.118)$

or equivalently,

$$\mathcal{L}_1: \quad (d, \Delta_1, \Delta_2, \Delta_3) \to (d+2, \Delta_1, \Delta_2+1, \Delta_3+1), \tag{3.119}$$

$$\mathcal{R}_1: \quad (d, \Delta_1, \Delta_2, \Delta_3) \to (d+2, \Delta_1+2, \Delta_2+1, \Delta_3+1), \tag{3.120}$$

and similarly under any permutation in the set $\{1, 2, 3\}$. To understand their expressions and actions, consider the following properties of Bessel functions

$$K_{\nu} = K_{-\nu},$$
 (3.121)

$$\frac{\partial}{\partial p} \left[p^{\nu} K_{\nu}(px) \right] = -x p^{\nu} K_{\nu-1}(px). \tag{3.122}$$

The first property (3.121), together with the definition of the triple-K integral (3.35), imply that

$$I_{\alpha\{-\beta_1\beta_2\beta_3\}} = p_1^{-2\beta_1} I_{\alpha\{\beta_1\beta_2\beta_3\}}, \qquad (3.123)$$

hence the operator $p_i^{-2\beta_i}$ sends $\beta_i \to -\beta_i$. This is effectively a shadow transformation, sending $\Delta_i \to d - \Delta_i$. From the property (3.122), we see that \mathcal{L}_1 acts on the triple-K integral as

$$\mathcal{L}_1 I_{\alpha\{\beta_1\beta_2\beta_3\}} = I_{\alpha+1\{\beta_1-1,\beta_2\beta_3\}}.$$
(3.124)

This can be seen by direct computation

$$\mathcal{L}_{1}I_{\alpha\{\beta_{1}\beta_{2}\beta_{3}\}} = c_{123} \int_{0}^{\infty} \mathrm{d}x \, x^{\alpha} \left[-\frac{1}{p_{1}} \frac{\partial}{\partial p_{1}} \left(p_{1}^{\beta_{1}} K_{\beta_{1}}(p_{1}x) \right) \right] p_{2}^{\beta_{2}} p_{3}^{\beta_{3}} K_{\beta_{2}}(p_{2}x) K_{\beta_{3}}(p_{3}x)$$
$$= c_{123} \int_{0}^{\infty} \mathrm{d}x \, x^{\alpha} \left(x \, p_{1}^{\beta_{1}-1} K_{\beta_{1}-1}(p_{1}x) \right) p_{2}^{\beta_{2}} p_{3}^{\beta_{3}} K_{\beta_{2}}(p_{2}x) K_{\beta_{3}}(p_{3}x)$$
$$= c_{123} \int_{0}^{\infty} \mathrm{d}x \, x^{\alpha+1} p_{1}^{\beta_{1}-1} p_{2}^{\beta_{2}} p_{3}^{\beta_{3}} K_{\beta_{1}-1}(p_{1}x) K_{\beta_{2}}(p_{2}x) K_{\beta_{3}}(p_{3}x). \quad (3.125)$$

Combining \mathcal{L}_1 and the shadow transform we obtain the operator \mathcal{R}_1

$$\mathcal{R}_1 = p_1^{2\beta_1 + 2} \,\mathcal{L}_1 \, p_1^{-2\beta_1}, \tag{3.126}$$

which acts on the triple-K as

$$\mathcal{R}_1 I_{\alpha\{\beta_1\beta_2\beta_3\}} = I_{\alpha+1\{\beta_1+1,\beta_2\beta_3\}},\tag{3.127}$$

since

$$(\beta_i, \alpha) \xrightarrow{p_i^{-2\beta_i}} (-\beta_i, \alpha) \xrightarrow{\mathcal{L}_i} (-\beta_i - 1, \alpha + 1) \xrightarrow{p_i^{-2(-\beta_i - 1)} = p_i^{2(\beta_i + 1)}} (\beta_i + 1, \alpha + 1).$$
(3.128)

It is interesting to note that the combination $\mathcal{L}_i \mathcal{R}_i$ amounts to the special conformal operator we introduced in (3.33),

$$\mathcal{L}_i \mathcal{R}_i = \mathcal{K}_i = \frac{\partial^2}{\partial p_i^2} + \frac{1 - 2\beta_i}{\partial} \frac{\partial}{\partial p_i}.$$
(3.129)

This operator acts on a triple-K integral to shift α up by two, since \mathcal{L}_i and \mathcal{R}_i shift β_i in opposite directions but both increase α by one:

$$K_i I_{\alpha\{\beta_1\beta_2\beta_3\}} = I_{\alpha+2\{\beta_1\beta_2\beta_3\}}, \qquad i = 1, 2, 3.$$
(3.130)

Triple-K recursion relations

Using the actions of the above operators above and the CWIs we can derive some additional recursion relations. From the dilatation Ward identity (3.25), expressing the operators $p_i \partial_{p_i}$ in terms of \mathcal{L}_i and \mathcal{R}_i , we have

$$(\alpha + 1 - \beta_t) I_{\alpha\{\beta_1\beta_2\beta_3\}} = \left(p_1^2 \mathcal{L}_1 + p_2^2 \mathcal{L}_2 + p_3^2 \mathcal{L}_3 \right) I_{\alpha\{\beta_1\beta_2\beta_3\}}, \tag{3.131}$$

$$(\alpha + 1 + \beta_t) I_{\alpha\{\beta_1\beta_2\beta_3\}} = (\mathcal{R}_1 + \mathcal{R}_2 + \mathcal{R}_3) I_{\alpha\{\beta_1\beta_2\beta_3\}}, \qquad (3.132)$$

and considering the action (3.124) of the operators \mathcal{L}_i and (3.127) of \mathcal{R}_i , we obtain

$$(\alpha + 1 - \beta_t)I_{\alpha\{\beta_1\beta_2\beta_3\}} = p_1^2 I_{\alpha+1\{\beta_1-1,\beta_2,\beta_3\}} + p_2^2 I_{\alpha+1\{\beta_1,\beta_2-1,\beta_3\}} + p_3^2 I_{\alpha+1\{\beta_1,\beta_2,\beta_3-1\}} (\alpha + 1 + \beta_t)I_{\alpha\{\beta_1\beta_2\beta_3\}} = I_{\alpha+1\{\beta_1+1,\beta_2,\beta_3\}} + I_{\alpha+1\{\beta_1,\beta_2+1,\beta_3\}} + I_{\alpha+1\{\beta_1,\beta_2,\beta_3+1\}}.$$
 (3.133)

Taking into account that

$$\mathcal{L}_1 I_{\alpha\{\beta_1\beta_2\beta_3\}} = \mathcal{R}_2 \mathcal{R}_3 I_{\alpha-1\{\beta_1-1\beta_2-1\beta_3-1\}},$$

$$\mathcal{R}_1 I_{\alpha\{\beta_1\beta_2\beta_3\}} = \mathcal{L}_2 \mathcal{L}_3 I_{\alpha-1\{\beta_1+1,\beta_2+1,\beta_3+1\}},\tag{3.134}$$

equations (3.131) and (3.132) are equivalent to the following identities

$$\left(\alpha + 1 - \beta_t\right)I_{\alpha\{\beta_1\beta_2\beta_3\}} = \left[p_1^2\mathcal{R}_2\mathcal{R}_3 + p_2^2\mathcal{R}_1\mathcal{R}_3 + p_3^2\mathcal{R}_1\mathcal{R}_2\right]I_{\alpha - 1\{\beta_1 - 1\beta_2 - 1\beta_3 - 1\}}, \quad (3.135)$$

$$(\alpha + 1 + \beta_t)I_{\alpha\{\beta_1\beta_2\beta_3\}} = (\mathcal{L}_1\mathcal{L}_2 + \mathcal{L}_1\mathcal{L}_3 + \mathcal{L}_2\mathcal{L}_3)I_{\alpha - 1\{\beta_1 + 1, \beta_2 + 1, \beta_3 + 1\}}.$$
(3.136)

3.6.2 Operators preserving d

In this section we discuss shift operators acting on solutions of the CWIs to shift the scaling dimensions while preserving the spacetime dimension. The best known operators in literature [30, 83] are those acting on the CWI solutions to shift two scaling dimensions Δ_i up or down by one unit. They are denoted by $\mathcal{W}_{ij}^{\sigma_i \sigma_j}$, with $\sigma_i = \pm 1$, and their action is to send

$$d \to d \qquad \Delta_i \to \Delta_i + \sigma_i, \qquad \Delta_j \to \Delta_j + \sigma_j,$$
(3.137)

or equivalently,

$$\alpha \to \alpha, \qquad \beta_i \to \beta_i + \sigma_i, \qquad \beta_j \to \beta_j + \sigma_j.$$
 (3.138)

One might wonder why the shift operators introduced until now act either to shift two parameters by one unit or act to shift one parameter by two units (see (3.130)). This is not a trivial question and we will address it in Chapter 5.

In the following we introduce the expressions of \mathcal{W} -operators and derive their action on the 3-point functions. We will consider $\mathcal{W}_{12}^{\pm\pm}$, since the expressions for general i, j can be obtained by permutation. According to recent understanding of these operators [36] in momentum space, their expressions read

$$\mathcal{W}_{12}^{--} = \frac{1}{2} \left(\frac{\partial}{\partial p_1^{\mu}} - \frac{\partial}{\partial p_2^{\mu}} \right) \left(\frac{\partial}{\partial p_{1,\mu}} - \frac{\partial}{\partial p_{2,\mu}} \right), \tag{3.139}$$

$$\mathcal{W}_{12}^{+-} = p_1^{2(\beta_1+1)} \mathcal{W}_{12}^{--} p_1^{-2\beta_1}, \qquad (3.140)$$

$$\mathcal{W}_{12}^{-+} = p_2^{2(\beta_2+1)} \mathcal{W}_{12}^{--} p_2^{-2\beta_2}, \qquad (3.141)$$

$$\mathcal{W}_{12}^{++} = p_1^{2(\beta_1+1)} p_2^{2(\beta_2+1)} \mathcal{W}_{12}^{--} p_1^{-2\beta_1} p_2^{-2\beta_2}.$$
(3.142)

We explain below the expression of \mathcal{W}_{12}^{--} , while the remaining ones are obtained by shadow transforming \mathcal{W}_{12}^{--} .

Lowering operator \mathcal{W}_{12}^{--}

The easiest way to derive equation (3.139), is to start from position space. In fact given the form of the general *n*-point solution (2.67), it is intuitive to find the position-space expression of the lowering operator \mathcal{W}_{12}^{--} . As we showed in Chapter 2, the general positionspace *n*-point function can be written as

$$\phi_n = \langle O(\boldsymbol{x}_1) ... O(\boldsymbol{x}_n) \rangle = \prod_{1 \le i < j \le n} x_{ij}^{2\alpha_{ij}} f(\boldsymbol{u}), \qquad (3.143)$$

and the parameters α_{ij} are related to the scaling dimensions Δ_i via

$$\Delta_i = -\sum_{j=1}^n \alpha_{ij}, \quad i = 1, 2, .., n \tag{3.144}$$

where $\alpha_{ij} = \alpha_{ji}$ and $\alpha_{ii} = 0$. It is straightforward to understand that x_{ij}^2 is an operator shifting Δ_i and Δ_j down by one. If we multiply (3.143) by x_{ij}^2 , for some specific choice of *i* and *j*, this serves to shift $\alpha_{ij} \to \alpha_{ij} + 1$ and hence $\Delta_i \to \Delta_i - 1$ and $\Delta_j \to \Delta_j - 1$. Multiplying by x_{ij}^2 thus acts as a lowering operator generating a new solution of the *n*point conformal Ward identities in which the dimensions Δ_i and Δ_j are reduced by one while preserving $f(\boldsymbol{u})$ and the spacetime dimension *d*.

To find the corresponding expression in momentum space we perform a Fourier transform. The Fourier transformed n-point function is

$$\Phi_n = \mathcal{F}[\phi_n] = \int \mathrm{d}\boldsymbol{x}_1 ... \mathrm{d}\boldsymbol{x}_n \,\mathrm{e}^{-i\sum_{j=1}^n \boldsymbol{x}_j \cdot \boldsymbol{p}_j} \phi_n(\boldsymbol{x}_1, ..., \boldsymbol{x}_n), \qquad (3.145)$$

where, taking into account momentum conservation,

$$\sum_{j=1}^{n} \boldsymbol{x}_{j} \cdot \boldsymbol{p}_{j} = \sum_{j=1}^{n-1} \boldsymbol{p}_{j} \boldsymbol{x}_{j} - \left(\sum_{j=1}^{n-1} \boldsymbol{p}_{j}\right) \boldsymbol{x}_{n} = \sum_{j=1}^{n-1} \boldsymbol{p}_{j} \boldsymbol{x}_{jn}.$$
(3.146)

Therefore, pulling down a factor of x_{12} is equivalent to acting on the Fourier transformed n-point function with the difference of derivatives with respect to p_1^{μ} and p_2^{μ} :

$$\mathcal{F}\left[\boldsymbol{x}_{12}\phi_{4}\right] = i\left(\frac{\partial}{\partial p_{1}^{\mu}} - \frac{\partial}{\partial p_{2}^{\mu}}\right)\Phi_{4},\tag{3.147}$$

hence

$$\mathcal{W}_{12}^{--} = \mathcal{F}\left[-\frac{1}{2}x_{12}^2\right] = \frac{1}{2}\left(\frac{\partial}{\partial p_1^{\mu}} - \frac{\partial}{\partial p_2^{\mu}}\right)\left(\frac{\partial}{\partial p_{1,\mu}} - \frac{\partial}{\partial p_{2,\mu}}\right),\tag{3.148}$$

where the factor -1/2 is purely conventional [30]. This is valid at *n*-point. However, here we focus on the 3-point function. Using the chain rule (3.24) we obtain the expression of the 3-point \mathcal{W}_{12}^{--} in terms of Mandelstam variables:

$$\mathcal{W}_{12}^{--} = \frac{1}{2} \bigg[\partial_1^2 + \partial_2^2 + \frac{d-1}{p_1} \partial_1 + \frac{d-1}{p_2} \partial_2 + \frac{p_1^2 + p_2^2 - p_3^2}{p_1 p_2} \partial_1 \partial_2 \bigg].$$
(3.149)

Let us now derive the action on the 3-point function. By construction, the lowering operator acts to generate a 3-point function with Δ_1 and Δ_2 lowered by one. To show this, we use the triangle representation (3.75). Using the equivalence between the 1-loop triangle integral and the triple-K representation, we then derive the action on the triple-K integral. To simplify the computation, we write the triangle integral by re-parametrising the loop momentum:

$$I_{d\{\alpha_{12},\alpha_{13},\alpha_{23}\}} = \int \frac{\mathrm{d}^{d} \boldsymbol{q}}{(2\pi)^{d}} \frac{1}{|\boldsymbol{p}_{1} + \boldsymbol{p}_{2} + \boldsymbol{q}|^{2\alpha_{23} + d} |\boldsymbol{q}|^{2\alpha_{13} + d} |\boldsymbol{p}_{1} + \boldsymbol{q}|^{2\alpha_{12} + d}},$$
(3.150)

and use the following identities

$$\left(\frac{\partial}{\partial p_1^{\mu}} - \frac{\partial}{\partial p_2^{\mu}}\right) f(\boldsymbol{p}_1 + \boldsymbol{p}_2) = 0, \qquad (3.151)$$

$$\left(\frac{\partial}{\partial p_1^{\mu}} - \frac{\partial}{\partial p_2^{\mu}}\right) f(\boldsymbol{p}_1) = \frac{\partial}{\partial p_1^{\mu}} f(\boldsymbol{p}_1), \qquad (3.152)$$

where f is a generic function. Then, we only need to compute

$$\delta^{\mu\nu} \frac{\partial}{\partial p_1^{\mu}} \frac{\partial}{\partial p_1^{\nu}} \frac{1}{|\boldsymbol{p}_1 + \boldsymbol{q}|^{2\alpha_{12} + d}} = \frac{2\left(\alpha_{12} + \frac{d}{2}\right)\left(\alpha_{12} + 1\right)}{|\boldsymbol{p}_1 + \boldsymbol{q}|^{2(\alpha_{12} + 1) + d}}.$$
(3.153)

Hence the action of \mathcal{W}_{12}^{--} on the triangle integral is

$$\mathcal{W}_{12}^{--}I_{d\{\alpha_{12},\alpha_{13},\alpha_{23}\}} = 2\left(\alpha_{12} + \frac{d}{2}\right)(\alpha_{12} + 1)I_{d\{\alpha_{12}+1,\alpha_{13},\alpha_{23}\}}.$$
(3.154)

Note that, according to (3.87), sending $\alpha_{12} \rightarrow \alpha_{12} + 1$ is equivalent to send $\beta_1 \rightarrow \beta_1 - 1$ and $\beta_2 \rightarrow \beta_2 - 1$. We derive the action on the triple-K by combining (3.154) with the triangle/triple-K equivalence (3.76)

$$\mathcal{W}_{12}^{--}I_{\alpha\{\beta_{1}\beta_{2}\beta_{3}\}} = 2\left(\alpha_{12} + \frac{d}{2}\right)(\alpha_{12} + 1)\frac{C_{T}}{C_{T}|_{\alpha_{12}+1}}I_{\alpha\{\beta_{1}-1,\beta_{2}-1,\beta_{3}\}}$$
$$= \frac{1}{2}\left[\beta_{3}^{2} - (\alpha - 1 + \beta_{1} + \beta_{2})^{2}\right]I_{\alpha,\{\beta_{1}-1,\beta_{2}-1,\beta_{3}\}}, \qquad (3.155)$$

where C_T is given in (3.77). Let us anticipate here that in the second part of this thesis, following the work [86], we explain how to derive the factor involving the parameters α and β_i by knowing the singularities of the triple-K.

Lowering-Raising \mathcal{W}_{12}^{-+} and raising operator \mathcal{W}_{12}^{++}

The momentum-space expressions of \mathcal{W}_{12}^{-+} and \mathcal{W}_{12}^{++} can be derived using the shadow transform as in (3.140-3.142). In fact, let us show that these definitions shift the parameters as we want, *i.e.*, $\beta_1 \rightarrow \beta_1 - 1$, $\beta_2 \rightarrow \beta_2 + 1$ and $\beta_1 \rightarrow \beta_1 + 1$, $\beta_2 \rightarrow \beta_2 + 1$ respectively :

$$\mathcal{W}_{12}^{-+}: \quad (\beta_1, \beta_2, \beta_3) \xrightarrow{p_2^{-2\beta_2}} (\beta_1, -\beta_2, \beta_3) \xrightarrow{\mathcal{W}_{12}^{--}} (\beta_1 - 1, -\beta_2 - 1, \beta_3) \xrightarrow{\frac{p_2^{-2\beta_2}}{p_2^{-(\beta_2+1)}}} (\beta_1 - 1, \beta_2 + 1, \beta_3), \quad (3.156)$$

$$\mathcal{W}_{12}^{++}: \quad (\beta_1, \beta_2, \beta_3) \xrightarrow{p_1^{-2\beta_1} p_2^{-2\beta_2}} (-\beta_1, -\beta_2, \beta_3) \xrightarrow{\mathcal{W}_{12}^{--}} (-\beta_1 - 1, -\beta_2 - 1, \beta_3) \xrightarrow{p_1^{2(\beta_1+1)} p_2^{2(\beta_2+1)}} (\beta_1 + 1, \beta_2 + 1, \beta_3). \quad (3.157)$$

By multiplying out the right-hand sides in (3.140)-(3.142), we obtain their explicit expressions

$$\mathcal{W}_{12}^{-+} = p_2^2 \mathcal{W}_{12}^{--} + 2\beta_2 \left(\beta_2 + 1 - \frac{d}{2} + p_2^{\mu} \partial_{12\mu}\right)$$
(3.158)

$$\mathcal{W}_{12}^{+-} = p_1^2 \mathcal{W}_{12}^{--} + 2\beta_1 \Big(\beta_1 + 1 - \frac{d}{2} - p_1^{\mu} \partial_{12\mu}\Big), \tag{3.159}$$

$$\mathcal{W}_{12}^{++} = p_1^2 p_2^2 \mathcal{W}_{12}^{--} + 2\beta_1 \beta_2 (p_1^2 + p_2^2 - p_3^2) + 2\beta_1 p_2^2 \Big(\beta_1 + 1 - \frac{d}{2} - p_1^{\mu} \partial_{12\mu}\Big) + 2\beta_2 p_1^2 \Big(\beta_2 + 1 - \frac{d}{2} + p_2^{\mu} \partial_{12\mu}\Big), \qquad (3.160)$$

where $\partial_{12\mu}$ denotes the difference $\partial/\partial p_1^{\mu} - \partial/\partial p_2^{\mu}$.

3.6.3 Intertwining relations

In this section we conclude the discussion on shift operators presenting an algebraic method to verify the action of a shift operator. Let $\mathcal{I}_{\{d,\Delta_i\}}$ be the solution of the CWIs (3.1)-(3.2) and X_i a generic shift operator acting on the solution to shift $d \to d'$ and $\Delta \to \Delta'_i$, for some *i*. Then X_i has to satisfy the following intertwining relation

$$\mathcal{K}^{\mu}[d', \Delta_{i}']X_{i} - X_{i}\mathcal{K}^{\mu}[d, \Delta_{i}] = \hat{O}_{1}\mathcal{K}^{\mu}[d, \Delta_{i}] + \hat{O}_{2}^{\mu}D[d, \Delta_{i}], \qquad (3.161)$$

where \mathcal{K}^{μ} is the special conformal operator in (3.4) and \hat{O}_j are some differential operators such that the homogeneity of the equation holds. Note that the right-hand side is an operator that annihilates the solution $\mathcal{I}_{\{d,\Delta_i\}}$. While the left-hand side holds, since

$$0 = \mathcal{K}^{\mu}[d', \Delta'_{i}]\mathcal{I}_{\{d', \Delta'_{i}\}} = \mathcal{K}^{\mu}[d', \Delta'_{i}]X_{i}\mathcal{I}_{\{d, \Delta_{i}\}} = X_{i}\mathcal{K}^{\mu}[d, \Delta_{i}]\mathcal{I}_{\{d, \Delta_{i}\}}.$$
 (3.162)

We can also consider the CWIs in Mandelstam variables, leading to the analogous relations:

$$K_{ij}[d',\Delta'_i]X_i - X_i K_{ij}[d,\Delta_i] = \hat{O}_1 K_{ij}[d,\Delta_i] + \hat{O}_2 D[d,\Delta_i] \quad i,j = 1, 2, 3.$$
(3.163)

By direct computation we verified that the *d*-shifting operators \mathcal{L}_i and \mathcal{R}_i satisfy the intertwining equation (3.163) with the right-hand side equal to zero, *i.e.*,

$$K_{12}[d+2, \Delta_1, \Delta_2+1, \Delta_3+1]\mathcal{L}_i = \mathcal{L}_i K_{12}[d, \Delta_1, \Delta_2, \Delta_3].$$
(3.164)

The weight-shifting operator W_{12} , instead, satisfies (3.163) with a non zero right-hand side of (3.163) different from zero. To be precise, its intertwining relation is

$$K_{12}[d, \Delta_1 - 1, \Delta_2 - 1, \Delta_3] \mathcal{W}_{12}^{--} - \mathcal{W}_{12}^{--} K_{12}[d, \Delta_1, \Delta_2, \Delta_3] = \left(\frac{1}{p_1}\partial_1 + \frac{1}{p_2}\partial_2\right) K_{12}. \quad (3.165)$$

3.7 Higher-point functions

In Chapter 2 we derived the general solution of the scalar 4- and *n*-point CWIs in position space. While 3-point functions are uniquely fixed by conformal symmetry, 4- and higher-point functions are less tightly constrained and depend on an arbitrary function of the cross ratios. In this section we review recent work on general *n*-point correlation functions

in momentum space [47, 46]. The representation found has the form of a Feynman integral with the topology of an (n - 1)-simplex, featuring an arbitrary function of momentum-space cross ratios. This will also be developed further in the following chapter.

3.7.1 4-point function

We first discuss 4-point correlation functions of scalar operators and then we generalise the results to *n*-point functions. The general 4-point function in momentum space has been shown to be expressible as a 3-loop Feynman integral, where the (massless) scalar propagators are raised to generalised powers $\alpha_{ij} + d/2$:

$$\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\mathcal{O}_4(\boldsymbol{p}_4)\rangle = \left(\int \prod_{i=1}^3 \frac{\mathrm{d}^d \boldsymbol{q}_i}{(2\pi)^d}\right) \frac{\hat{f}(\hat{u},\hat{v})}{\mathrm{Den}(\boldsymbol{q}_j,\boldsymbol{p}_k)} (2\pi)^d \delta\Big(\sum_{j=1}^4 \boldsymbol{p}_j\Big), \quad (3.166)$$

where the denominator is

$$Den(\boldsymbol{q}_{j}, \boldsymbol{p}_{k}) = q_{3}^{2\alpha_{12}+d} q_{2}^{2\alpha_{13}+d} q_{1}^{2\alpha_{23}+d} |\boldsymbol{p}_{1} + \boldsymbol{q}_{2} - \boldsymbol{q}_{3}|^{2\alpha_{14}+d} \\ \times |\boldsymbol{p}_{2} + \boldsymbol{q}_{3} - \boldsymbol{q}_{1}|^{2\alpha_{24}+d} |\boldsymbol{p}_{3} + \boldsymbol{q}_{1} - \boldsymbol{q}_{2}|^{2\alpha_{34}+d}$$
(3.167)

and \hat{f} is an arbitrary function that depends on the two dependent momentum-space cross ratios

$$\hat{u} = \frac{q_2^2 |\boldsymbol{p}_2 + \boldsymbol{q}_3 - \boldsymbol{q}_1|^2}{q_3^2 |\boldsymbol{p}_3 + \boldsymbol{q}_1 - \boldsymbol{q}_2|^2}, \qquad \hat{v} = \frac{q_1^2 |\boldsymbol{p}_1 + \boldsymbol{q}_2 - \boldsymbol{q}_3|^2}{q_2^2 |\boldsymbol{p}_2 + \boldsymbol{q}_3 - \boldsymbol{q}_1|^2}.$$
(3.168)

They play a role analogous to the position-space cross ratios (2.65), however they depend on integration variables q_j , and so are not independent conformal invariants in their own right. In the following, we show the representation (3.166) is conformally invariant. First, we Fourier transform the position-space 4-point function (2.66). Then, we briefly discuss conformal invariance from a purely momentum-space perspective.

We already computed the Fourier transform of the 2-point function and this will be useful in our discussion. As a warm up example, we first Fourier transform the positionspace 3-point function (2.61) that we denote with ϕ_3 here

$$\phi_3 = x_{12}^{2\alpha_{12}} x_{13}^{2\alpha_{13}} x_{23}^{2\alpha_{23}}. \tag{3.169}$$

To Fourier transform ϕ_3 , we use the convolution theorem as follows

$$\mathcal{F}[\phi_3] = c_{123} \mathcal{F}\left[x_{12}^{2\alpha_{12}}\right] * \mathcal{F}\left[x_{13}^{2\alpha_{13}}x_{23}^{2\alpha_{23}}\right]$$
$$= c_{123} \int \frac{\mathrm{d}^d \boldsymbol{q}_1}{(2\pi)^d} \frac{\mathrm{d}^d \boldsymbol{q}_2}{(2\pi)^d} \mathcal{F}\left[x_{12}^{2\alpha_{12}}\right] (\boldsymbol{p}_j - \boldsymbol{q}_j) \mathcal{F}\left[x_{13}^{2\alpha_{13}}x_{23}^{2\alpha_{23}}\right] (\boldsymbol{p}_j), \quad j = 1, 2, 3. \quad (3.170)$$

Using equation (3.22) and taking into account that the Fourier transform in the integrand also depends on p_3 , we have

$$\mathcal{F}\left[x_{12}^{2\alpha_{12}}\right](\boldsymbol{p}_1, \boldsymbol{p}_2, \boldsymbol{p}_3) = (2\pi)^d \delta(\boldsymbol{p}_1 + \boldsymbol{p}_2) \delta(\boldsymbol{p}_3) \frac{C_{12}}{p_1^{2\alpha_{12}+d}}, \qquad (3.171)$$

and

$$\mathcal{F}\left[x_{13}^{2\alpha_{13}}x_{23}^{2\alpha_{23}}\right](\boldsymbol{p}_{1},\boldsymbol{p}_{2},\boldsymbol{p}_{3}) = (2\pi)^{d}\delta(\boldsymbol{p}_{1}+\boldsymbol{p}_{2}+\boldsymbol{p}_{3})\frac{C_{13}C_{23}}{p_{1}^{2\alpha_{13}+d}p_{2}^{2\alpha_{13}+d}},$$
(3.172)

where $\mathcal{F}[x_{ij}^{2\alpha_{ij}}...]$ denotes the Fourier transform over p_1 , p_2 and p_3 . The parameters α_{ij} are given in (2.60) and

$$C_{ij} = \frac{\pi^{d/2} 2^{2\alpha_{ij}+d}}{\Gamma(-\alpha_{ij})} \Gamma\left(\frac{d}{2} + \alpha_{ij}\right).$$
(3.173)

Then, substituting (3.171) and (3.172) in (3.170), we find

$$\mathcal{F}[\phi_3] = c_{123}C_{12}C_{13}C_{23} \int \frac{\mathrm{d}^d \boldsymbol{q}_1}{(2\pi)^d} \frac{\mathrm{d}^d \boldsymbol{q}_2}{(2\pi)^d} \frac{(2\pi)^d \delta(\boldsymbol{p}_3 + \boldsymbol{q}_1 + \boldsymbol{q}_2)}{q_1^{2\alpha_{13}+d} q_2^{2\alpha_{23}+d}} \frac{(2\pi)^d \delta(\boldsymbol{p}_1 + \boldsymbol{p}_2 + \boldsymbol{p}_3)}{|\boldsymbol{p}_1 - \boldsymbol{q}_1|^{2\alpha_{12}+d}}.$$
(3.174)

Integrating over q_2 and sending $q_1 \rightarrow -q_1 + p_1$ we recover the 1-loop triangle integral (3.75):

$$\mathcal{F}[\phi_3] = c_{123}C_{12}C_{13}C_{23} \int \frac{\mathrm{d}^d \boldsymbol{q}}{(2\pi)^d} \frac{(2\pi)^d \delta(\boldsymbol{p}_1 + \boldsymbol{p}_2 + \boldsymbol{p}_3)}{|\boldsymbol{q}|^{2\alpha_{12}+d}|\boldsymbol{q} - \boldsymbol{p}_1|^{2\alpha_{13}+d}|\boldsymbol{q} + \boldsymbol{p}_2|^{2\alpha_{23}+d}}.$$
(3.175)

Now, we move to 4-point functions. To find the simplex representation of 4-point functions (3.166), we Fourier transform the position-space 4-point function in identical fashion to the 3-point function. Note that in this case we also have to consider the arbitrary function. First, let us assume a monomial position-space arbitrary function $f(u, v) = u^{\alpha}v^{\beta}$ in equation (2.66), then the position-space 4-point function is just a product of powers of x_{ij}^2 :

$$\phi_4(\alpha,\beta) = x_{12}^{2(\alpha_{12}+\alpha)} x_{13}^{2(\alpha_{13}-\alpha+\beta)} x_{14}^{2(\alpha_{14}-\beta)} x_{23}^{2(\alpha_{23}-\beta)} x_{24}^{2(\alpha_{24}-\alpha+\beta)} x_{34}^{2(\alpha_{34}+\alpha)} = \prod_{1 \le i < j \le n} x_{ij}^{2\gamma_{ij}},$$
(3.176)

where

$$\gamma_{12} = \alpha_{12} + \alpha, \qquad \gamma_{13} = \alpha_{13} - \alpha + \beta, \qquad \gamma_{14} = \alpha_{14} - \beta, \\
\gamma_{23} = \alpha_{23} - \beta, \qquad \gamma_{24} = \alpha_{24} - \alpha + \beta, \qquad \gamma_{34} = \alpha_{34} + \alpha.$$
(3.177)

Then, we use the convolution theorem by grouping the powers of x_{ij} in a way that a recursive approach can be used:

$$\mathcal{F}[\phi_4(\alpha,\beta)] = \mathcal{F}\left[x_{14}^{2\gamma_{14}}x_{24}^{2\gamma_{24}}x_{34}^{2\gamma_{34}}\right] * \mathcal{F}\left[x_{12}^{2\gamma_{12}}x_{13}^{2\gamma_{13}}x_{23}^{2\gamma_{23}}\right]$$
$$= \mathcal{F}\left[x_{14}^{2\gamma_{14}}x_{24}^{2\gamma_{24}}x_{34}^{2\gamma_{34}}\right] * \mathcal{F}[\phi_3](2\pi)^d\delta(\boldsymbol{p_4}), \tag{3.178}$$

where the second factor on the right-hand side is given by the 3-point Fourier transform (3.175). This makes the the recursive structure evident and will be useful for the general-

isation to n-point functions. Taking into account that

$$\mathcal{F}\left[x_{14}^{2\gamma_{14}}x_{24}^{2\gamma_{24}}x_{34}^{2\gamma_{34}}\right] = \frac{(2\pi)^d \delta\left(\sum_{j=1}^4 \boldsymbol{p}_j\right)\prod_{i=1}^4 C_{i4}}{p_1^{2\gamma_{14}+d}p_2^{2\gamma_{24}+d}p_3^{2\gamma_{34}+d}},$$
(3.179)

using the result (3.175) and applying the convolution theorem we find

$$\mathcal{F}[\phi_4(\alpha,\beta)] = \int \frac{\mathrm{d}^d \boldsymbol{q}_1}{(2\pi)^d} \frac{\mathrm{d}^d \boldsymbol{q}_2}{(2\pi)^d} \frac{\mathrm{d}^d \boldsymbol{q}_3}{(2\pi)^d} \frac{\prod_{1 \le i < j \le 4} C_{ij}}{\mathrm{Den}_3^{\alpha\beta}(\boldsymbol{q}_j, \boldsymbol{p}_k)} (2\pi)^d \delta\Big(\sum_{j=1}^4 \boldsymbol{p}_j\Big), \tag{3.180}$$

with

$$Den_{3}^{\alpha\beta}(\boldsymbol{q}_{j},\boldsymbol{p}_{k}) = q_{3}^{2\gamma_{12}+d} q_{2}^{2\gamma_{13}+d} q_{1}^{2\gamma_{23}+d} |\boldsymbol{p}_{1} + \boldsymbol{q}_{2} - \boldsymbol{q}_{3}|^{2\gamma_{14}+d} \times |\boldsymbol{p}_{2} + \boldsymbol{q}_{3} - \boldsymbol{q}_{1}|^{2\gamma_{24}+d} |\boldsymbol{p}_{3} + \boldsymbol{q}_{1} - \boldsymbol{q}_{2}|^{2\gamma_{34}+d}, \qquad (3.181)$$

where the γ_{ij} depend on α and β as in (3.177). Note that this is the right-hand side of (3.166) with the momentum-space arbitrary function

$$\hat{f}(\hat{u}, \hat{v}) = \prod_{1 \le i < j \le 4} C_{ij} \hat{u}^{\alpha} \hat{v}^{\beta}.$$
 (3.182)

So far we have shown that a 3-simplex integral (3.166) with \hat{f} in (3.182) is a solution of CWIs. Next, we want to show that this is also valid for any arbitrary function $\hat{f}(\hat{u}, \hat{v})$. To see this, we take into account that (3.180) is a solution of CWIs and use the inverse Mellin transform to express a general \hat{f} :

$$\hat{f}(\hat{u},\hat{v}) = \frac{1}{(2\pi i)^2} \int_{a-i\infty}^{a+i\infty} \int_{b-i\infty}^{b+i\infty} \mathrm{d}\alpha \,\mathrm{d}\beta \,\rho(\alpha,\beta) \hat{u}^{\alpha} \hat{v}^{\beta},\tag{3.183}$$

for some appropriate choice of integration contour specified by a and b. Since equation (3.166) can be written as

$$\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\mathcal{O}_4(\boldsymbol{p}_4)\rangle = \frac{1}{(2\pi i)^2} \int_{a-i\infty}^{a+i\infty} \int_{b-i\infty}^{b+i\infty} \mathrm{d}\alpha \,\mathrm{d}\beta \,\rho(\alpha,\beta)W_{\alpha,\beta},\qquad(3.184)$$

where we defined

$$W_{\alpha,\beta} = \frac{1}{\prod_{1 \le i < j \le 4} C_{ij}} \mathcal{F}[\phi_4(\alpha,\beta)], \qquad (3.185)$$

we showed that (3.166) is the general conformal 4-point function in momentum space. Note that the solution (3.166) and the Fourier transform (3.180) differs by a factor consisting in the product of the constants C_{ij} defined in (3.173).

Conformal invariance via total derivative

In this section we discuss the solution (3.166) from the point of view of momentum space only. We will briefly discuss why (3.166) is a solution of the momentum-space CWIs (3.1), (3.2).

The DWI for the reduced *n*-point conformal correlators given in (3.1) constrains their expressions to scale as $\Delta_t - (n-1)d$. For n = 4, it means that the conformal 4-point function must scale as $\Delta_t - 3d$. By power counting, we find that this is indeed the scaling of the simplex representation (3.166) of 4-point functions. In fact, the three integrals contribute with dimension 3d, and each propagator scales as $-2\alpha_{ij}-d$, with $1 \le i < j \le 4$. Hence the scaling dimension of (3.166) is

$$-2\sum_{1\leq i< j\leq 4}\alpha_{ij} - 6d + 3d = -\sum_{i,j=1}^{4}\alpha_{ij} - 3d = \Delta_t - 3d, \qquad (3.186)$$

where in the last equality we used the equation (2.64) relating α_{ij} and the scaling dimensions of scalar operators.

Let us now move to the special conformal Ward identity. First, we recall that this is given in (3.4), and in this case $\mathcal{K}^{\mu} = \mathcal{K}^{\mu}_1 + \mathcal{K}^{\mu}_2$, with

$$\mathcal{K}_{j}^{\mu} = 2(\Delta_{j} - d)\frac{\partial}{\partial p_{j\mu}} - 2p_{j}^{\nu}\frac{\partial}{\partial p_{j}^{\nu}}\frac{\partial}{\partial p_{j\mu}} + p_{j}^{\mu}\frac{\partial}{\partial p_{j}^{\nu}}\frac{\partial}{\partial p_{j\nu}}.$$
(3.187)

To prove that (3.166) satisfies the 4-point special conformal Ward identity we have to show that the action of the SCWI operator \mathcal{K}^{μ} on the integrand of the simplex representation corresponds to a total derivative. Taking into account that \mathcal{K}^{μ} is a second-order differential operator, we expect its action on the integrand of (3.166) to be of the form

$$\mathcal{K}^{\mu}\left(\frac{\hat{f}(\hat{u},\hat{v})}{\operatorname{Den}(\boldsymbol{q}_{j},\boldsymbol{p}_{k})}\right) = \sum_{j=1}^{2} \frac{\partial}{\partial q_{j}^{k}} \left(A_{j}^{k\mu}\hat{f} + B_{j}^{k\mu}\frac{\partial\hat{f}}{\partial\hat{u}} + C_{j}^{k\mu}\frac{\partial\hat{f}}{\partial\hat{v}}\right),$$
(3.188)

for some coefficients $A_j^{k\mu}$, $B_j^{k\mu}$ and $C_j^{k\mu}$ that are *independent* of the arbitrary function \hat{f} . Their explicit computation and expressions can be found in [47] and [46]. In the following chapter we will prove the validity of the *n*-point CWIs using a new scalar parametrisation of the simplex integral. Hence, we will omit further details here of the explicit computation of these coefficients.

Before showing some special solutions of 4-point CWIs, we first extend the simplex representation to the n-point functions. This is, indeed, a generalisation of what we analysed in this section.

3.7.2 *n*-point function: the simplex representation

In this section we present the conformal *n*-point function of scalar operators known as *simplex integral*. This is a Feynman integral with the topology of an (n - 1)-simplex, featuring an arbitrary function of momentum-space cross ratios. This generalises the 4-point solution we introduced in the previous section and reads

$$\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle = \prod_{1 \le i < j \le n} \int \frac{\mathrm{d}^d \boldsymbol{q}_{ij}}{(2\pi)^d} \frac{\hat{f}(\hat{\boldsymbol{u}})}{q_{ij}^{2\alpha_{ij}+d}} \prod_{k=1}^n (2\pi)^d \delta\Big(\boldsymbol{p}_k + \sum_{l=1}^n \boldsymbol{q}_{lk}\Big), \qquad (3.189)$$

where q_{ij} is the momentum running in the oriented edge from vertex *i* to *j* so that $q_{ij} = -q_{ji}$ and $q_{jj} = 0$. The parameters α_{ij} are the same appearing in the position-space solution and satisfy the condition (2.64). We denoted the set of independent cross ratios with $\hat{\boldsymbol{u}}$, collecting the independent momentum-space cross ratios:

$$\hat{u}_{[pqrs]} = \frac{q_{pq}^2 q_{rs}^2}{q_{pr}^2 q_{qs}^2}.$$
(3.190)

Since the simplex integral is related to the position-space *n*-point solution through a Fourier transform, their number follows from the position-space argument explained at the end of chapter 2, *i.e.*, there are n(n-3)/2 for $n \leq d+2$ and nd - (d+2)(d+1)/2 when n > d+2. Moreover, as noted for the 4-point functions, momentum-space cross ratios depend on the integration variables.

Note that the integrand in (3.189) contains a product of n delta functions. To define the *reduced simplex integral* (denoted with the double-brackets), we set aside the delta function corresponding to momentum conservation

$$\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle = (2\pi)^d \delta \Big(\sum_{i=1}^n \boldsymbol{p}_i \Big) \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle\!\rangle,$$
 (3.191)

so that the reduced simplex integral depends only on n-1 independent external momenta. Then, we are left with n-1 delta functions. We choose the loop-parametrisation of the simplex by integrating over the variables q_{in} for i = 1, 2, ..., n-1. We obtain the following reduced simplex integral

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle\!\rangle = \prod_{1 \le i < j \le n-1} \int \frac{\mathrm{d}^d \boldsymbol{q}_{ij}}{(2\pi)^d} \frac{\hat{f}(\hat{\boldsymbol{u}})}{\mathrm{Den}_n(\boldsymbol{\alpha})}$$
(3.192)

where the denominator reads

$$Den_n(\alpha) = \prod_{1 \le i < j \le n-1} q_{ij}^{2\alpha_{ij}+d} \times \prod_{m=1}^{n-1} |\boldsymbol{l}_m - \boldsymbol{p}_m|^{2\alpha_{mn}+d}$$
(3.193)

and l_m depends only on the remaining internal momenta,

$$\boldsymbol{l}_{m} = -\boldsymbol{q}_{mn} + \boldsymbol{p}_{m} = \sum_{j=1}^{n-1} \boldsymbol{q}_{mj} = -\sum_{j=1}^{m-1} \boldsymbol{q}_{jm} + \sum_{j=m+1}^{n-1} \boldsymbol{q}_{mj}.$$
 (3.194)

The integral (3.192) displays (n-1)(n-2)/2 d-dimensional integrals. In the following chapter we describe instead new purely scalar parametrisations for the simplex which, amongst other advantages, feature fewer integrals.

Conformal invariance via Fourier transform

Here we want to understand the simplex integral (3.189) from Fourier transforming the position-space *n*-point function (3.143). The idea is to use the recursive structure already illustrated in the previous section. A more direct proof of conformal invariance for the

simplex integral will be given in the next chapter.

We follow the same strategy used in the previous section, *i.e.*, we first consider a monomial position-space arbitrary function f and derive the Fourier transform of the associated *n*-point function. We will see that this corresponds to the simplex integral with a monomial momentum-space arbitrary function. To show that for any generic arbitrary function the simplex is conformally invariant, we then take an inverse Mellin transform. Let F_n be the position-space *n*-point correlator (3.143) with a monomial arbitrary function. This is expressed as a product of powers of the independent cross ratios

$$F_n(\boldsymbol{\alpha}; \boldsymbol{x}_1, \dots, \boldsymbol{x}_n) = \prod_{1 \le i < j \le n} x_{ij}^{2\alpha_{ij}}, \qquad (3.195)$$

where the parameters α_{ij} satisfy (2.64). We want to show that its Fourier transform is

$$\mathcal{F}_{(n)}[F_n] = \prod_{1 \le i < j \le n} C_{ij} \int \frac{\mathrm{d}^d \boldsymbol{q}_{ij}}{(2\pi)^d} \frac{1}{q_{ij}^{2\alpha_{ij}+d}} \prod_{k=1}^n (2\pi)^d \delta\Big(\boldsymbol{p}_k + \sum_{l=1}^n \boldsymbol{q}_{lk}\Big), \tag{3.196}$$

where C_{ij} is given in (3.173). This integral is referred to as the mesh integral in [104], corresponding to a generalised Feynman integral with n points and n(n-1)/2 generalised (scalar, massless) propagators with every pair of points connected. Note that with 'generalised' we intend that the powers α_{ij} are not necessarily equal to one. This integral also corresponds to the simplex integral (3.189) when $\hat{f} = \prod_{1 \leq i < j \leq n} C_{ij}$. First, let us note that (3.195) is satisfied for n = 2:

$$\mathcal{F}_{(2)}[x_{12}^{2\alpha_{12}}](\alpha_{12};\boldsymbol{p}_1,\boldsymbol{p}_2) = C_{12} \int \frac{\mathrm{d}^d \boldsymbol{q}_{12}}{(2\pi)^d} \frac{1}{q_{12}^{2\alpha_{12}+d}} (2\pi)^d \delta(\boldsymbol{p}_1 - \boldsymbol{q}_{12}) (2\pi)^d \delta(\boldsymbol{p}_2 + \boldsymbol{q}_{12})$$
$$= (2\pi)^d \delta(\boldsymbol{p}_1 + \boldsymbol{p}_2) \frac{C_{12}}{p_1^{2\alpha_{12}+d}}.$$
(3.197)

This result is in agreement with the Fourier transform of the 2-point function (3.22). Before we perform the Fourier transform of (3.195) and make the recursive structure manifest, let us introduce the following notation

$$\mathcal{F}_{(n)}[F_n] = \int \mathrm{d}^d \boldsymbol{x}_1 \dots \mathrm{d}^d \boldsymbol{x}_n \mathrm{e}^{-i\sum_{j=1}^n \boldsymbol{x}_j \cdot \boldsymbol{p}_j} F_n(\boldsymbol{\alpha}; \boldsymbol{x}_1, \dots, \boldsymbol{x}_n)$$
$$\mathcal{F}_{(n)}[F_{n-1}] = \int \mathrm{d}^d \boldsymbol{x}_1 \dots \mathrm{d}^d \boldsymbol{x}_n \mathrm{e}^{-i\sum_{j=1}^n \boldsymbol{x}_j \cdot \boldsymbol{p}_j} F_{n-1}(\boldsymbol{\alpha}; \boldsymbol{x}_1, \dots, \boldsymbol{x}_{n-1}), \qquad (3.198)$$

and one can show that

$$\mathcal{F}_{(n)}[F_{n-1}] = \mathcal{F}_{(n-1)}[F_{n-1}](2\pi)^d \delta(\boldsymbol{p}_n).$$
(3.199)

We are now ready to prove equation (3.196) by induction. The key point is the recursive structure in n of $\mathcal{F}_{(n)}[F_n]$. We already showed this property for $n \leq 4$ at the beginning of this section. We discuss it more systematically here, for any n. Let us consider the position-space function F_n and factorise it to display a recursive structure (see figure 3.3):

$$F_n = x_{1n}^{2\alpha_{1n}} x_{2n}^{2\alpha_{2n}} \dots x_{n-1,n}^{2\alpha_{n-1,n}} \times F_{n-1}.$$
(3.200)



Figure 3.3: Illustration of the recursive structure of F_n (3.200). The continuous lines denote F_{n-1} , and the dashed lines correspond to x_{jn}^2 (j = 1, ..., n - 1).

Then, we Fourier transform it by applying the convolution theorem:

$$\begin{aligned} \mathcal{F}_{(n)}[F_n] &= \mathcal{F}_{(n)}[x_{1n}^{2\alpha_{1n}}x_{2n}^{2\alpha_{2n}}\dots x_{n-1,n}^{2\alpha_{n-1,n}}] * \mathcal{F}_{(n)}[F_{n-1}] \\ &= \left[\frac{(2\pi)^d \delta\left(\sum_{j=1}^n \boldsymbol{p}_j\right) \prod_{i=1}^{n-1} C_{in}}{p_1^{2\alpha_{1n}+d} p_2^{2\alpha_{2n}+d} \dots p_{n-1}^{2\alpha_{n-1,n}+d}} \right] * \left[\mathcal{F}_{(n-1)}[F_{n-1}](2\pi)^d \delta(\boldsymbol{p}_n) (2\pi)^d \delta(\boldsymbol{p}_n) \right] \\ &= \prod_{i=1}^{n-1} C_{in} \int \frac{\mathrm{d}^d \boldsymbol{q}_i}{(2\pi)^d} \frac{\mathcal{F}_{(n-1)}[F_{n-1}](\boldsymbol{\alpha}; \boldsymbol{p}_1 - \boldsymbol{q}_1, \dots, \boldsymbol{p}_{n-1} - \boldsymbol{q}_{n-1})}{q_1^{2\alpha_{1n}+d} q_2^{2\alpha_{2n}+d} \dots q_{n-1}^{2\alpha_{n-1,n}+d}} (2\pi)^d \delta\left(\boldsymbol{p}_n + \sum_{j=1}^{n-1} \boldsymbol{q}_j\right) \end{aligned}$$
(3.201)

Now, first we rename $q_j \to q_{jn}$ (j = 1, ..., n-1). Then, taking into account that equation (3.196) holds for n = 2, equation (3.201) implies that (3.196) holds for all $n \ge 2$ by induction. This proves that the simplex integral (3.189) is conformally invariant when \hat{f} is a monomial in the momentum-space cross ratios. In fact, this is equivalent to shifting the α_{ij} by some amount and re-defining $\hat{f} = 1$, as we showed for n = 4. In other words, the simplex integral with monomial \hat{f} satisfies the CWIs just as the simplex with $\hat{f} = 1$.

Finally, to prove that the simplex integral is conformally invariant for any arbitrary function, we express \hat{f} as an inverse Mellin transform in identical fashion we did in equation (3.183) for n = 4.

Conformal invariance via CWIs

To directly prove that the simplex integral (3.189) satisfies the CWIs we proceed in an analogous way of n = 4. First, to show the scale invariance of (3.192), we need to verify that it scales as $\Delta_t - (n-1)d$. Each integration increases the dimension by d, while each propagator decreases the dimension by $2\alpha_{ij} + d$. Taking into account that the number of integrals is (n-1)(n-2)/2 and the number of propagators appearing in the denominator

(3.193) is (n-1)(n-2)/2 + (n-1), the total scaling is

$$\frac{1}{2}(n-1)(n-2)d - \left[\frac{1}{2}(n-1)(n-2) - (n-1)\right]d - \sum_{1 \le i < j \le n} 2\alpha_{ij} = \Delta_t - (n-1)d, \quad (3.202)$$

where we used (2.64). Hence, the DWI (3.1) is satisfied.

To prove that the simplex solves the SCWIs, one strategy is to show that the SCWI operator $\mathcal{K}^{\mu} = \sum_{j=1}^{n-1} \mathcal{K}^{\mu}_{j}$ acting on the integrand of (3.192) is equivalent to a sum of total derivatives with respect to q_{ij} . This has been shown in [104], here we cite the result:

$$\mathcal{K}^{\kappa}(\mathbf{\Delta})\left[\frac{\hat{f}(\hat{\boldsymbol{u}})}{\operatorname{Den}_{n}(\boldsymbol{\alpha})}\right] = \sum_{\substack{i,j=1\\i\neq j}}^{n-1} \frac{\partial}{\partial q_{ij}^{\mu}} \left[\Gamma_{ij}^{\kappa\mu}(\boldsymbol{\alpha})\hat{f}(\hat{\boldsymbol{u}}) + \sum_{I\in\mathcal{U}}\Gamma_{ij,I}^{\kappa\mu}(\boldsymbol{\alpha})\frac{\partial\hat{f}(\hat{\boldsymbol{u}})}{\partial\hat{u}_{I}}\right],\tag{3.203}$$

with

$$\Gamma_{ij}^{\kappa\mu}(\boldsymbol{\alpha}) = (2\alpha_{in} + d) \times \frac{A_{ij}^{\kappa\mu}}{\mathrm{Den}_n(\boldsymbol{\alpha})},$$
(3.204)

$$\Gamma_{ij,[pqrs]}^{\kappa\mu}(\boldsymbol{\alpha}) = 2(\delta_{ip}\delta_{rn} + \delta_{iq}\delta_{sn} - \delta_{ip}\delta_{qn} - \delta_{ir}\delta_{sn}) \times \frac{A_{ij}^{c\mu}\hat{u}_{[pqrs]}}{\mathrm{Den}_n(\boldsymbol{\alpha})},$$
(3.205)

and

$$A_{ij}^{\kappa\mu} = \left(\delta^{\kappa\mu}\delta_{\alpha\beta} + \delta^{\kappa}_{\beta}\delta^{\mu}_{\alpha} - \delta^{\kappa}_{\alpha}\delta^{\mu}_{\beta}\right) \frac{q_{ij}^{\alpha}(\boldsymbol{l}_i - \boldsymbol{p}_i)^{\beta}}{(\boldsymbol{l}_i - \boldsymbol{p}_i)^2}.$$
(3.206)

In the second part of this thesis, in Chapter 4, we will show in a more direct manner that the simplex solves the CWIs by using new scalar representations of the simplex.

3.7.3 4-point Ward identities: an example

In the previous section we showed that the general 4-point function depends on an arbitrary function \hat{f} of momentum-space cross ratios. This means that different 4-point functions associated to different \hat{f} exist. Certain classes of solutions for the 4-point CWIs are known, including Witten diagrams and free fields [30, 36, 29, 95, 105, 106]. The corresponding simplex representations of these solutions including the form of the function \hat{f} have been analysed in [36, 46]. Amongst the conformal integrals appearing in the second part of this thesis, we study contact and exchange Witten diagrams [107, 108]. Therefore, we conclude this chapter with an exercise: proving that these 4-point Witten diagrams solve the CWIs. Both diagrams consist of integrals of multiple Bessel functions.

First let us derive the 4-point CWIs in terms of the scalar variables p_1, p_2, p_3, p_4, s, t , with

$$s^{2} = (\boldsymbol{p}_{1} + \boldsymbol{p}_{2})^{2}, \qquad t^{2} = (\boldsymbol{p}_{2} + \boldsymbol{p}_{3})^{2}.$$
 (3.207)

Using the chain rule,

$$\frac{\partial}{\partial p_{1\mu}} = \frac{p_1^{\mu}}{p_1} \frac{\partial}{\partial p_1} + \frac{p_1^{\mu} + p_2^{\mu} + p_3^{\mu}}{p_4} \frac{\partial}{\partial p_4} + \frac{p_1^{\mu} + p_2^{\mu}}{s} \frac{\partial}{\partial s},$$

$$\frac{\partial}{\partial p_{2\mu}} = \frac{p_2^{\mu}}{p_2} \frac{\partial}{\partial p_2} + \frac{p_1^{\mu} + p_2^{\mu} + p_3^{\mu}}{p_4} \frac{\partial}{\partial p_4} + \frac{p_1^{\mu} + p_2^{\mu}}{s} \frac{\partial}{\partial s} + \frac{p_2^{\mu} + p_3^{\mu}}{t} \frac{\partial}{\partial t},$$

$$\frac{\partial}{\partial p_{3\mu}} = \frac{p_3^{\mu}}{p_3} \frac{\partial}{\partial p_3} + \frac{p_1^{\mu} + p_2^{\mu} + p_3^{\mu}}{p_4} \frac{\partial}{\partial p_4} + \frac{p_2^{\mu} + p_3^{\mu}}{t} \frac{\partial}{\partial t},$$
(3.208)

we obtain the DWI

$$0 = \left[-\Delta_t + 3d + \sum_{i=1}^4 p_i \partial_i + s \partial_s + t \partial_t \right] \left\langle \! \left\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2(\boldsymbol{p}_2) \mathcal{O}_3(\boldsymbol{p}_3) \mathcal{O}_4(\boldsymbol{p}_4) \right\rangle \! \right\rangle \tag{3.209}$$

and the SCWIs [29]

$$0 = \mathcal{D}_{ij} \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2(\boldsymbol{p}_2) \mathcal{O}_3(\boldsymbol{p}_3) \mathcal{O}_4(\boldsymbol{p}_4) \rangle\!\rangle, \quad i, j = 1, ..., 4,$$
(3.210)

where

$$\mathcal{D}_{12} = K_{12} + (L_1 - L_2 - L_3 + L_4) \frac{1}{t} \partial_t + (-p_3^2 + p_4^2) \frac{1}{st} \partial_s \partial_t, \qquad (3.211)$$

$$\mathcal{D}_{23} = K_{23} + (L_1 + L_2 - L_3 - L_4) \frac{1}{s} \partial_s + (p_1^2 - p_4^2) \frac{1}{st} \partial_s \partial_t, \qquad (3.212)$$

$$\mathcal{D}_{34} = K_{34} + (-L_1 + L_2 + L_3 - L_4)\frac{1}{t}\partial_t + (-p_1^2 + p_2^2)\frac{1}{st}\partial_s\partial_t, \qquad (3.213)$$

are the independent special conformal operators and we defined $L_i = p_i \partial_i - \Delta_i$. Now, let us assume we seek a solution to these equations that does not depend on s and t. Then the SCWIs simplify to

$$0 = \mathbf{K}_{ij} \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \mathcal{O}_2(\boldsymbol{p}_2) \mathcal{O}_3(\boldsymbol{p}_3) \mathcal{O}_4(\boldsymbol{p}_4) \rangle\!\rangle, \quad i, j = 1, ..., 4.$$
(3.214)

By comparing these equations with the 3-point SCWIs in (3.34), it becomes evident that they can also be solved by separation of variables, giving an integral of four Bessel-K functions. Since the DWI must also be satisfied, we find

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\mathcal{O}_4(\boldsymbol{p}_4)\rangle\!\rangle = c \int_0^\infty \mathrm{d}x \, x^{d-1} \prod_{j=1}^4 p_j^{\beta_j} K_{\beta_j}(p_j x).$$
(3.215)

This indeed coincides with the 4-point contact Witten diagram, $i_{[d; \Delta_1, \Delta_2; \Delta_3, \Delta_4;]}$, when

$$c = \left(\prod_{j=1}^{4} 2^{\beta_j - 1} \Gamma\left(\beta_j\right)\right)^{-1}.$$
 (3.216)

Let us now move to the 4-point s-channel exchange diagram. This is also an integral of multiple Bessel functions and it reads

$$i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = \int_0^\infty dz \, z^{-d-1} \mathcal{K}_{[\Delta_1]}(z,p_1) \mathcal{K}_{[\Delta_2]}(z,p_2)$$

$$\times \int_0^\infty d\zeta \, \zeta^{-d-1} \mathcal{G}_{[\Delta_x]}(z,s;\zeta) \mathcal{K}_{[\Delta_3]}(\zeta,p_3) \mathcal{K}_{[\Delta_4]}(\zeta,p_4),$$
(3.217)

where Δ_x is the dimension of the exchanged operator and $\mathcal{G}_{[\Delta_x]}$ denotes the bulk-to-bulk propagator

$$\mathcal{G}_{[\Delta_x]}(z,s;\zeta) = \begin{cases} (z\zeta)^{\frac{d}{2}} I_{\beta_x}(sz) K_{\beta_x}(s\zeta) & \text{for } z < \zeta, \\ (z\zeta)^{\frac{d}{2}} K_{\beta_x}(sz) I_{\beta_x}(s\zeta) & \text{for } z > \zeta, \end{cases}$$
(3.218)

with I_{β} and K_{β} representing modified Bessel functions and $\beta_x = \Delta_x - d/2$. We will not derive this result here, which is a result known in holographic CFT. Our goal here is, instead, to show that (3.217) solves the 4-point CWIs (3.209), (3.210). First, we note that s-channel exchange diagrams don't depend on t. Hence the SCWIs operators reduces to

$$\mathcal{D}_{12} = K_{12},$$

$$\mathcal{D}_{23} = K_{23} + (L_1 + L_2 - L_3 - L_4) \frac{1}{s} \partial_s,$$

$$\mathcal{D}_{34} = K_{34}.$$
(3.219)

Taking into account that $K_{12} = K_1 - K_2$, with K_i the Bessel operator in (3.40), by direct computation we find that the following SCWIs are satisfied

$$0 = K_{12} i_{[d; \Delta_1, \Delta_2; \Delta_3, \Delta_4; \Delta_x]} = K_{34} i_{[d; \Delta_1, \Delta_2; \Delta_3, \Delta_4; \Delta_x]}.$$
(3.220)

To prove the second SCWI involving $\mathcal{D}_{23} = 0$, we first introduce the Casimir operator. This is also useful to derive the action of the operator \mathcal{W}_{12}^{--} on both contact and exchange diagrams. The quadratic Casimir operator in momentum-space reads

$$\mathcal{C}_{12} = (\boldsymbol{p}_1 \cdot \boldsymbol{p}_2 \delta^{\mu\nu} - 2p_1^{\mu} p_2^{\nu}) \partial_{12}^{\mu} \partial_{12}^{\nu} + 2[(\Delta_1 - d)p_2^{\mu} - (\Delta_2 - d)p_1^{\mu}] \partial_{12,\mu} + (\Delta_1 + \Delta_2 - 2d)(\Delta_1 + \Delta_2 - d).$$
(3.221)

In Mandelstam variables, omitting terms involving derivatives with respect to t, this is

$$\mathcal{C}_{12} = \frac{1}{2}(s^2 + p_1^2 - p_2^2)K_1 + \frac{1}{2}(s^2 + p_2^2 - p_1^2)K_2 - (L_1 + L_2 + \frac{3d}{2})^2 + \frac{d^2}{4} + O(\partial_t). \quad (3.222)$$

Then the action of the Casimir operator on the exchange diagram amounts to the action of the following reduced operator [36]

$$\tilde{\mathcal{C}}_{12} = \frac{s^2}{2}(K_1 + K_2) - (L_1 + L_2 + \frac{3d}{2})^2 + \frac{d^2}{4}, \qquad (3.223)$$

since the exchange diagram satisfies $K_{12}i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = K_{34}i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = 0$. This reduced operator has the property that it sends an exchange diagram to a contact diagram as follows

$$(\mathcal{C}_{12} + m_x^2)i_{[d;\,\Delta_1,\Delta_2;\,\Delta_3,\Delta_4;\,\Delta_x]} = i_{[d;\,\Delta_1,\Delta_2;\,\Delta_3,\Delta_4]},\tag{3.224}$$

where $m_x^2 = \Delta_x(\Delta_x - d)$. We are now ready to prove that the *s*-channel exchange diagram satisfies the remaining SCWI $\mathcal{D}_{23i[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = 0$. Taking into account \mathcal{D}_{23} in (3.219) and using the DWI (dropping the derivative with respect to *t*) to eliminate the derivative with respect to s, we have

$$s^{2}\mathcal{D}_{23} = s^{2}K_{23} + (L_{1} + L_{2} - L_{3} - L_{4})(-3d - L_{1} - L_{2} - L_{3} - L_{4}), \qquad (3.225)$$

and by rearranging, we find

$$s^{2}\mathcal{D}_{23} = s^{2}K_{23} - (L_{1} + L_{2} + 3d/2)^{2} + (L_{3} + L_{4} + 3d/2)^{2} = \tilde{C}_{12} - \tilde{C}_{34}, \qquad (3.226)$$

where we used (3.223). Thus, when acting on the exchange diagram

$$s^{2} \mathcal{D}_{23} i_{[d;\Delta_{1},\Delta_{2};\Delta_{3},\Delta_{4};\Delta_{x}]} = (\tilde{C}_{12} - \tilde{C}_{34}) i_{[d;\Delta_{1},\Delta_{2};\Delta_{3},\Delta_{4};\Delta_{x}]} = 0, \qquad (3.227)$$

where in the last equality we used the action of the reduced Casimir operator (3.224) and took into account that \tilde{C}_{12} and \tilde{C}_{34} when acting on the exchange diagram give the same contact diagram, hence the action of their difference on the latter vanishes.

In section 3.6.2 we introduced the weight-shifting operators $\mathcal{W}_{12}^{\pm\pm}$ acting on any *n*-point functions. We computed the action of \mathcal{W}_{12}^{--} on the 3-point function (3.155), showing that it generates a shifted 3-point function. While the 3-point function is unique, 4-point functions are not. For instance, here we considered two types of 4-point functions, the contact and *s*-channel Witten diagrams. As a consequence, the action of the weightshifting operator $\mathcal{W}_{12}^{\pm\pm}$ on a 4-point function does not a priori generate the same function with shifted parameters. In fact, in [36] it has been shown that the operator $\mathcal{W}_{12}^{\pm\pm}$ acts on an exchange Witten diagram to generate a linear combination of a shifted exchange and a shifted contact diagrams, or equivalently a shifted exchange diagram but with derivative vertices. Hence, it does not generate the same function with shifted parameters. A natural question then arises: is there a weight-shifting operator that when acting on 4-point Witten diagrams preserves the form of the function and only shifts the parameters? We show the answer in Chapter 5.

3.8 Discussion

In this chapter we gave an overview of conformal field theory in momentum space, focusing on the scalar sector. We derived the *n*-point CWIs and discussed their solutions. We devoted considerable space to the 3-point function, presenting its equivalent representations and singularities. We constructed the shift operators \mathcal{L}_i and \mathcal{R}_i (3.117), which connect 3-point functions in spacetime dimensions *d* differing by two. Moreover, we presented the shift operators $\mathcal{W}_{ij}^{\pm\pm}$ that connect *n*-point functions with shifted scaling dimensions but same *d*. We then derived the general solutions of 4- and *n*-point CWIs that were found recently in terms of the simplex integral (3.189).

While various studies of *n*-point functions yielded special classes of solutions to the 4-point CWIs, the simplex integral provides the *general* solution. Several questions arise. Is there a scalar representation of the simplex integral that simplifies the study of *n*-point functions? What is the generalisation of the shift operators \mathcal{L}_i at *n* points? Is there a representation of the simplex integral that helps us to find this class of operators? Moreover, proving that the simplex integral satisfies the CWIs was cumbersome. Is there a representation that simplifies this computation? We address these questions in the next chapter, where we find new scalar parametrisations of the integral by using insights from

the physics of electrical circuits.

We concluded this chapter by showing that 4-point contact and exchange Witten diagrams solve the CWIs and quoted the action of the shift operators $\mathcal{W}_{ij}^{\pm\pm}$ on such solutions. Unlike the 3-point case, these operators act on 4-point functions connecting solutions of CWIs with shifted parameters but do not leave the form of the functions unchanged. Thus, further questions about shift operators arise. Is there a shift operator that when acting on Witten diagrams (and more generally on a certain integral) preserves the form of the integral while shifting the parameters? Does the inverse operator of \mathcal{L}_i exist? We discuss these questions and find a class of such operators in Chapter 5, using the formalism of a class of multivariable hypergeometric functions known as GKZ systems.

Part III

Integral representations and shift operators

Chapter 4

Shift operators from the simplex representation in momentum-space CFT

4.1 Introduction

Understanding the general form of correlation functions in momentum-space conformal field theory is an important goal. Working in momentum space is natural for many applications, particularly inflationary cosmology (see, *e.g.*, [19–32]), and reveals features inherited from scattering amplitudes that would otherwise be hidden, for example double-copy structure and colour/kinematics duality [37–41]. Momentum-space methods are moreover well suited for renormalisation [45, 34–36], and are of growing interest for the conformal bootstrap [109–111].

In position space, the structure of general scalar *n*-point functions has been understood for over fifty years [7]. A correspondingly general solution in momentum space was proposed only recently in [47, 46]. This takes the form of a generalised Feynman integral with the topology of an (n-1)-simplex,

$$\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle = \prod_{1 \le i < j \le n} \int \frac{\mathrm{d}^d \boldsymbol{q}_{ij}}{(2\pi)^d} \frac{f(\hat{\boldsymbol{q}})}{q_{ij}^{2\alpha_{ij}+d}} \prod_{k=1}^n (2\pi)^d \delta\Big(\boldsymbol{p}_k + \sum_{l=1}^n \boldsymbol{q}_{lk}\Big), \tag{4.1}$$

where the integration is taken over the internal momenta q_{ij} running between vertices of the simplex. Here $q_{ij} = -q_{ji}$ runs from vertex *i* to *j*, while the external momenta p_i enter only via momentum conservation as imposed by the delta function inserted at each vertex. Each propagator corresponds to an edge of the simplex, as illustrated in figure 4.1, and is raised to a power specified by the parameter α_{ij} . Together, these satisfy the constraints

$$\Delta_i = -\sum_{j=1}^n \alpha_{ij},\tag{4.2}$$

where Δ_i is the scaling dimension of the operator \mathcal{O}_i . To simplify the writing of such sums we define $\alpha_{ii} = 0$ and $\alpha_{ji} = \alpha_{ij}$. Euclidean signature will be assumed throughout.



Figure 4.1: Structure of the simplex integral, illustrated for the 5-point function.

The distinguishing feature of the simplex representation (4.1) is the presence of an *arbitrary function* $f(\hat{q})$ of the independent momentum-space cross ratios

$$\hat{q}_{[ijkl]} = \frac{q_{ij}^2 q_{kl}^2}{q_{ik}^2 q_{il}^2},\tag{4.3}$$

denoted collectively by the vector \hat{q} . As the simplex representation can be derived from the general position-space solution [47, 46], the number of independent cross ratios is the same in both cases, *i.e.*, n(n-3)/2 for $n \leq d+2$ and nd - (d+2)(d+1)/2 for n > d+2. For $n \geq 4$, the solution of the constraints (4.2) for the α_{ij} is not unique, but making a different choice simply multiplies $f(\hat{q})$ by a product of powers of the cross ratios (4.3). Since $f(\hat{q})$ is arbitrary, the solution of (4.2) chosen is therefore immaterial.

In this chapter, we explore scalar parametric representations of the simplex integral (4.1) obtained by integrating out the internal momenta. This offers several advantages:

- The original integral (4.1) features n(n-1)/2 d-dimensional loop integrations and we have (n-1) delta functions to help us, with one remaining behind to enforce overall momentum conservation. This leaves the equivalent of (n-1)(n-2)d/2scalar integrals to perform. In contrast, the parametrisations we derive feature fewer integrals: only n(n-1)/2 scalar parametric integrals, one for each edge of the simplex.
- By inverting the graph polynomials that arise, we construct novel weight-shifting operators connecting solutions of the conformal Ward identities in spacetime dimension d to new solutions in dimension d + 2. Remarkably, these operators have a determinantal structure based on the Cayley-Menger matrix familiar from distance geometry. In contrast, the well-known weight-shifting operators introduced in [83] preserve the spacetime dimension. Operators mapping $d \rightarrow d + 2$ are we believe

known only for 3-point functions, where their existence can be seen from the triple- K representation in momentum space [85]¹, and for 4-point conformal blocks in position space (the operator \mathcal{E}_+ in [74]). The new $d \to d+2$ operators we obtain can be viewed as a natural generalisation of the 3-point operators of [85] to arbitrary n-point correlators.

The plan of this chapter is as follows. In section 4.2, we show that all graph polynomials for the simplex integral (4.1) can be constructed from the corresponding Gram matrix. The standard parametric representations for Feynman integrals then follow. Alternatively, by regarding the Schwinger parameters as resistances in an electrical network, we can compute the *effective* resistances between all vertices of the simplex. This latter set of variables dramatically simplifies the structure of the Schwinger exponential. In section 4.3, we use these effective resistances to construct new $d \rightarrow d + 2$ shift operators for the general npoint function. The cases n = 3, 4 are discussed in detail, and we verify the action of all operators independently through computation of their intertwining relations with the conformal Ward identities. The actions of the *d*-preserving weight-shifting operators of [83] are also demonstrated from this scalar parametric perspective. In section 4.4, we prove that the new parametric representations indeed solve the conformal Ward identities. In contrast to the vectorial representation (4.1) (for which the Ward identities are analysed in [47, 46]), for the new scalar parametric representations the Ward identities can be verified directly without use of recursive arguments in the number of points n. As we show in section 4.5, the validity of the conformal Ward identities, as well as the action of the dpreserving weight-shifting operators, can also be seen from the position-space counterpart of the simplex. Section 4.6 concludes with a summary of results and open directions.

4.2 Parametric representations of the simplex

This section investigates scalar parametric representations for the simplex integral (4.1). In the following, we identify the necessary graph polynomials (section 4.2.1), standard parametric representations (section 4.2.2), and introduce new variables analogous to the effective resistances between nodes of the simplex (section 4.2.3). To re-formulate the simplex integral in these variables, we solve the inverse problem to express the original Schwinger parameters in terms of the effective resistances (section 4.2.4). The reparametrised integral, which will be the basis of our new shift operators, then follows (section 4.2.5).

4.2.1 Graph polynomials

Exponentiating all propagators via Schwinger parametrisation, the internal momenta can be integrated out reducing the simplex integral to various scalar parametrisations. The structure of the resulting Symanzik polynomials is clearest however when expressed in terms of the *inverse* of the usual variables. For this reason, we use the inverse Schwinger parametrisation

$$\frac{1}{q_{ij}^{2\alpha_{ij}+d}} = \frac{1}{\Gamma(\alpha_{ij}+d/2)} \int_0^\infty \mathrm{d}v_{ij} \, v_{ij}^{-d/2-\alpha_{ij}-1} e^{-q_{ij}^2/v_{ij}}.$$
(4.4)

¹These $d \rightarrow d+2$ operators also enable the construction of d-dimensional tensorial correlators [42, 34, 35].

The resulting polynomials \mathcal{U} and \mathcal{F} are then related to the standard Symanzik polynomials U and F by

$$\mathcal{U}(v_{ij}) = \left(\prod_{i< j}^{n} v_{ij}\right) U\left(\frac{1}{v_{ij}}\right), \qquad \mathcal{F}(v_{ij}) = \left(\prod_{i< j}^{n} v_{ij}\right) F\left(\frac{1}{v_{ij}}\right). \tag{4.5}$$

For the simplex, the structure of \mathcal{U} and \mathcal{F} can be expressed in terms of two matrices. The first is the $(n-1) \times (n-1)$ Gram matrix $G_{ij} = \mathbf{p}_i \cdot \mathbf{p}_j$. For our purposes, the most convenient parametrisation is

$$G_{ij} = \begin{cases} \sum_{k=1}^{n} V_{ik} & i = j \\ -V_{ij} & i \neq j \end{cases} \quad i, j = 1, \dots, n-1,$$
(4.6)

where

$$V_{ij} = \begin{cases} -\boldsymbol{p}_i \cdot \boldsymbol{p}_j & i \neq j \\ 0 & i = j \end{cases} \quad i, j = 1, \dots, n.$$

$$(4.7)$$

Here the V_{ij} provide a full set of n(n-1)/2 symmetric and independent Lorentz invariants. To write the diagonal entries in the Gram matrix, we used momentum conservation to express $p_i^2 = -\sum_{k\neq i}^n \boldsymbol{p}_i \cdot \boldsymbol{p}_k$. The second matrix is simply the image of the Gram matrix under the mapping $V_{ij} \rightarrow v_{ij}$, namely

$$g_{ij} = \begin{cases} \sum_{k=1}^{n} v_{ik} & i = j, \\ -v_{ij} & i \neq j, \end{cases} \quad i, j = 1, \dots, n-1.$$
(4.8)

Since the v_{ij} correspond to the edges of the simplex we define, as we did for the V_{ij} ,

$$v_{ij} = v_{ji}, \qquad v_{ii} = 0.$$
 (4.9)

As shown in appendix B.1, the graph polynomials are now

$$\mathcal{U} = |g|, \qquad \mathcal{F} = \operatorname{tr}(\operatorname{adj}(g) \cdot G), \qquad \frac{\mathcal{F}}{\mathcal{U}} = \operatorname{tr}(g^{-1} \cdot G), \qquad (4.10)$$

where $|g| = \det g$, $\operatorname{adj} g = |g| g^{-1}$ is the adjugate matrix and g^{-1} the inverse matrix. The derivation proceeds by expressing the delta functions of (4.1) in Fourier form and integrating out all internal momenta. Only after this step has been performed are the Fourier integrals for the delta functions then evaluated. As the Gram determinant |G| is proportional to the squared volume of the simplex spanned by the independent momenta, the polynomial \mathcal{U} describes the image of this squared volume under the mapping $V_{ij} \to v_{ij}$. Alternatively, by the matrix tree theorem (see *e.g.*, [65]), \mathcal{U} is the Kirchhoff polynomial encoding the sum of spanning trees on the simplex.

A second useful expression for \mathcal{F} can be derived from Jacobi's identity,

$$\partial_{v_{ij}}|g| = \operatorname{tr}(\operatorname{adj}(g) \cdot \partial_{v_{ij}}g), \qquad (4.11)$$
in combination with the relation

$$G_{ij} = \sum_{k
(4.12)$$

This last relation follows from the linearity of the G_{ij} in the V_{kl} , as we saw in (4.6), and the mapping of $G_{ij} \rightarrow g_{ij}$ under $V_{kl} \rightarrow v_{kl}$. The sums run over all k and l such that k < l, corresponding to all edges of the simplex. Substituting (4.12) into (4.10) then using (4.11),

$$\mathcal{F} = \sum_{i < j}^{n} \frac{\partial |g|}{\partial v_{ij}} V_{ij}, \qquad \frac{\mathcal{F}}{\mathcal{U}} = \sum_{i < j}^{n} \frac{\partial \ln |g|}{\partial v_{ij}} V_{ij}, \qquad (4.13)$$

or in terms of the raw momenta,

$$\mathcal{F} = -\sum_{i < j}^{n} \frac{\partial |g|}{\partial v_{ij}} \mathbf{p}_{i} \cdot \mathbf{p}_{j}, \qquad \frac{\mathcal{F}}{\mathcal{U}} = -\sum_{i < j}^{n} \frac{\partial \ln |g|}{\partial v_{ij}} \mathbf{p}_{i} \cdot \mathbf{p}_{j}.$$
(4.14)

4.2.2 Parametric representations of the *n*-point correlator

To express correlators compactly, we extract the overall delta function of momentum conservation as

$$\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle = \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle\!\rangle (2\pi)^d \delta(\sum_{i=1}^n \boldsymbol{p}_n).$$
 (4.15)

We also define an arbitrary function $f(\hat{v})$ whose arguments, denoted collectively by the vector \hat{v} , are the independent inverse Schwinger parameter cross ratios

$$v_{[ijkl]} = \frac{v_{ij}v_{kl}}{v_{ik}v_{jl}}.$$
(4.16)

The simplex integral (4.1) can now be written in a variety of standard forms using the polynomials \mathcal{U} and \mathcal{F} defined in (4.10) or (4.14). Among the most useful are:

1. Schwinger parametrisation:

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \Big(\prod_{i< j}^n \int_0^\infty \mathrm{d}v_{ij} \, v_{ij}^{-\alpha_{ij}-1}\Big) f(\hat{\boldsymbol{v}}) \, \mathcal{U}^{-d/2} e^{-\mathcal{F}/\mathcal{U}} \tag{4.17}$$

Here, the $v_{ij}^{-d/2}$ factors in (4.4) cancel with those associated with $\mathcal{U}^{-d/2}$ via (4.5).

2. Lee-Pomeransky parametrisation [112]:

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \Big(\prod_{i< j}^n \int_0^\infty \mathrm{d}v_{ij} \, v_{ij}^{-\alpha_{ij}-1}\Big) f(\hat{\boldsymbol{v}}) \, (\mathcal{U}+\mathcal{F})^{-d/2} \tag{4.18}$$

3. Feynman parametrisation:

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \Big(\prod_{i< j}^n \int_0^\infty \mathrm{d}v_{ij} \, v_{ij}^{-\alpha_{ij}-1}\Big)\delta\Big(1 - \sum_{i< j}^n \kappa_{ij} v_{ij}\Big)f(\hat{\boldsymbol{v}}) \,\mathcal{U}^{\omega-d/2}\mathcal{F}^{-\omega}$$

$$(4.19)$$

where $\omega = (n-1)d/2 + \sum_{i < j}^{n} \alpha_{ij} = (-\Delta_t + (n-1)d)/2$ and the constants $\kappa_{ij} \ge 0$ can be chosen arbitrarily provided they are not all zero.² If we choose all $\kappa_{ij} = 1$ then the integration region is a simplex in the space spanned by the v_{ij} . Alternatively, we can set a single κ_{ij} to unity and the rest to zero which trivialises one of the integrations at the cost of obscuring permutation invariance.

These representations are all equivalent up to numerical factors; for clarity, we have re-absorbed these into the arbitrary functions. For analysing the action of weight-shifting operators and verifying the conformal Ward identities, we will focus exclusively on the Schwinger parametrisation (4.17). Nevertheless, the Lee-Pomeransky representation (4.18) is well suited for studying the Landau singularities, as discussed in appendix B.3, and the Feynman parametrisation (4.19) has the virtue that one integral can be performed using the delta function.

Example: As a quick illustration, the 4-point function in Schwinger parametrisation is

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\mathcal{O}_2(\boldsymbol{p}_2)\mathcal{O}_3(\boldsymbol{p}_3)\mathcal{O}_4(\boldsymbol{p}_4)\rangle\!\rangle = \Big(\prod_{i< j}^4 \int_0^\infty \mathrm{d}v_{ij} \, v_{ij}^{-\alpha_{ij}-1}\Big) f\Big(\frac{v_{12}v_{34}}{v_{13}v_{24}}, \, \frac{v_{14}v_{23}}{v_{13}v_{24}}\Big) \,|g|^{-d/2} e^{-\mathrm{tr}\,(g^{-1}\cdot G)}$$
(4.20)

where $G_{ij} = \boldsymbol{p}_i \cdot \boldsymbol{p}_j$ is the 3 × 3 Gram matrix and g is its image

$$g = \begin{pmatrix} v_{12} + v_{13} + v_{14} & -v_{12} & -v_{13} \\ -v_{12} & v_{12} + v_{23} + v_{24} & -v_{23} \\ -v_{13} & -v_{23} & v_{13} + v_{23} + v_{34} \end{pmatrix}.$$
 (4.21)

The determinant is

$$|g| = v_{12}v_{13}v_{14} + v_{12}v_{14}v_{23} + v_{13}v_{14}v_{23} + v_{12}v_{13}v_{24} + v_{13}v_{14}v_{24} + v_{12}v_{23}v_{24} + v_{13}v_{23}v_{24} + v_{14}v_{23}v_{24} + v_{12}v_{13}v_{34} + v_{12}v_{14}v_{34} + v_{12}v_{23}v_{34} + v_{13}v_{23}v_{34} + v_{14}v_{23}v_{34} + v_{12}v_{24}v_{34} + v_{13}v_{24}v_{34} + v_{14}v_{24}v_{34}$$

$$(4.22)$$

and the equivalence of (4.10) and (4.14) can be verified directly.

² The Feynman parametrisation follows from the Schwinger parametrisation by setting $v_{ij} = y_{ij}/\sigma$ subject to the constraint $\sum_{i< j}^{n} \kappa_{ij} y_{ij} = 1$. The \mathcal{U} and \mathcal{F} are homogeneous polynomials of weights (n-1)and (n-2) respectively, meaning that $\mathcal{F}(v_{ij})/\mathcal{U}(v_{ij}) = \sigma \mathcal{F}(y_{ij})/\mathcal{U}(y_{ij})$ while the Jacobian can be evaluated as per appendix B of [46]. We then perform the scale integral over σ and relabel the $y_{ij} \to v_{ij}$.

4.2.3 The effective resistances

Thus far, we have expressed the Kirchhoff polynomial \mathcal{U} as the determinant of g, the image under $V_{ij} \rightarrow v_{ij}$ of the Gram matrix, where \mathbf{p}_n is eliminated using momentum conservation. However, since all vertices of the simplex are equivalent, \mathcal{U} ought also to be expressible in terms of the $n \times n$ matrix \tilde{g} corresponding to the image of the extended Gram matrix $\tilde{G}_{ij} = \mathbf{p}_i \cdot \mathbf{p}_j$ for $i, j = 1, \ldots, n$. This is simply the Laplacian matrix for the simplex:

$$\tilde{g}_{ij} = \begin{cases} \sum_{k=1}^{n} v_{ik}, & i = j, \\ -v_{ij}, & i \neq j, \end{cases} \quad i, j = 1, \dots, n.$$
(4.23)

As every row and column sum of the Laplacian matrix is zero its determinant vanishes identically, but its cofactors (*i.e.*, signed first minors) are in fact all equal to \mathcal{U} . To see this, consider the diagonal minor $|\tilde{g}^{(n,n)}|$ formed by deleting row n and column n then taking the determinant. Comparing with (4.8), we then see that $|\tilde{g}^{(n,n)}| = |g| = \mathcal{U}$. As any diagonal minor is equal to its cofactor, \mathcal{U} is likewise the (n, n) cofactor. However, by elementary row and column operations one can show that all cofactors of the Laplacian matrix are equal.³ Thus, every cofactor (and every diagonal minor) is equal to \mathcal{U} . Note this also confirms that our choice of eliminating p_n in section 4.2.1 was immaterial.

Let us now turn to an electrical analogy involving a simplicial network of resistors. Here, the Laplacian matrix naturally encodes the external current \mathcal{I}_i flowing into node *i*, since

$$\mathcal{I}_i = \sum_{j \neq i} v_{ij} (\mathcal{V}_i - \mathcal{V}_j) = \sum_{j \neq i} \tilde{g}_{ij} \mathcal{V}_j, \qquad (4.24)$$

where v_{ij} is the conductivity of the edge connecting nodes *i* and *j* and \mathcal{V}_j is the voltage of node *j*. Given this identification of the v_{ij} with the conductivities, a natural question to ask is what are the corresponding *effective resistances* between the nodes? From Kirchhoff, the effective resistance s_{ij} between nodes *i* and *j* is given by the ratio of minors [113, 114]

$$s_{ij} = \frac{|\tilde{g}^{(ij,ij)}|}{|\tilde{g}^{(j,j)}|},\tag{4.25}$$

where $|\tilde{g}^{(I,J)}|$ indicates the minor formed by deleting the set of rows I and columns J then taking the determinant. Thus, $|\tilde{g}^{(ij,ij)}|$ is the second minor formed by deleting rows i and j as well as columns i and j, while $|\tilde{g}^{(j,j)}|$ is the first minor corresponding to deleting row and column j. From (4.23), the element v_{ij} appears only in the row and columns (i, i), (i, j), (j, i) and (j, j) of \tilde{g} . Forming the first minor $|\tilde{g}^{(j,j)}|$ by deleting row and column j, v_{ij} then appears only once in the (i, i) position. The derivative $\partial |\tilde{g}^{(j,j)}|/\partial v_{ij}$ is thus equal

³For example, add one to every element of \tilde{g}_{ij} then add all rows to the first row, and all columns to the first column. The top left entry is now n^2 while all remaining entries of the first row and column are n. Taking the determinant, we first extract an overall factor of n from the top row, then subtract the new top row (whose leftmost entry is now n with all other entries one) from all the other rows. The resulting matrix has zeros in all entries of the first column apart from the top one which is n, and all entries other than those in the first row and column are \tilde{g}_{ij} (since we added one then subtracted one). The determinant of \tilde{g}_{ij} plus the all ones matrix is therefore n^2 times the (1, 1) cofactor of \tilde{g}_{ij} . Repeating the exercise for any other choice of row and column yields the same result with the corresponding cofactor, hence all cofactors are equal. Note this also shows that \mathcal{U} is n^{-2} times the determinant of the Laplacian plus the all-ones matrix.

to the second minor $|\tilde{g}^{(ij,ij)}|$ formed by additionally deleting row and column *i* in $|\tilde{g}^{(j,j)}|$. Since $|\tilde{g}^{(j,j)}| = |g|$ as above, we have

$$s_{ij} = \frac{\partial \ln |g|}{\partial v_{ij}}, \qquad \frac{\mathcal{F}}{\mathcal{U}} = -\sum_{i

$$(4.26)$$$$

where the second result follows immediately from (4.14). The Schwinger exponent in (4.17) thus encodes the effective resistances s_{ij} between all vertices. Moreover, both \mathcal{U} and \mathcal{F} have been related to minors of the Laplacian: \mathcal{U} is any diagonal first minor (or cofactor), while the coefficients of the \mathcal{F} polynomial correspond to the second minors: from (4.14), the coefficient of $V_{ij} = -\mathbf{p}_i \cdot \mathbf{p}_j$ (for i < j) is $\partial |g| / \partial v_{ij} = |\tilde{g}^{(ij,ij)}|$.

Earlier, we noted that $\mathcal{U} = |g|$ is proportional to the squared volume of the (n-1)simplex formed by the independent momenta under the map $V_{ij} \to v_{ij}$. By the same token, each coefficient $|\tilde{g}^{(ij,ij)}|$ of the \mathcal{F} polynomial thus corresponds to the image of $|\tilde{G}^{(ij,ij)}|$, the second minor of the extended Gram matrix. However, this minor is simply the determinant of the reduced Gram matrix formed from all the momenta apart from p_i and p_j . Thus, the coefficient of V_{ij} in the \mathcal{F} polynomial is proportional to the squared volume of the (n-2)-simplex, formed from all the momenta except for p_i and p_j , under the map $V_{ij} \to v_{ij}$. Similarly, the effective resistance s_{ij} is proportional to the ratio of the squared volume of this (n-2)-simplex to the squared volume of the full (n-1)-simplex.

4.2.4 Re-parametrising the simplex

The original Schwinger parametrisation (4.17) is complicated by the non-linear dependence of the exponent on the v_{ij} . As we saw in (4.26), however, the coefficients of the $V_{ij} = -\mathbf{p}_i \cdot \mathbf{p}_j$ in \mathcal{F}/\mathcal{U} are simply the effective resistances s_{ij} between nodes. The next step is thus to invert the relation (4.26) to find the v_{ij} in terms of the s_{ij} , *i.e.*, to express the conductivities in terms of the effective resistances. The simplex integral can then be fully re-parametrised in terms of the s_{ij} , with the linearity of the Schwinger exponent giving a Fourier-style duality between the V_{ij} and the s_{ij} . This duality means that all momentum derivatives acting on the simplex, and all momenta, can be trivially exchanged for operators constructed from the s_{ij} and derivatives $\partial/\partial s_{ij}$. The latter can then be integrated by parts. This strategy will repeatedly prove useful to us later.

We start by applying Jacobi's relation to further evaluate (4.26),

$$s_{ij} = \frac{1}{|g|} \frac{\partial |g|}{\partial v_{ij}} = \operatorname{tr} \left(g^{-1} \cdot \frac{\partial g}{\partial v_{ij}} \right) = \begin{cases} (g^{-1})_{ii} + (g^{-1})_{jj} - 2(g^{-1})_{ij}, & i < j < n \\ (g^{-1})_{ii}, & i < j = n \end{cases}$$
(4.27)

where the matrices $\partial g / \partial v_{ij}$ are easily evaluated from (4.8). Defining $s_{ii} = 0$ for convenience (as we similarly defined $v_{ii} = 0$) and re-arranging, we find

$$(g^{-1})_{ij} = \frac{1}{2}(s_{in} + s_{jn} - s_{ij}), \qquad i, j = 1, \dots, n-1$$
(4.28)

where the diagonal entries reduce to $(g^{-1})_{ii} = s_{in}$. Inverting this matrix will now give us back the matrix g, as defined in (4.8), but re-expressed in terms of the s_{ij} . The desired

expressions for the v_{ij} in terms of the s_{ij} can then be read off from the appropriate entries.

In fact, it is sufficient simply to know the determinant $|g^{-1}|$. For i < j < n, the (i, j) minor formed by deleting row i and column j of g^{-1} is $|(g^{-1})^{(i,j)}| = -(-1)^{i+j}\partial |g^{-1}|/\partial s_{ij}$, since from (4.28) s_{ij} appears (with coefficient minus one-half) only in the positions (i, j) and (j, i) of the symmetric matrix g^{-1} . The off-diagonal entries of the adjugate matrix are thus

$$\operatorname{adj}(g^{-1})_{ij} = (-1)^{i+j} |(g^{-1})^{(i,j)}| = -\frac{\partial |g^{-1}|}{\partial s_{ij}}, \qquad i < j < n$$
(4.29)

 \mathbf{SO}

$$v_{ij} = -g_{ij} = -\frac{1}{|g^{-1}|} \operatorname{adj}(g^{-1})_{ij} = \frac{\partial \ln |g^{-1}|}{\partial s_{ij}}, \qquad i < j < n.$$
(4.30)

Similarly, s_{in} appears in every entry of the i^{th} row of g^{-1} , and in every entry of the i^{th} column. The coefficients for the off-diagonal entries are all one-half, while that for the diagonal entry is one. The derivative $\partial |g^{-1}|/\partial s_{in}$ then corresponds to summing one-half times the signed minors both along the i^{th} row and down the i^{th} column such that the diagonal entry is counted twice. As g^{-1} is symmetric, however, these two sums are equal so we can simply sum along the i^{th} row only with coefficient one. This gives

$$\frac{\partial \ln |g^{-1}|}{\partial s_{in}} = \sum_{j=1}^{n-1} \frac{(-1)^{i+j}}{|g^{-1}|} |(g^{-1})^{(i,j)}| = \sum_{j=1}^{n-1} \frac{1}{|g^{-1}|} \operatorname{adj}(g^{-1})_{ij} = \sum_{j=1}^{n-1} g_{ij} = v_{in}, \quad (4.31)$$

where in the final step we used (4.8) to identify the sum of the first n-1 entries along the i^{th} row of the Laplacian as v_{in} . The relation (4.30) thus holds for all $i < j \le n$.

To simplify this formula further, we observe that $|g^{-1}|$ can be re-expressed in terms of the determinant of the $(n+1) \times (n+1)$ Cayley-Menger matrix,

$$m = \begin{pmatrix} 0 & s_{12} & s_{13} & \dots & s_{1n} & 1\\ s_{12} & 0 & s_{23} & \dots & s_{2n} & 1\\ s_{13} & s_{23} & 0 & \dots & s_{3n} & 1\\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots\\ s_{1n} & s_{2n} & s_{3n} & \dots & 0 & 1\\ 1 & 1 & 1 & \dots & 1 & 0 \end{pmatrix}.$$
(4.32)

When evaluating the determinant, if we subtract the n^{th} column from the first (n-1) columns, and then the n^{th} row from the first (n-1) rows, we find

$$|m| = \begin{vmatrix} -2s_{1n} & s_{12} - s_{1n} - s_{2n} & s_{13} - s_{1n} - s_{3n} & \dots & s_{1n} & 0 \\ s_{12} - s_{1n} - s_{2n} & -2s_{2n} & s_{23} - s_{2n} - s_{3n} & \dots & s_{2n} & 0 \\ s_{13} - s_{1n} - s_{3n} & s_{23} - s_{2n} - s_{3n} & -2s_{3n} & \dots & s_{3n} & 0 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ s_{1n} & s_{2n} & s_{3n} & \dots & 0 & 1 \\ 0 & 0 & 0 & \dots & 1 & 0 \end{vmatrix}$$
(4.33)

Comparing with (4.28), the upper-left $(n-1) \times (n-1)$ submatrix is $-2g^{-1}$. Laplace

expanding along the $(n+1)^{\text{th}}$ row and then the $(n+1)^{\text{th}}$ column thus gives

$$|m| = -(-2)^{n-1}|g^{-1}|. (4.34)$$

Equations (4.30) and (4.31) can now be cleanly re-expressed in terms of the Cayley-Menger determinant:

$$v_{ij} = \frac{\partial \ln |m|}{\partial s_{ij}}, \qquad i < j \le n.$$
 (4.35)

This is our desired result expressing all the v_{ij} in terms of the s_{ij} , inverting (4.26). A few additional relations also follow. Jacobi's relation allows us to write

$$v_{ij} = \frac{1}{|m|} \frac{\partial |m|}{\partial s_{ij}} = \frac{1}{|m|} \operatorname{tr}\left(\operatorname{adj}(m) \cdot \frac{\partial m}{\partial s_{ij}}\right) = \operatorname{tr}\left(m^{-1} \cdot \frac{\partial m}{\partial s_{ij}}\right) = 2(m^{-1})_{ij}, \quad (4.36)$$

since $\partial m_{kl}/\partial s_{ij} = 2\delta_{i(k}\delta_{l)j}$ from (4.32). As the off-diagonal entries of the Laplacian matrix are $\tilde{g}_{ij} = -v_{ij}$, this means that

$$\tilde{g}_{ij} = -2(m^{-1})_{ij}, \quad i, j \le n.$$
(4.37)

In fact, as indicated, this equation also holds for the diagonal elements with $i = j \leq n$, since if we multiply the (n + 1)th row of m by column i of m^{-1} we find

$$0 = \sum_{j=1}^{n} m_{ij}^{-1}, \qquad i \le n \tag{4.38}$$

and since all row and column sums of the Laplacian matrix vanish,

$$\tilde{g}_{ii} = -\sum_{j \neq i}^{n} \tilde{g}_{ij} = \sum_{j \neq i}^{n} 2(m^{-1})_{ij} = -2(m^{-1})_{ii}.$$
(4.39)

Thus, the $n \times n$ upper-left submatrix of the inverse Cayley-Menger matrix is minus onehalf the Laplacian matrix, using either (4.26) or (4.35) to convert between the s_{ij} and v_{ij} .⁴

The appearance of the Cayley-Menger matrix in our analysis is not a total surprise: in Euclidean distance geometry, the Cayley-Menger determinant is proportional to the squared volume of the simplex whose squared side lengths are given by the s_{ij} . Here, the map $v_{ij} \rightarrow V_{ij}$ sends $(g^{-1})_{ij}$ to the inverse Gram matrix $G_{ij}^{-1} = (\tilde{p}_i \cdot \tilde{p}_j)$, which is itself the Gram matrix formed from the independent *dual* momentum vectors $\tilde{p}_i = G_{ij}^{-1} p_j$ satisfying $\tilde{p}_i \cdot p_j = \delta_{ij}$. The determinant $|g^{-1}|$ is thus proportional to the squared volume of the dual (n-1)-simplex spanned by the independent \tilde{p}_i , and by (4.34), the s_{ij} are then the squared side lengths of this dual simplex. This provides an alternative (dual) geometrical interpretation for the s_{ij} , besides the volume ratio discussed at the end of section 4.2.3.

⁴ To the best of our knowledge, this result, along with a geometrical interpretation of the remaining $(n+1)^{\text{th}}$ row and column of the Cayley-Menger inverse, was first obtained by Fiedler, see [115, 116].

Example: All the relations above are easily checked for small values of n, and the s_{ij} are always rational functions of the v_{ij} and vice versa. For the 4-point function, we find, *e.g.*,

$$s_{12} = \frac{\partial \ln |g|}{\partial v_{12}} = (v_{13}v_{14} + v_{14}v_{23} + v_{13}v_{24} + v_{23}v_{24} + v_{13}v_{34} + v_{14}v_{34} + v_{23}v_{34} + v_{24}v_{34})|g|^{-1}$$
$$v_{12} = \frac{\partial \ln |m|}{\partial s_{12}} = (-2s_{13}s_{23} + 2s_{14}s_{23} + 2s_{13}s_{24} - 2s_{14}s_{24} - 4s_{12}s_{34} + 2s_{13}s_{34} + 2s_{14}s_{34} + 2s_{23}s_{34} + 2s_{24}s_{34} - 2s_{3}^{2})|m|^{-1}, \qquad (4.40)$$

where |g| was evaluated in (4.22) and

$$|m| = \begin{vmatrix} 0 & s_{12} & s_{13} & s_{14} & 1 \\ s_{12} & 0 & s_{23} & s_{24} & 1 \\ s_{13} & s_{23} & 0 & s_{34} & 1 \\ s_{14} & s_{24} & s_{34} & 0 & 1 \\ 1 & 1 & 1 & 1 & 0 \end{vmatrix}$$
(4.41)

evaluates to

$$|m| = -2s_{12}s_{13}s_{23} + 2s_{12}s_{14}s_{23} + 2s_{13}s_{14}s_{23} - 2s_{14}^2s_{23} - 2s_{14}s_{23}^2 + 2s_{12}s_{13}s_{24} - 2s_{13}^2s_{24} - 2s_{13}s_{14}s_{24} + 2s_{13}s_{23}s_{24} + 2s_{14}s_{23}s_{24} - 2s_{13}s_{24}^2 - 2s_{12}^2s_{34} + 2s_{12}s_{13}s_{34} + 2s_{12}s_{14}s_{34} - 2s_{13}s_{14}s_{34} + 2s_{12}s_{23}s_{34} + 2s_{12}s_{23}s_{34} + 2s_{12}s_{23}s_{34} + 2s_{12}s_{23}s_{34} + 2s_{12}s_{23}s_{24} - 2s_{12}s_{23}s_{34} + 2s_{12}s_{2$$

4.2.5 Cayley-Menger parametrisation of the *n*-point correlator

Using the results above, we can re-express the various parametrisations of the simplex integral in terms of the effective resistances s_{ij} . If we write the external momenta in Cayley-Menger form,

$$M = m \Big|_{s_{ij} \to \boldsymbol{p}_i \cdot \boldsymbol{p}_j} \tag{4.43}$$

the Schwinger exponent can be written as

$$-\frac{\mathcal{F}}{\mathcal{U}} = \sum_{i$$

where the constant term n just produces an overall scaling which can be re-absorbed into the arbitrary function of cross-ratios. Moreover, as shown in appendix B.2.1, the determinant of the Jacobian is

$$\left|\frac{\partial s}{\partial v}\right| = \left|\frac{\partial^2 \ln |g|}{\partial v \partial v}\right| \propto |g|^{-n} \propto |m|^n, \tag{4.45}$$

where the constant of proportionality can again be absorbed into the arbitrary function.

The Schwinger form (4.17) of the simplex integral now becomes

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \Big(\prod_{i< j}^n \int_0^\infty \mathrm{d}s_{ij} \left(\frac{\partial \ln|m|}{\partial s_{ij}}\right)^{-\alpha_{ij}-1} \Big) f(\hat{\boldsymbol{v}}) \,|m|^{d/2-n} \, e^{\frac{1}{2}\mathrm{tr}(M\cdot m)} \quad (4.46)$$

where the cross-ratios \hat{v} are rational functions of the s_{ij} as defined via (4.16) and (4.35). An alternative expression can be given in terms of the Cayley-Menger minors, since from Jacobi's relation

$$\frac{\partial |m|}{\partial s_{ij}} = 2(-1)^{i+j} |m^{(i,j)}|, \qquad (4.47)$$

where $|m^{(i,j)}|$ is the minor formed by taking the determinant after deleting row *i* and column *j*. After absorbing numerical factors into the arbitrary function, this gives

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \left(\prod_{i< j}^n \int_0^\infty \mathrm{d}s_{ij} \,|m^{(i,j)}|^{-\alpha_{ij}-1}\right) f(\hat{\boldsymbol{v}}) \,|m|^\alpha \, e^{\frac{1}{2}\mathrm{tr}(M\cdot m)} \tag{4.48}$$

where

$$\alpha = \frac{1}{2} \Big(d + n(n-3) - \sum_{i=1}^{n} \Delta_i \Big), \qquad \hat{\boldsymbol{v}}_{[ijkl]} = \frac{v_{ij}v_{kl}}{v_{ik}v_{jl}} = \frac{|\boldsymbol{m}^{(i,j)}||\boldsymbol{m}^{(k,l)}|}{|\boldsymbol{m}^{(i,k)}||\boldsymbol{m}^{(j,l)}|}.$$
(4.49)

Analogous expressions can be obtained for the Lee-Pomeransky and Feynman representations (4.18) and (4.19), but the Schwinger parametrisations (4.46) and (4.48) are particularly convenient. As noted, the diagonal Schwinger exponent means differential operators in the momenta can easily be traded for equivalent differential operators in the s_{ij} acting on the exponential, whose action can be further evaluated through integration by parts.

4.3 Weight-shifting operators

New weight-shifting operators now follow from the Cayley-Menger parametrisation (4.48). Acting on the Schwinger exponent (4.44) with an appropriate polynomial differential operator in the momenta pulls down a corresponding polynomial in the s_{ij} . Choosing these polynomials to be the Cayley-Menger determinant and its minors, we obtain shift operators either increasing α or decreasing one of the α_{ij} by integer units. We discuss these new operators in section 4.3.1, showing their effect is to increase the spacetime dimension by two while performing assorted shifts of the operator dimensions. Further weight-shifting operators can then be constructed by conjugating these operators with shadow transforms as shown in section 4.3.2. Explicit examples are given for the 3- and 4-point functions in section 4.3.3. Then, in section 4.3.4, we turn to analyse the weight-shifting operators proposed in [83]. These preserve the spacetime dimension but their action can nevertheless be understood using our parametric representations.

4.3.1 New operators sending $d \rightarrow d+2$

Let us begin with the V_{ij} defined in (4.7) as our independent momentum variables. Acting on the Schwinger exponent (4.44), for any i < j

$$-\frac{\partial}{\partial V_{ij}}e^{\frac{1}{2}\operatorname{tr}(M\cdot m)} = s_{ij}\,e^{\frac{1}{2}\operatorname{tr}(M\cdot m)}, \qquad -V_{ij}\,e^{\frac{1}{2}\operatorname{tr}(M\cdot m)} = \frac{\partial}{\partial s_{ij}}e^{\frac{1}{2}\operatorname{tr}(M\cdot m)} \tag{4.50}$$

allowing differential operators in the momenta to be traded for equivalent operators in the integration variables s_{ij} . The shift operators

$$\mathcal{S}_{ij}^{++} = |m^{(i,j)}|\Big|_{s_{ij} \to -\partial/\partial V_{ij}}, \qquad \mathcal{S} = |m|\Big|_{s_{ij} \to -\partial/\partial V_{ij}}, \qquad (4.51)$$

then serve to pull down factors of $|m^{(i,j)}|$ and |m| respectively, thus their action is to send

$$\mathcal{S}_{ij}^{++}: \quad \alpha_{ij} \to \alpha_{ij} - 1, \qquad \mathcal{S}: \quad \alpha \to \alpha + 1.$$
 (4.52)

From (4.2) and (4.49), this is equivalent to shifting

$$\mathcal{S}_{ij}^{++}: d \to d+2, \quad \Delta_i \to \Delta_i+1, \quad \Delta_j \to \Delta_j+1,$$

$$(4.53)$$

$$\mathcal{S}: \quad d \to d+2, \tag{4.54}$$

and so the superscript on \mathcal{S}_{ij}^{++} is chosen to indicate its action of raising Δ_i and Δ_j by one.

While the Cayley-Menger structure of \mathcal{S}_{ij}^{++} and \mathcal{S} is manifest in the variables V_{ij} , where convenient these operators can easily be rewritten in terms of other scalar invariants (*e.g.*, Mandelstam variables) via the chain rule. We will discuss this for 3- and 4-point functions shortly in section 4.3.3.

Alternatively, we can express S_{ij}^{++} and S in terms of *vectorial* derivatives with respect to independent momentum p_i for i = 1, ..., n - 1. For S, we find

$$S = -\frac{(n-1)!}{|G|} p_1^{[\mu_1} \dots p_{n-1}^{\mu_{n-1}]} \frac{\partial}{\partial p_1^{\mu_1}} \dots \frac{\partial}{\partial p_{n-1}^{\mu_{n-1}}}$$
(4.55)

where $|G| = |\mathbf{p}_i \cdot \mathbf{p}_j|$ is the Gram determinant and the μ_i are Lorentz indices. (We leave all Lorentz indices upstairs to avoid confusion with the momentum labels, given we are working on a flat Euclidean metric.) The equivalence of (4.55) to (4.51) can be established either by direct calculation for specific n, or else by considering its action on the Schwinger exponential of the representation (4.17). This representation is the appropriate one since, from (4.10), it involves only dot products of the *independent* momenta. Evaluating, we find

$$S\left(e^{-\sum_{i,j}^{n-1}(g^{-1})_{ij}\boldsymbol{p}_{i}\cdot\boldsymbol{p}_{j}}\right) = -\frac{(n-1)!}{|G|} p_{1}^{[\mu_{1}} \dots p_{n-1}^{\mu_{n-1}]}$$

$$\times \left(-2\sum_{j_{1}}^{n-1} (g^{-1})_{1j_{1}} p_{j_{1}}^{\mu_{1}} \right) \dots \left(-2\sum_{j_{n-1}}^{n-1} (g^{-1})_{n-1,j_{n-1}} p_{j_{n-1}}^{\mu_{n-1}} \right) e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}}$$

$$= \frac{-(-2)^{n-1} (n-1)!}{|G|} \sum_{j_{1},k_{1}}^{n-1} \dots \sum_{j_{n-1},k_{n-1}}^{n-1} (g^{-1})_{1j_{1}} \dots (g^{-1})_{n-1,j_{n-1}} (\boldsymbol{p}_{1} \cdot \boldsymbol{p}_{k_{1}}) \dots (\boldsymbol{p}_{n-1} \cdot \boldsymbol{p}_{k_{n-1}})$$

$$\times \delta_{j_{1}}^{[k_{1}} \dots \delta_{j_{n-1}}^{k_{n-1}]} e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}}$$

$$= -(-2)^{n-1} |g|^{-1} e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}}$$

$$= |m| e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}},$$

$$(4.56)$$

where in the last step we used the Levi-Civita identity $(n-1)! \, \delta_{j_1}^{[k_1} \dots \delta_{j_{n-1}}^{k_{n-1}]} = \varepsilon^{j_1 \dots j_{n-1}} \varepsilon_{k_1 \dots k_{n-1}}$ to generate a product of determinants $|g^{-1}||G|$, with the |G| then cancelling. Referring back to (4.17), since $\mathcal{U}^{-d/2} = |g|^{-d/2}$ we see the action of \mathcal{S} is thus indeed to raise $d \to d+2$. Through similar manipulations, we find

Through similar manipulations, we find

$$\mathcal{S}_{in}^{++} = (-1)^{i+n} \, \frac{(n-1)!}{|G|} \, p_1^{[\mu_1} \dots p_{n-1}^{\mu_{n-1}]} p_n^{\mu_i} \prod_{k \neq i}^{n-1} \frac{\partial}{\partial p_k^{\mu_k}}.$$
(4.57)

Relative to (4.55), the derivative $\partial/\partial p_i^{\mu_i}$ has been replaced by the dependent momentum $p_n^{\mu_i} = -\sum_{j_i=1}^{n-1} p_{j_i}^{\mu_i}$ positioned to the left of all derivatives. This leads to

$$\begin{aligned} \mathcal{S}_{in}^{++} \left(e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}} \right) \\ &= \frac{(-1)^{i+n} (n-1)!}{|G|} p_{1}^{[\mu_{1}} \dots p_{n-1}^{\mu_{n-1}]} \\ &\times \left(-2\sum_{j_{1}}^{n-1} (g^{-1})_{1j_{1}} p_{j_{1}}^{\mu_{1}} \right) \dots \left(-2\sum_{j_{i}}^{n-1} p_{j_{i}}^{\mu_{i}} \right) \dots \left(-2\sum_{j_{n-1}}^{n-1} (g^{-1})_{n-1,j_{n-1}} p_{j_{n-1}}^{\mu_{n-1}} \right) e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}} \\ &= (-1)^{i} 2^{n-2} \sum_{j_{i}=1}^{n-1} \frac{\partial |g^{-1}|}{\partial (g^{-1})_{ij_{i}}} e^{-\sum_{i,j}^{n-1} (g^{-1})_{ij} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j}}, \end{aligned}$$

$$(4.58)$$

since, relative to our previous calculation, the matrix element $(g^{-1})_{ij_i}$ is missing in the product on the middle line. Instead of obtaining the full determinant $|g^{-1}|$, we then get the derivative of this with respect to the missing element. As in (4.31), we can now rewrite

$$\sum_{j_i=1}^{n-1} \frac{\partial |g^{-1}|}{\partial (g^{-1})_{ij_i}} = \sum_{j_i=1}^{n-1} (\operatorname{adj} g^{-1})_{ij_i} = \sum_{j_i=1}^{n-1} g_{ij_i} |g^{-1}| = v_{in} |g|^{-1} = (-1)^i \, 2^{2-n} |m^{(i,n)}| \quad (4.59)$$

using (4.47) in the last step. The action of S_{in}^{++} in (4.57) on the exponential is thus to pull down a factor of $v_{ij}|g|^{-1}$. From the representation (4.17), this has precisely the required action of sending $\alpha_{ij} \to \alpha_{ij} - 1$ and $d \to d + 2$. Finally, since the choice of dependent momentum is immaterial, (4.57) generalises to

$$\mathcal{S}_{ij}^{++} = (-1)^{i+j} \, \frac{(n-1)!}{|G|} \, p_1^{[\mu_1} \dots \hat{p}_j^{\hat{\mu}_j} \dots p_n^{\mu_n]} p_j^{\mu_i} \prod_{k \neq i,j}^n \frac{\partial}{\partial p_k^{\mu_k}}$$
(4.60)

where the hats $\hat{p}_{j}^{\hat{\mu}_{j}}$ indicates that this factor and index are omitted in the antisymmetrised product, and we take $p_{j}^{\mu_{j}} = -\sum_{k_{j}\neq j}^{n} p_{k_{j}}^{\mu_{j}}$ as the dependent momentum. In principle these last few derivations allow use of the s_{ij} variables to be avoided entirely, although in practice the form of the operators (4.55) and (4.60) would be hard to anticipate.

4.3.2 Further shift operators from shadow conjugation

Additional $d \to d + 2$ shift operators can now be constructed – at no expense – by conjugating S_{ij}^{++} and S by a pair of shadow transforms. This idea was discussed recently for *d*-preserving weight-shifting operators in [36].

In momentum space, the shadow transform $\Delta_i \to d - \Delta_i$ (leaving d invariant) simply corresponds to multiplying by $p_i^{d-2\Delta_i}$. First, notice that attempting to conjugate \mathcal{S}_{ij}^{++} by shadow transforms on either of Δ_i or Δ_j has no effect: for example, the action of the operator $p_i^{2\Delta_i-d}\mathcal{S}_{ij}^{++}p_i^{d-2\Delta_i}$ corresponds to the successive parameter shifts

$$(\Delta_{i}, \Delta_{j}, d) \xrightarrow{p_{i}^{d-2\Delta_{i}}} (d - \Delta_{i}, \Delta_{j}, d) \\ \xrightarrow{S_{ij}^{++}} (d - \Delta_{i} + 1, \Delta_{j} + 1, d + 2) \\ \xrightarrow{p_{i}^{(d+2)-2(d-\Delta_{1}+1)} = p_{i}^{2\Delta_{i}-d}}} ((d+2) - (d - \Delta_{i} + 1), \Delta_{j} + 1, d + 2) \\ = (\Delta_{i} + 1, \Delta_{j} + 1, d + 2)$$
(4.61)

which is equivalent to the action of S_{ij}^{++} alone. Further computations confirm that the shadow transform on Δ_i or Δ_j commutes with S_{ij}^{++} .

However, we do obtain new operators if we shadow conjugate S_{ij}^{++} on any index $k \neq i, j$. For example, the action of

$$p_k^{2\Delta_k+2-d}\mathcal{S}_{ij}^{++}p_k^{d-2\Delta_k} \tag{4.62}$$

corresponds to the successive parameter shifts

$$(\Delta_i, \Delta_j, \Delta_k, d) \xrightarrow{p_k^{d-2\Delta_k}} (\Delta_i, \Delta_j, d - \Delta_k, d)$$

$$\xrightarrow{\mathcal{S}_{ij}^{i++}} (\Delta_i + 1, \Delta_j + 1, d - \Delta_k, d + 2)$$

$$\xrightarrow{p_k^{(d+2)-2(d-\Delta_k)} = p_k^{2\Delta_k+2-d}} (\Delta_i + 1, \Delta_j + 1, \Delta_k + 2, d + 2).$$
(4.63)

Thus, in addition to the shifts produced by S_{ij}^{++} alone, we have also shifted Δ_k up by two. Shadow conjugating on further variables has the same effect, for example,

$$p_k^{2\Delta_k+2-d} p_l^{2\Delta_l+2-d} \mathcal{S}_{ij}^{++} p_k^{d-2\Delta_k} p_l^{d-2\Delta_l}$$
(4.64)

for any $(k, l) \neq (i, j)$ sends $(\Delta_i, \Delta_j, \Delta_k, \Delta_l, d) \rightarrow (\Delta_i + 1, \Delta_j + 1, \Delta_k + 2, \Delta_l + 2, d + 2)$. We can also apply similar considerations to S. The action of

$$p_i^{2\Delta_i+2-d} \mathcal{S} p_i^{d-2\Delta_i} \tag{4.65}$$

corresponds to the shifts

$$(\Delta_i, d) \xrightarrow{p_i^{d-2\Delta_i}} (d - \Delta_i, d) \xrightarrow{\mathcal{S}} (d - \Delta_i, d+2) \xrightarrow{p_i^{(d+2)-2(d-\Delta_i)} = p_i^{2\Delta_i + 2-d}} (\Delta_i + 2, d+2).$$

$$(4.66)$$

Shadow conjugating on further momenta p_k leads similarly to shifting $\Delta_k \to \Delta_k + 2$.

With all these operators obtained by shadow conjugation, notice we can always obtain an equivalent differential operator with purely polynomial coefficients (*i.e.*, an operator in the Weyl algebra) by commuting the inner $p_k^{d-2\Delta_k}$ shadow factors through the differential operator S or S_{ij}^{++} , whereupon all non-integer powers cancel with those from the outer shadow transform.

4.3.3 Examples at three and four points

To illustrate the general discussion in the two preceding subsections, let us now compute the explicit form of these $d \rightarrow d + 2$ shift operators for 3- and 4-point functions.

3-point shift operators

For the 3-point function, it is convenient to use the three squared momentum magnitudes as variables. Defining

$$P_i = p_i^2, \qquad D_i = \frac{\partial}{\partial P_i} = \frac{1}{2p_i} \frac{\partial}{\partial p_i}, \qquad i = 1, 2, 3$$

$$(4.67)$$

via momentum conservation we have

$$P_{i} = -\sum_{j \neq i}^{3} \boldsymbol{p}_{i} \cdot \boldsymbol{p}_{j} = \sum_{j \neq i}^{3} V_{ij}, \qquad \frac{\partial}{\partial V_{ij}} = \sum_{k=1}^{3} \frac{\partial P_{k}}{\partial V_{ij}} \frac{\partial}{\partial P_{k}} = D_{i} + D_{j}$$
(4.68)

From (4.51), writing $D_i D_j = D_{ij}$ for short, we then find

$$S = -4(D_{12} + D_{23} + D_{13}),$$

$$S_{12}^{++} = -2D_3, \quad S_{23}^{++} = -2D_1, \quad S_{13}^{++} = 2D_2.$$
(4.69)

The various signs on the second line reflect our choice to use the Cayley minors in (4.48) and (4.51): had we used instead the cofactors or $\partial |m|/\partial s_{ij}$ as per (4.47) then all signs would be the same. Generally, any overall coefficient in S or the S_{ij}^{++} can be eliminated by rescaling the corresponding prefactor in the definition of the simplex integral.

As noted in the introduction, these 3-point operators (and their shadow conjugates) are already known from the triple-K representation of the 3-point function. In [42, 85],

the Bessel shift operators

$$\mathcal{L}_{i} = -\frac{1}{p_{i}}\frac{\partial}{\partial p_{i}}, \qquad \mathcal{R}_{i} = 2\beta_{i} - p_{i}\frac{\partial}{\partial p_{i}} = p_{i}^{2\beta_{i}+2}\mathcal{L}_{i}p_{i}^{-2\beta_{i}}, \qquad \beta_{i} = \Delta_{i} - \frac{d}{2}$$
(4.70)

where shown to act on the 3-point function by sending

$$\mathcal{L}_i: \quad \beta_i \to \beta_i - 1, \quad d \to d + 2, \qquad \mathcal{R}_i: \quad \beta_i \to \beta_i + 1, \quad d \to d + 2, \tag{4.71}$$

or equivalently,

$$\mathcal{L}_1: \quad (d, \Delta_1, \Delta_2, \Delta_3) \to (d+2, \Delta_1, \Delta_2+1, \Delta_3+1), \tag{4.72}$$

$$\mathcal{R}_1: \quad (d, \Delta_1, \Delta_2, \Delta_3) \to (d+2, \Delta_1+2, \Delta_2+1, \Delta_3+1), \tag{4.73}$$

and similarly under permutations. This is consistent with our analysis here, since

$$(\mathcal{L}_1, \mathcal{L}_2, \mathcal{L}_3) = (\mathcal{S}_{23}^{++}, -\mathcal{S}_{13}^{++}, \mathcal{S}_{12}^{++})$$
(4.74)

and \mathcal{S}_{ij}^{++} augments Δ_i and Δ_j by one and d by two. The \mathcal{R}_i operators are then their shadow conjugates as defined in (4.62), producing the expected shifts (4.63). Finally,

$$\mathcal{S} = -\mathcal{L}_1 \mathcal{L}_2 - \mathcal{L}_2 \mathcal{L}_3 - \mathcal{L}_3 \mathcal{L}_1 \tag{4.75}$$

does not appear explicitly in [85], but can be derived as follows. Writing the 3-point function as the triple-K integral $I_{d/2-1,\{\beta_1,\beta_2,\beta_3\}}$, from (4.71) we have

$$-\mathcal{S}I_{d/2-1,\{\beta_1,\beta_2,\beta_3\}} = I_{d/2+1,\{\beta_1-1,\beta_2-1,\beta_3\}} + I_{d/2+1,\{\beta_1,\beta_2-1,\beta_3-1\}} + I_{d/2+1,\{\beta_1-1,\beta_2,\beta_3-1\}}$$
$$= (\mathcal{R}_1 + \mathcal{R}_2 + \mathcal{R}_3)I_{d/2,\{\beta_1-1,\beta_2-1,\beta_3-1\}}$$
$$= \left(\frac{d}{2} + \beta_t + 4\right)I_{d/2,\{\beta_1-1,\beta_2-1,\beta_3-1\}}$$
(4.76)

where the final line follows by eliminating the sum of \mathcal{R}_i operators using the dilatation Ward identity. The effect of \mathcal{S} is thus to increase $d \to d+2$ and all $\beta_i \to \beta_i - 1$. All dimensions $\Delta_i = \beta_i + d/2$ are then preserved, consistent with (4.53).

4-point shift operators

The 3-point calculations above provide a first consistency check, but to obtain genuinely new shift operators we now turn to the 4-point function.

To write our results, we introduce the Mandelstam variables,

$$P_I = \{p_1^2, p_2^2, p_3^2, p_4^2, s^2, t^2\}, \qquad I = 1, \dots, 6$$
(4.77)

where $s^2 = (\boldsymbol{p}_1 + \boldsymbol{p}_2)^2$ and $t^2 = (\boldsymbol{p}_2 + \boldsymbol{p}_3)^2$, and define the derivative operators

$$D_I = \frac{\partial}{\partial P_I}, \qquad D_{IJ} = D_I D_J, \qquad D_{IJK} = D_I D_J D_K.$$
 (4.78)

Defining $S_{ij}^{++} = 4(-1)^{i+j}S_{ij}^{++}$ and S = -8S to suppress trivial numerical factors, from

(4.51) and (4.41), using the chain rule analogous to (4.68), we obtain the operators

$$S_{12}^{++} = D_{34} + D_{45} + D_{35} + D_{56},$$

$$S_{13}^{++} = D_{24} - D_{56},$$

$$S_{14}^{++} = D_{23} + D_{26} + D_{36} + D_{56},$$

$$S_{23}^{++} = D_{14} + D_{16} + D_{46} + D_{56},$$

$$S_{24}^{++} = D_{13} - D_{56},$$

$$S_{34}^{++} = D_{12} + D_{15} + D_{25} + D_{56},$$
(4.79)

and

$$S = D_{456} + D_{356} + D_{346} + D_{256} + D_{246} + D_{245} + D_{235} + D_{234} + D_{156} + D_{145} + D_{136} + D_{135} + D_{134} + D_{126} + D_{124} + D_{123}.$$
 (4.80)

As per (4.53), the S_{ij}^{++} increase Δ_i and Δ_j by one and d by two, while S increases d by two.

Following section 4.3.2, we can obtain further shift operators by shadow conjugation. As noted earlier, shadow conjugating each S_{ij}^{++} on either of the (i, j) indices has no effect: from (4.79), S_{ij}^{++} contains neither D_i or D_j hence these shadow factors commute through the operator. Instead, we must shadow conjugate each S_{ij}^{++} with respect to indices other than (i, j). At four points, once a pair of insertions (i, j) is specified, the remaining set also form a pair $(k, l) \neq (i, j)$. Shadow conjugating each S_{ij}^{++} on the opposite pair (k, l)then defines

$$\bar{S}_{ij}^{++} = p_k^{2(\beta_k+1)} p_l^{2(\beta_l+1)} S_{ij}^{++} p_k^{-2\beta_k} p_l^{-2\beta_l}, \qquad (k,l) \neq (i,j)$$
(4.81)

where $\beta_i = \Delta_i - d/2$. Expressed in terms of the variables (4.77), we find

$$\bar{S}_{12}^{++} = \beta_3\beta_4 - \beta_4P_3D_3 - \beta_3P_4D_4 - (\beta_3P_4 + \beta_4P_3)D_5 + P_3P_4S_{12}^{++},
\bar{S}_{13}^{++} = \beta_2\beta_4 - \beta_4P_2D_2 - \beta_2P_4D_4 + P_2P_4S_{13}^{++},
\bar{S}_{14}^{++} = \beta_2\beta_3 - \beta_3P_2D_2 - \beta_2P_3D_3 - (\beta_2P_3 + \beta_3P_2)D_6 + P_2P_3S_{14}^{++},
\bar{S}_{23}^{++} = \beta_1\beta_4 - \beta_4P_1D_1 - \beta_1P_4D_4 - (\beta_1P_4 + \beta_4P_1)D_6 + P_1P_4S_{23}^{++},
\bar{S}_{24}^{++} = \beta_1\beta_3 - \beta_3P_1D_1 - \beta_1P_3D_3 + P_1P_3S_{24}^{++},
\bar{S}_{34}^{++} = \beta_1\beta_2 - \beta_2P_1D_1 - \beta_1P_2D_2 - (\beta_1P_2 + \beta_2P_1)D_5 + P_1P_2S_{34}^{++}.$$
(4.82)

The action of each operator \bar{S}_{ij}^{++} is to shift $d \to d+2$, $\Delta_{i,j} \to \Delta_{i,j}+1$ and $\Delta_{k,l} \to \Delta_{k,l}+2$. This leaves β_i and β_j invariant while raising β_k and β_l by one. Heuristically, these \bar{S}_{ij}^{++} are then the 4-point generalisation of the 3-point \mathcal{R}_i operators in (4.70). Likewise, the S_{ij}^{++} in (4.79) leave β_i and β_j invariant but lower β_k and β_l by one, and represent the 4-point generalisation of the 3-point \mathcal{L}_i operators.

Besides shadow conjugating S_{ij}^{++} with respect to the pair (k, l), one can of course also conjugate with respect to only a single index k to find operators sending $d \to d+2$, $\Delta_{i,j} \to \Delta_{i,j} + 1$ and $\Delta_k \to \Delta_k + 2$ only. One can also apply the shadow conjugation procedure to the $d \to d+2$ operator S. All these operators can be evaluated similarly to the \bar{S}_{ij}^{++} above and we will not write them explicitly. One case of particular interest, however, corresponds to acting with \bar{S}_{ij}^{++} followed by S_{ij}^{++} , which produces an overall shift of $d \to d + 4$ while increasing all operator dimensions by two. The same shift is produced when acting with these operators in the opposite order (remembering to shift $\beta_{k,l} \to \beta_{k,l} - 1$ in \bar{S}_{ij}^{++} to account for the prior action of S_{ij}^{++}). By subtracting, we then obtain a shift operator of only *second* order in derivatives, rather than fourth. For example,

$$\bar{S}_{24}^{++}\Big|_{\beta_1-1,\beta_3-1}S_{24}^{++} - S_{24}^{++}\bar{S}_{24}^{++}\Big|_{\beta_1,\beta_3} = (\beta_1+\beta_3)D_{56}$$
(4.83)

and so D_{56} shifts $d \to d + 4$ while sending all $\Delta_i \to \Delta_i + 2$ and preserving the β_i .

Finally, let us emphasise that the action of all these shift operators is general and not in any way tied to the simplex representation: *any* solution of the 4-point conformal Ward identities is mapped to an appropriately shifted solution.⁵ We have confirmed this explicitly by computing all the relevant intertwining relations between the shift operators in this section and the conformal Ward identities, whose form in Mandelstam variables can be found in *e.g.*, [36, 29]. Thus, for example,

$$\mathcal{K}(\Delta_1 + 1, \Delta_2 + 1, \Delta_3, \Delta_4, d+2)S_{12}^{++} = S_{12}^{++}\mathcal{K}(\Delta_1, \Delta_2, \Delta_3, \Delta_4, d)$$
(4.84)

where $\mathcal{K}(\{\Delta_i\}, d)$ represents schematically any of the special conformal or dilatation Ward identities with the operator and spacetime dimensions as indicated. Applying this relation to any CFT correlator with dimensions $(\{\Delta_i\}, d)$, the right-hand side vanishes and the lefthand side then indicates that the action of S_{12}^{++} produces a solution of the shifted Ward identities. Intertwining relations such as these⁶ allow the shift action of operators to be established independently of any integral representation for the correlator.

4.3.4 Operators preserving d

A different class of weight-shifting operators that preserve the spacetime dimension d while shifting the Δ_i was identified in [83]. In momentum space, these operators have been applied to de Sitter correlators in [29, 30]. With the aid of shadow conjugation, we can write them in the compact form [36]

$$\mathcal{W}_{ij}^{--} = \frac{1}{2} \left(\frac{\partial}{\partial p_i^{\mu}} - \frac{\partial}{\partial p_j^{\mu}} \right) \left(\frac{\partial}{\partial p_{i\mu}} - \frac{\partial}{\partial p_{j\mu}} \right)
\mathcal{W}_{ij}^{+-} = p_i^{2(\beta_i+1)} \mathcal{W}_{ij}^{--} p_i^{-2\beta_i}
\mathcal{W}_{ij}^{-+} = p_j^{2(\beta_j+1)} \mathcal{W}_{ij}^{--} p_j^{-2\beta_j}
\mathcal{W}_{ij}^{++} = p_i^{2(\beta_i+1)} p_j^{2(\beta_j+1)} \mathcal{W}_{ij}^{--} p_i^{-2\beta_i} p_j^{-2\beta_j},$$
(4.85)

 $^{{}^{5}}$ Up to a technical caveat (common to all shift operators) that where divergences occur, one must work in a suitable dimensional regularisation scheme. In some cases the shift operator then only yields the leading divergences of the shifted correlator, see the discussion in [36].

⁶More generally, the right-hand side of (4.84) could feature *any* operator in the left ideal of the conformal Ward identities, since all that matters is that it vanishes when acting on a solution with dimensions $(\{\Delta_i\}, d)$.

where $\beta_i = \Delta_i - d/2$ and $1 \le i < j \le n - 1$ so p_n is taken as the dependent momentum. Their action is to shift

$$\mathcal{W}_{ij}^{\sigma_i\sigma_j}: \quad \Delta_i \to \Delta_i + \sigma_i, \quad \Delta_j \to \Delta_j + \sigma_j, \quad d \to d, \quad \{\sigma_i, \sigma_j\} \in \pm 1.$$
(4.86)

In this section, our goal is to understand the action of the simplest of these operators, W_{ij}^{--} , from the simplex perspective. The action of the others then follows via shadow conjugation, or else can be shown explicitly: for example, we analyse W_{ij}^{-+} in section 4.5.2.

We begin by writing the Schwinger exponential (4.10) in the form

$$-\operatorname{tr}(g^{-1} \cdot G) = \sum_{k < l}^{n} s_{kl} \, \boldsymbol{p}_k \cdot \boldsymbol{p}_l = -\sum_{k < l}^{n-1} (s_{kn} + s_{ln} - s_{kl}) (\boldsymbol{p}_k \cdot \boldsymbol{p}_l) - \sum_{k}^{n-1} s_{kn} \, p_k^2.$$
(4.87)

As only the independent momenta feature in this last expression, the action of \mathcal{W}_{ij}^{--} on the Schwinger exponential can be rewritten as a differential operator in the s_{kl} . We will do this in several steps. First, notice that

$$\frac{\partial}{\partial p_i^{\mu}} e^{-\operatorname{tr}(g^{-1} \cdot G)} = -\left(2s_{in}p_i^{\mu} + \sum_{k \neq i}^{n-1}(s_{in} + s_{kn} - s_{ik})p_k^{\mu}\right)e^{-\operatorname{tr}(g^{-1} \cdot G)}$$
$$= -\sum_k^{n-1}(s_{in} + s_{kn} - s_{ik})p_k^{\mu}e^{-\operatorname{tr}(g^{-1} \cdot G)}, \qquad (4.88)$$

where in the second line s_{ik} vanishes for i = k. This gives

$$\left(\frac{\partial}{\partial p_i^{\mu}} - \frac{\partial}{\partial p_j^{\mu}}\right) e^{-\operatorname{tr}(g^{-1} \cdot G)} = \sum_{k}^{n-1} (s_{ik} - s_{jk} - s_{in} + s_{jn}) p_k^{\mu} e^{-\operatorname{tr}(g^{-1} \cdot G)},$$
(4.89)

and hence

$$\mathcal{W}_{ij}^{--}e^{-\operatorname{tr}(g^{-1}\cdot G)}$$

$$= \left(-ds_{ij} + \frac{1}{2}\sum_{k,l}^{n-1}(s_{ik} - s_{jk} - s_{in} + s_{jn})(s_{il} - s_{jl} - s_{in} + s_{jn})\boldsymbol{p}_k \cdot \boldsymbol{p}_l\right)e^{-\operatorname{tr}(g^{-1}\cdot G)}.$$

$$(4.90)$$

To rewrite these momentum dot products as derivatives with respect to the s_{kl} , we now rearrange this sum as follows. Using momentum conservation $p_k^2 = -\sum_{l \neq k}^n \boldsymbol{p}_k \cdot \boldsymbol{p}_l$, for any generic coefficient A_k such that $A_n = 0$, we have

$$\sum_{k,l}^{n-1} A_k A_l \boldsymbol{p}_k \cdot \boldsymbol{p}_l = \sum_{\substack{k,l \\ k \neq l}}^{n-1} A_k A_l \boldsymbol{p}_k \cdot \boldsymbol{p}_l + \sum_{k}^{n-1} A_k^2 p_k^2$$
$$= \sum_{\substack{k,l \\ k \neq l}}^{n-1} A_k (A_l - A_k) \boldsymbol{p}_k \cdot \boldsymbol{p}_l - \sum_{k}^{n-1} A_k^2 \boldsymbol{p}_k \cdot \boldsymbol{p}_n$$

$$= -\frac{1}{2} \sum_{\substack{k,l \\ k \neq l}}^{n-1} (A_l - A_k)^2 \, \boldsymbol{p}_k \cdot \boldsymbol{p}_l - \sum_k^{n-1} A_k^2 \, \boldsymbol{p}_k \cdot \boldsymbol{p}_n$$

$$= -\sum_{k < l}^n (A_l - A_k)^2 \, \boldsymbol{p}_k \cdot \boldsymbol{p}_l$$
(4.91)

where in the final line the sum runs up to n. Setting $A_k = s_{ik} - s_{jk} - s_{in} + s_{jn}$, we find

$$\mathcal{W}_{ij}^{--}e^{-\operatorname{tr}(g^{-1}\cdot G)} = \left(-ds_{ij} - \frac{1}{2}\sum_{k
$$= \left(-ds_{ij} - \frac{1}{2}\sum_{k
$$= \left(-ds_{ij} + 2\partial_{v_{ij}}\right)e^{-\operatorname{tr}(g^{-1}\cdot G)}$$
$$= 2|g|^{d/2} \partial_{v_{ij}}\left(|g|^{-d/2} e^{-\operatorname{tr}(g^{-1}\cdot G)}\right).$$
(4.92)$$$$

In the second line here, we exchanged $\boldsymbol{p}_k \cdot \boldsymbol{p}_l$ for $\partial_{s_{kl}}$ using the first expression in (4.87). The change of variables from $\partial_{s_{kl}}$ to $\partial_{v_{ij}}$ in the third line then comes from the Jacobian evaluated in appendix B.2.2, and in the final line we used (4.26).

The action of \mathcal{W}_{ij}^{--} on the full simplex integral (4.17) now follows. First, the outer factor of $|g|^{d/2}$ in (4.92) cancels with the factor $\mathcal{U}^{-d/2} = |g|^{-d/2}$ in (4.17). Integrating by parts with respect to v_{ij} , assuming the boundary terms vanish,⁷ the derivative then acts on the prefactors as

$$-2\partial_{v_{ij}}\left(\prod_{k(4.93)$$

Here, the terms coming from $\partial_{v_{ij}}$ hitting the cross-ratios (4.16) inside the arbitrary function $f(\hat{\boldsymbol{v}})$, as well as those from hitting $v_{ij}^{-\alpha_{ij}-1}$, have been repackaged in the form $v_{ij}^{-1}\tilde{f}(\hat{\boldsymbol{v}})$ for some new function of cross-ratios $\tilde{f}(\hat{\boldsymbol{v}})$. Thus, overall, we find

$$\mathcal{W}_{ij}^{--} \Big(\prod_{k(4.94)$$

The action of \mathcal{W}_{ij}^{--} on the simplex is therefore to send $\alpha_{ij} \to \alpha_{ij} + 1$, up to changes of the arbitrary function. The latter is of no account as far as mapping one solution of the conformal Ward identities to another is concerned.⁸ From (4.2), we now confirm

⁷For the upper limit this is automatic for momentum configurations with non-vanishing Gram determinant thanks to the decaying exponential. The lower limit vanishes provided $\alpha_{ij} < 0$.

⁸An exception is if \mathcal{W}_{ij}^{-1} maps us from a finite correlator to a singular one, corresponding to a solution of the conditions $d + \sum_{i=1}^{n} \sigma_i (\Delta_i - d/2) = -2k$ for some non-negative integer k and a choice of signs

that sending $\alpha_{ij} \to \alpha_{ij} + 1$ while keeping the remaining α_{kl} fixed is equivalent to sending $\Delta_i \to \Delta_i - 1$ and $\Delta_j \to \Delta_j - 1$ while preserving d, in perfect agreement with (4.86).

4.4 Verifying the conformal Ward identities

In this section, we prove that the parametric representation of the simplex integral (4.17) satisfies the conformal Ward identities for any arbitrary function of cross-ratios. The corresponding result for the vectorial simplex integral (4.1) was established in [47, 46]. Working purely in momentum space, our approach is to show that the action of the Ward identities on the simplex integral reduces to a total derivative. With a degree of hindsight, the structure of this total derivative, obtained in (4.118), can also be understood from somewhat simpler position-space arguments. We will return to these in section 4.5.1.

As the dilatation Ward identity can be verified by power counting, we focus on the special conformal Ward identities

$$0 = \sum_{j=1}^{n-1} \left(p_j^{\mu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\nu}} - 2p_j^{\nu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\mu}} + 2(\Delta_j - d) \frac{\partial}{\partial p_j^{\mu}} \right) \langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle\!\rangle, \qquad (4.95)$$

treating p_n as the dependent momentum. As a first step, we rewrite the action of each individual term in (4.95) on the Schwinger exponential as an equivalent differential operator in v_{ij} . From (4.10), we have

$$\sum_{j}^{n-1} 2(\Delta_j - d) \frac{\partial}{\partial p_j^{\mu}} e^{-\operatorname{tr}(g^{-1} \cdot G)} = \sum_{j}^{n-1} p_j^{\mu} \left(-4 \sum_{k}^{n-1} (\Delta_k - d) g_{jk}^{-1} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)},$$
(4.96)

$$\sum_{j}^{n-1} p_{j}^{\mu} \frac{\partial}{\partial p_{j}^{\nu}} \frac{\partial}{\partial p_{j}^{\nu}} e^{-\operatorname{tr}(g^{-1} \cdot G)} = \sum_{j}^{n-1} p_{j}^{\mu} \Big(-2dg_{jj}^{-1} + 4\sum_{k,l}^{n-1} g_{jk}^{-1} g_{jl}^{-1} \boldsymbol{p}_{k} \cdot \boldsymbol{p}_{l} \Big) e^{-\operatorname{tr}(g^{-1} \cdot G)}.$$
(4.97)

Using (4.28) for the inverse metric and the manipulation (4.91), this last expression can be rewritten analogously to (4.92):

$$\sum_{j}^{n-1} p_{j}^{\mu} \frac{\partial}{\partial p_{j}^{\nu}} \frac{\partial}{\partial p_{j}^{\nu}} e^{-\operatorname{tr}(g^{-1} \cdot G)} = \sum_{j}^{n-1} p_{j}^{\mu} \Big(-2ds_{jn} - \sum_{k
$$= \sum_{j}^{n-1} p_{j}^{\mu} \Big(-2ds_{jn} - \sum_{k
$$= \sum_{j}^{n-1} p_{j}^{\mu} \Big(-2ds_{jn} + 4\partial_{v_{jn}} \Big) e^{-\operatorname{tr}(g^{-1} \cdot G)}. \tag{4.98}$$$$$$

 $\{\sigma_i\} \in \pm 1$, see [47]. In such cases, the arbitrary function $\tilde{f}(\hat{v})$ vanishes. In dimensional regularisation, this zero then cancels the pole coming from the divergent correlator such that the result is finite, see [36].

Next, we must deal with

$$\sum_{j}^{n-1} \left(-2p_{j}^{\nu} \frac{\partial}{\partial p_{j}^{\nu}} \frac{\partial}{\partial p_{j}^{\mu}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$

$$= 4 \sum_{j}^{n-1} p_{j}^{\mu} \left(g_{jj}^{-1} - \sum_{k,l}^{n-1} (g_{jk}^{-1} + g_{jl}^{-1}) g_{kl}^{-1} \boldsymbol{p}_{k} \cdot \boldsymbol{p}_{l} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$

$$= 4 \sum_{j}^{n-1} p_{j}^{\mu} \left(g_{jj}^{-1} - 2 \sum_{k

$$(4.99)$$$$

Using (4.28) and momentum conservation, the p_k^2 terms in this final sum can be rewritten

$$-2\sum_{k}^{n-1} g_{jk}^{-1} g_{kk}^{-1} p_{k}^{2} = \sum_{k}^{n} s_{kn} (s_{jn} + s_{kn} - s_{jk}) \sum_{l \neq k}^{n} \boldsymbol{p}_{k} \cdot \boldsymbol{p}_{l}$$
$$= \sum_{k < l}^{n} \left(s_{kn} (s_{jn} + s_{kn} - s_{jk}) + s_{ln} (s_{jn} + s_{ln} - s_{jl}) \right) \boldsymbol{p}_{k} \cdot \boldsymbol{p}_{l}.$$
(4.100)

In the first line here, notice we extended the sum over k to run up to n, which is possible since the additional term with k = n vanishes as $s_{nn} = 0$. To get the second line, we then re-expressed the terms for which k > l by swapping $k \leftrightarrow l$. For convenience, it is useful to define

$$\hat{g}_{ij}^{-1} = \frac{1}{2}(s_{in} + s_{jn} - s_{ij}) = \begin{cases} g_{ij}^{-1} & i, j \le n - 1, \\ 0 & i = n \text{ and/or } j = n, \end{cases}$$
(4.101)

effectively extending the $(n-1) \times (n-1)$ matrix g_{ij}^{-1} to an $n \times n$ matrix \hat{g}_{ij}^{-1} by adding a final row and column of zeros. This allows us to compactly rewrite (4.99) and (4.100) as

$$\sum_{j}^{n-1} \left(-2p_{j}^{\nu} \frac{\partial}{\partial p_{j}^{\nu}} \frac{\partial}{\partial p_{j}^{\mu}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$

$$= \sum_{j}^{n-1} p_{j}^{\mu} \left(4\hat{g}_{jj}^{-1} + 8\sum_{k

$$(4.102)$$$$

Here, the sum over l for the $\mathbf{p}_k \cdot \mathbf{p}_l$ terms in (4.99) has similarly been extended to run up to n, noting the additional l = n term vanishes. We then replaced $\mathbf{p}_k \cdot \mathbf{p}_l$ by a derivative with respect to s_{kl} using (4.87). The result now simplifies further upon exchanging

$$\partial_{s_{kl}} = \sum_{a < b}^{n} \frac{\partial v_{ab}}{\partial s_{kl}} \, \partial_{v_{ab}} = -\sum_{a < b}^{n} \tilde{g}_{a(k} \tilde{g}_{l)b} \, \partial_{v_{ab}}, \tag{4.103}$$

where \tilde{g}_{ab} is the Laplacian matrix (4.23) and the Jacobian is evaluated in appendix B.2.2.

First, we write

$$8\sum_{k(4.104)$$

where the sum over k < l of (k, l)-symmetric terms has been rewritten as half the sum over all k and l, noting the terms with k = l explicitly cancel. The final two terms now vanish since all row and column sums of the Laplacian matrix \tilde{g} are zero:

$$\sum_{k,l}^{n} \tilde{g}_{ak} \, \tilde{g}_{lb} \, \hat{g}_{kk}^{-1} \, \hat{g}_{jk}^{-1} = \sum_{k}^{n} \tilde{g}_{ak} \, \hat{g}_{kk}^{-1} \, \hat{g}_{jk}^{-1} \sum_{l}^{n} \, \tilde{g}_{lb} = 0, \qquad (4.105)$$

$$\sum_{k,l}^{n} \tilde{g}_{ak} \, \tilde{g}_{lb} \, \hat{g}_{ll}^{-1} \, \hat{g}_{jl}^{-1} = \sum_{k}^{n} \tilde{g}_{lb} \, \hat{g}_{ll}^{-1} \, \hat{g}_{jl}^{-1} \sum_{k}^{n} \, \tilde{g}_{ak} = 0.$$
(4.106)

For the first two terms in (4.104), we use the identity

$$\sum_{k}^{n} \tilde{g}_{ik} \, \hat{g}_{kj}^{-1} = \delta_{ij} - \delta_{in}, \qquad i, j \le n.$$
(4.107)

To derive this, note the sum over k restricts to $k \leq n-1$ from (4.101), then for $i, j \leq n-1$ we have $\tilde{g}_{ik}\hat{g}_{kj}^{-1} = g_{ik}g_{kj}^{-1}$. For $i = n, j \leq n-1$ we use $\tilde{g}_{nk}\hat{g}_{kj}^{-1} = -\sum_{l}^{n-1}g_{lk}g_{kj}^{-1}$ and for j = n and any *i* the sum vanishes from (4.101). With the aid of this identity, we then find

$$\sum_{j}^{n-1} \left(-2p_{j}^{\nu} \frac{\partial}{\partial p_{j}^{\nu}} \frac{\partial}{\partial p_{j}^{\mu}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$
$$= 4 \sum_{j}^{n-1} p_{j}^{\mu} \left(s_{jn} - \partial_{v_{jn}} - \sum_{a < b}^{n} (\hat{g}_{ja}^{-1} + \hat{g}_{jb}^{-1}) \theta_{v_{ab}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$
(4.108)

where we used $\tilde{g}_{ab} = -v_{ab}$ for a < b to obtain the Euler operator $\theta_{v_{ab}} = v_{ab}\partial_{v_{ab}}$.

Assembling the pieces above, the action of the conformal Ward identity is now

$$\sum_{j=1}^{n-1} \left(p_j^{\mu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\nu}} - 2p_j^{\nu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\mu}} + 2(\Delta_j - d) \frac{\partial}{\partial p_j^{\mu}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$

$$= 4 \sum_{j}^{n-1} p_j^{\mu} \left(\left(1 - \frac{d}{2} \right) s_{jn} - \sum_{k}^{n-1} (\Delta_k - d) g_{jk}^{-1} - \sum_{a < b}^{n} (\hat{g}_{ja}^{-1} + \hat{g}_{jb}^{-1}) \theta_{v_{ab}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$

$$= 4 \sum_{j}^{n-1} p_j^{\mu} \left(\left(1 - \frac{d}{2} \right) s_{jn} + d \sum_{a}^{n} \hat{g}_{ja}^{-1} + \sum_{a \neq b}^{n} \hat{g}_{ja}^{-1} (\alpha_{ab} - \theta_{v_{ab}}) \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$

$$(4.109)$$

using (4.2) in the last line. Finally, we need two further identities:

$$\sum_{a\neq b}^{n} \theta_{v_{ab}} \hat{g}_{ja}^{-1} = -s_{jn}, \qquad \sum_{a\neq b}^{n} \hat{g}_{ja}^{-1} v_{ab} s_{ab} = -s_{jn} + 2\sum_{a}^{n} \hat{g}_{ja}^{-1}$$
(4.110)

To establish the first of these, we write

$$\sum_{a\neq b}^{n} \theta_{v_{ab}} \hat{g}_{ja}^{-1} = -\sum_{a\neq b}^{n-1} g_{ab} \frac{\partial g_{ja}^{-1}}{\partial v_{ab}} + \sum_{a}^{n-1} v_{an} \frac{\partial g_{ja}^{-1}}{\partial v_{an}} = -\sum_{a\neq b}^{n-1} g_{ab} \frac{\partial g_{ja}^{-1}}{\partial v_{ab}} + \sum_{a,b}^{n-1} g_{ab} \frac{\partial g_{ja}^{-1}}{\partial v_{an}}$$
(4.111)

then use the chain rule, which for $i, j, k, l \leq n - 1$ gives

$$\frac{\partial g_{ij}^{-1}}{\partial v_{kl}} = -\sum_{a,b}^{n-1} g_{i(a}^{-1} g_{b)j}^{-1} \frac{\partial g_{ab}}{\partial v_{kl}} = (g_{ik}^{-1} - g_{il}^{-1})(g_{jl}^{-1} - g_{jk}^{-1}), \qquad (4.112)$$

$$\frac{\partial g_{ij}^{-1}}{\partial v_{kn}} = -\sum_{a,b}^{n-1} g_{i(a}^{-1} g_{b)j}^{-1} \frac{\partial g_{ab}}{\partial v_{kn}} = -g_{ik}^{-1} g_{kj}^{-1}.$$
(4.113)

Inserting these into (4.111), the sum over $a \neq b$ can be extended to run over all a, b since the term with a = b vanishes. The only non-cancelling term is then $-g_{jj}^{-1} = -s_{jn}$ as required.

For the second identity in (4.110), we use (4.28) to rewrite

$$\sum_{a\neq b}^{n} \hat{g}_{ja}^{-1} v_{ab} s_{ab} = -\sum_{a\neq b}^{n-1} g_{ja}^{-1} g_{ab} s_{ab} + \sum_{a}^{n-1} g_{ja}^{-1} v_{an} s_{an}$$
$$= -\sum_{a\neq b}^{n-1} g_{ja}^{-1} g_{ab} (g_{aa}^{-1} + g_{bb}^{-1} - 2g_{ab}^{-1}) + \sum_{a}^{n-1} g_{ja}^{-1} \left(\sum_{b}^{n-1} g_{ab}\right) g_{aa}^{-1}.$$
(4.114)

The sum over $a \neq b$ can then be extended to run over all a, b as the term with a = b cancels, after which the first and the last terms cancel and the result follows.

With the aid of the identities (4.110), we find that (4.109) becomes

$$\sum_{j=1}^{n-1} \left(p_j^{\mu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\nu}} - 2p_j^{\nu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\mu}} + 2(\Delta_j - d) \frac{\partial}{\partial p_j^{\mu}} \right) e^{-\operatorname{tr}(g^{-1} \cdot G)}$$
$$= -4 \sum_j^{n-1} p_j^{\mu} |g|^{d/2} \Omega^{-1} \sum_{a \neq b}^n \partial_{v_{ab}} \left(v_{ab} \, \hat{g}_{ja}^{-1} |g|^{-d/2} \Omega \, e^{-\operatorname{tr}(g^{-1} \cdot G)} \right)$$
(4.115)

where $\Omega = \prod_{k < l}^{n} v_{kl}^{-\alpha_{kl}-1}$. Recalling that the simplex representation (4.17) is

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \Big(\prod_{k< l}^n \int_0^\infty \mathrm{d}v_{kl} \, v_{kl}^{-\alpha_{kl}-1} \Big) f(\hat{\boldsymbol{v}}) |g|^{-d/2} \, e^{-\mathrm{tr}\,(g^{-1}\cdot G)},\tag{4.116}$$

we note that

$$\sum_{\substack{b\\b\neq a}}^{n} \theta_{v_{ab}} f(\hat{\boldsymbol{v}}) = 0 \tag{4.117}$$

since whenever the index *a* appears in a cross ratio $\hat{v}_{[acde]} = v_{ac}v_{de}/v_{ad}v_{ce}$ it enters with equal weight in the numerator and the denominator producing a cancellation. Acting with the Ward identity thus yields a total derivative:

$$\sum_{j=1}^{n-1} \left(p_j^{\mu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\nu}} - 2p_j^{\nu} \frac{\partial}{\partial p_j^{\nu}} \frac{\partial}{\partial p_j^{\mu}} + 2(\Delta_j - d) \frac{\partial}{\partial p_j^{\mu}} \right) \langle \langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle \rangle$$
$$= -4 \sum_j^{n-1} p_j^{\mu} \left(\prod_{k < l}^n \int_0^\infty \mathrm{d} v_{kl} \right) \sum_{a \neq b}^n \partial_{v_{ab}} \left(v_{ab} \, \hat{g}_{ja}^{-1} \, f(\hat{\boldsymbol{v}}) |g|^{-d/2} \Omega \, e^{-\mathrm{tr} \, (g^{-1} \cdot G)} \right). \tag{4.118}$$

The boundary terms vanish under reasonable assumptions: for generic momentum configurations with non-vanishing Gram determinant, the upper limit is suppressed by the decay of the Schwinger exponential; the lower limit is zero provided $v_{ab}^{-\alpha_{ab}} f(\hat{v})$ vanishes as $v_{ab} \to 0$, which is satisfied whenever the simplex representation itself converges. The simplex integral thus solves the special conformal Ward identity.

4.5 Insight from position space

Thus far, our analysis has been entirely in momentum space. However, as noted above, the form of the total derivative produced by the action of the special conformal Ward identity in (4.118) can also be understood through independent position-space arguments. We present these in section 4.5.1. Then, in section 4.5.2, we show how similar position-space arguments can be applied to verify the action of *d*-preserving shift operators such as \mathcal{W}_{12}^{-+} .

4.5.1 The conformal Ward identities

To Fourier transform the simplex representation (4.17) to position space, we compute

$$\langle \mathcal{O}(\boldsymbol{x}_{1}) \dots \mathcal{O}(\boldsymbol{x}_{n}) \rangle = \prod_{k}^{n-1} \int \frac{\mathrm{d}^{d} \boldsymbol{p}_{k}}{(2\pi)^{d}} e^{i\boldsymbol{p}_{k} \cdot \boldsymbol{x}_{kn}} \langle \langle \mathcal{O}(\boldsymbol{p}_{1}) \dots \mathcal{O}(\boldsymbol{p}_{n}) \rangle$$

$$= \left(\prod_{i < j}^{n} \int \mathrm{d} v_{ij} v_{ij}^{-\alpha_{ij}-1}\right) f(\hat{\boldsymbol{v}}) |g|^{-d/2} \left(\prod_{k}^{n-1} \int \frac{\mathrm{d}^{d} \boldsymbol{p}_{k}}{(2\pi)^{d}}\right) \exp\left(\sum_{k}^{n-1} i\boldsymbol{p}_{k} \cdot \boldsymbol{x}_{kn} - \sum_{k,l}^{n-1} g_{kl}^{-1} \boldsymbol{p}_{k} \cdot \boldsymbol{p}_{l}\right)$$

$$= \left(\prod_{i < j}^{n} \int \mathrm{d} v_{ij} v_{ij}^{-\alpha_{ij}-1}\right) \tilde{f}(\hat{\boldsymbol{v}}) \exp\left(-\frac{1}{4} \sum_{k,l}^{n-1} g_{kl} \boldsymbol{x}_{kn} \cdot \boldsymbol{x}_{ln}\right)$$

$$(4.119)$$

where $x_{ij} = x_i - x_j$, and for the Gaussian integral over momenta we completed the square:

$$\sum_{k}^{n-1} i oldsymbol{p}_k \cdot oldsymbol{x}_{kn} - \sum_{k,l}^{n-1} g_{kl}^{-1} \, oldsymbol{p}_k \cdot oldsymbol{p}_l$$

$$= -\sum_{k,l}^{n-1} g_{kl}^{-1} (\boldsymbol{p}_k - \frac{i}{2} \sum_{a}^{n-1} g_{ka} \boldsymbol{x}_{an}) \cdot (\boldsymbol{p}_l - \frac{i}{2} \sum_{b}^{n-1} g_{lb} \boldsymbol{x}_{bn}) - \frac{1}{4} \sum_{k,l}^{n-1} g_{kl} \boldsymbol{x}_{kn} \cdot \boldsymbol{x}_{ln}. \quad (4.120)$$

The numerical factor from the integration can then be re-absorbed into the arbitrary function by setting $(4\pi)^{(1-n)d/2} f(\hat{\boldsymbol{v}}) = \tilde{f}(\hat{\boldsymbol{v}})$. The exponent in (4.119) now simplifies to⁹

$$-\frac{1}{4}\sum_{k,l}^{n-1} g_{kl} \boldsymbol{x}_{kn} \cdot \boldsymbol{x}_{ln} = -\frac{1}{4}\sum_{k,l}^{n} \tilde{g}_{kl} \boldsymbol{x}_{k} \cdot \boldsymbol{x}_{l} = -\frac{1}{4}\sum_{i< j}^{n} v_{ij} x_{ij}^{2}, \qquad (4.121)$$

and hence the simplex representation in position space is

$$\langle \mathcal{O}(\boldsymbol{x}_1) \dots \mathcal{O}(\boldsymbol{x}_n) \rangle = \left(\prod_{i < j}^n \int \mathrm{d}v_{ij} \, v_{ij}^{-\alpha_{ij}-1} e^{-\frac{1}{4}v_{ij}x_{ij}^2} \right) \tilde{f}(\hat{\boldsymbol{v}}). \tag{4.122}$$

If the arbitrary function $\tilde{f}(\hat{v})$ is a product of powers, this expression reduces to the conformal correlator $\prod_{i < j}^{n} x_{ij}^{2\tilde{\alpha}_{ij}}$ where the $\tilde{\alpha}_{ij}$ satisfy $\sum_{j \neq i} \tilde{\alpha}_{ij} = -\Delta_i$. More generally, wherever $\tilde{f}(\hat{v})$ admits a Mellin-Barnes representation, we recover $\prod_{i < j}^{n} x_{ij}^{2\alpha_{ij}}$ times a function of position-space cross ratios as shown in [46]. However, the most straightforward way to check that (4.122) solves the conformal Ward identities is to note that, when acting on a function $F = F(\{x_{kl}^2\})$ of the squared coordinate separations,

$$\sum_{i}^{n} \left(2x_{i}^{\mu}x_{i}^{\nu}\frac{\partial}{\partial x_{i}^{\nu}} - x_{i}^{2}\frac{\partial}{\partial x_{i}^{\mu}} + 2\Delta_{i}x_{i}^{\mu} \right)F = \sum_{i}^{n} 2x_{i}^{\mu} \left(\Delta_{i} + \sum_{\substack{j\\j\neq i}}^{n} x_{ij}^{2}\frac{\partial}{\partial(x_{ij}^{2})} \right)F.$$
(4.123)

It then follows that

$$\sum_{i}^{n} \left(2x_{i}^{\mu}x_{i}^{\nu}\frac{\partial}{\partial x_{i}^{\nu}} - x_{i}^{2}\frac{\partial}{\partial x_{i}^{\mu}} + 2\Delta_{i}x_{i}^{\mu} \right) \left(\prod_{k

$$= \sum_{i}^{n} 2x_{i}^{\mu} \left(\prod_{k

$$= \sum_{i}^{n} 2x_{i}^{\mu} \left(\prod_{k$$$$$$

where in the last line we integrated by parts¹⁰ then used (4.2). The middle line here accounts for the form of the total derivative we found earlier in (4.118). Multiplying by -i and Fourier transforming, the first line yields the momentum-space conformal Ward identity acting on the momentum-space simplex representation (*i.e.*, the left-hand side of

⁹Recall the analogous relation in a resistor network of simplex topology, namely, that the power dissipated is $\sum_{i< j}^{n} v_{ij}(V_i - V_j)^2 = \sum_{i,j}^{n} \tilde{g}_{ij}V_iV_j$, where v_{ij} is the conductivity and V_i the voltage at node *i*.

¹⁰As previously, the boundary terms vanish provided $v_{kl}^{-\alpha_{kl}}\tilde{f}(\hat{v})$ as $v_{kl} \to 0$.

(4.118)), while the middle line yields

$$\sum_{i}^{n-1} 2 \frac{\partial}{\partial p_{i}^{\mu}} \left(\left(\prod_{k

$$= \sum_{i}^{n-1} \left(\prod_{k

$$= -4 \sum_{a}^{n-1} p_{a}^{\mu} \left(\prod_{k

$$(4.125)$$$$$$$$

where in the second line we evaluated the momentum derivative of the exponential and pushed the factors of $\Omega = \prod_{k < l} v_{kl}^{-\alpha_{kl}-1}$, $f(\hat{v})$ and v_{ij} inside the v_{ij} -derivative which cancels the Δ_i term via (4.2). In the final line, we extended the sum over *i* to run up to *n* by replacing g_{ia}^{-1} with \hat{g}_{ia}^{-1} and combined it with the sum over *j*. Up to a relabelling of indices, this final line is now the total derivative appearing on the right-hand side of (4.118).

The manipulations above illustrate a general theme: given the simplicity of the positionspace simplex representation (4.122), it is often profitable to work with the position-space equivalents of differential operators in order to evaluate their action in terms of the v_{ij} variables. Both sides can then be Fourier transformed back to momentum space in order to deduce the action of the corresponding momentum-space operator on the momentumspace simplex in terms of the v_{ij} variables. In many cases this is more straightforward than working in momentum space throughout.

4.5.2 Action of \mathcal{W}_{12}^{-+}

As a further illustration of this approach, let us evaluate the action of the shift operator \mathcal{W}_{12}^{-+} defined in (4.85). After expanding out the derivative, this operator can easily be Fourier transformed to position space where it reads

$$\mathcal{W}_{12}^{-+} = \frac{1}{2} x_{12}^2 \frac{\partial}{\partial x_2^{\mu}} \frac{\partial}{\partial x_{2\mu}} + 2(\beta_2 + 1) \Big(\beta_2 + \frac{d}{2} - x_{12}^{\mu} \frac{\partial}{\partial x_2^{\mu}}\Big).$$
(4.126)

Acting on a function $F = F(\{x_{kl}^2\})$ of the squared coordinate separations, we find via the chain rule

$$\mathcal{W}_{12}^{-+}F = \sum_{\substack{i,j\\i,j\neq 2}}^{n} x_{12}^2 (x_{2i}^2 + x_{2j}^2 - x_{ij}^2) \frac{\partial^2 F}{\partial(x_{2i}^2)\partial(x_{2j}^2)} + \sum_{i\neq 2}^{n} \left(2(\beta_2 + 1)(x_{12}^2 - x_{1i}^2 + x_{2i}^2) + dx_{12}^2 \right) \frac{\partial F}{\partial(x_{2i}^2)} + 2(\beta_2 + 1) \left(\beta_2 + \frac{d}{2}\right) F$$

$$(4.127)$$

Acting on the Schwinger exponent appearing in the position-space simplex representation (4.122), this can be translated into v_{ij} -derivatives as

$$\mathcal{W}_{12}^{-+} \Big(\prod_{k

$$= 2 \Big(\prod_{k

$$+ \sum_{i=3}^{n} v_{2i} \Big((\beta_{2}+1 + \theta_{v_{12}}) (\partial_{v_{12}} + \partial_{v_{2i}} - \partial_{v_{1i}}) + \Big(\frac{d}{2} + \theta_{v_{2i}} \Big) \partial_{v_{12}} \Big)$$

$$+ \sum_{3\leq i< j}^{n} v_{2i} v_{2j} (\partial_{v_{2i}} + \partial_{v_{2j}} - \partial_{v_{ij}}) \partial_{v_{12}} \Big] e^{-\frac{1}{4} \sum_{k

$$(4.128)$$$$$$$$

where $\partial_{v_{ij}} = \partial/\partial v_{ij}$ and $\theta_{v_{ij}} = v_{ij}\partial_{v_{ij}}$. Integrating by parts, we find

$$\mathcal{W}_{12}^{-+} \Big(\prod_{k

$$(4.129)$$$$

We now rewrite the first part of the last line as

$$\begin{bmatrix} \sum_{3 \le i < j}^{n} v_{2i} v_{2j} (\partial_{v_{2i}} + \partial_{v_{2j}}) \partial_{v_{12}} \end{bmatrix} \Omega \tilde{f}(\hat{\boldsymbol{v}}) = \begin{bmatrix} \sum_{i=3}^{n} \left(\sum_{\substack{j=3\\j \ne i}}^{n} v_{2j} \right) \theta_{v_{2i}} \partial_{v_{12}} \left(\partial_{v_{12}} + \sum_{i \ne 2}^{n} \theta_{v_{2i}} \right) - \sum_{j=3}^{n} v_{2i} \theta_{v_{2i}} \partial_{v_{12}} \end{bmatrix} \Omega \tilde{f}(\hat{\boldsymbol{v}})$$
$$= \begin{bmatrix} -\sum_{i=3}^{n} v_{2i} \left((\theta_{v_{12}} + 1) - \beta_2 - \frac{d}{2} + (n-1) + \theta_{v_{2i}} \right) \partial_{v_{12}} \end{bmatrix} \Omega \tilde{f}(\hat{\boldsymbol{v}})$$
(4.130)

where in the final step we rewrote $\partial_{v_{12}}\theta_{v_{12}} = (\theta_{v_{12}} + 1)\partial_{v_{12}}$ and used $\sum_{i\neq 2}^{n} \theta_{v_{2i}}\tilde{f}(\hat{v}) = 0$, as follows from (4.117), along with (4.2) with $\Delta_2 = \beta_2 + d/2$ to replace

$$\left(\sum_{i\neq 2}^{n} \theta_{v_{2i}}\right) \Omega \tilde{f}(\hat{\boldsymbol{v}}) = (\beta_2 + d/2 - (n-1)) \Omega \tilde{f}(\hat{\boldsymbol{v}}).$$

$$(4.131)$$

Substituting (4.130) into (4.129) and making further use of (4.131), we find the result

$$\mathcal{W}_{12}^{-+} \Big(\prod_{k

$$(4.132)$$$$

Equivalently, acting on the position-space simplex with W_{12}^{-+} corresponds to acting on the arbitrary function $\tilde{f}(\hat{v})$ with the operator

$$\tilde{\mathcal{W}}_{12}^{-+} = -2\Omega^{-1} \Big[(\theta_{v_{12}} - \beta_2) \sum_{i=3}^n v_{2i} \partial_{v_{1i}} + \sum_{3 \le i < j}^n v_{2i} v_{2j} \partial_{v_{ij}} \partial_{v_{12}} \Big] \Omega.$$
(4.133)

The same remains true when we Fourier transform back to momentum space, giving

$$\mathcal{W}_{12}^{-+} \Big(\prod_{k(4.134)$$

Finally, it remains to check that the action of $\tilde{\mathcal{W}}_{12}^{-+}$ on the arbitrary function produces the required shift in dimensions $\Delta_1 \to \Delta_1 - 1$ and $\Delta_2 \to \Delta_2 + 1$. Since

$$\partial_{v_{ij}}\Omega = -(\alpha_{ij}+1)\frac{\Omega}{v_{ij}}, \qquad \partial_{v_{ij}}f(\hat{\boldsymbol{v}}) = \frac{h(\hat{\boldsymbol{v}})}{v_{ij}}, \tag{4.135}$$

where $h(\hat{\boldsymbol{v}})$ is also function of the cross ratios, we see that

$$\tilde{\mathcal{W}}_{12}^{-+}f(\hat{\boldsymbol{v}}) = \sum_{i=3}^{n} \frac{v_{2i}}{v_{1i}} h_i(\hat{\boldsymbol{v}}) + \sum_{3 \le i < j}^{n} \frac{v_{2i}v_{2j}}{v_{ij}v_{12}} h_{ij}(\hat{\boldsymbol{v}})$$
(4.136)

where $h_i(\hat{\boldsymbol{v}})$ and $h_{ij}(\hat{\boldsymbol{v}})$ are specific functions of the cross ratios. Each term in the first sum then corresponds to a simplex integral with the shifts

$$\alpha_{2i} \to \alpha_{2i} - 1, \qquad \alpha_{1i} \to \alpha_{1i} + 1, \tag{4.137}$$

while each term in the second sum corresponds to a simplex integral with the shifts

$$\alpha_{2i} \to \alpha_{2i} - 1, \qquad \alpha_{2j} \to \alpha_{2j} - 1, \qquad \alpha_{ij} \to \alpha_{ij} + 1, \qquad \alpha_{12} \to \alpha_{12} + 1. \tag{4.138}$$

From (4.2), both (4.137) and (4.138) correspond to shifting $\Delta_1 \rightarrow \Delta_1 - 1$ and $\Delta_2 \rightarrow \Delta_2 + 1$ leaving all other operator dimensions fixed. The action of \mathcal{W}_{12}^{-+} on the simplex thus produces an appropriately shifted simplex integral, whose function of cross ratios is obtained through the action of the operator (4.133).

4.6 Discussion

Our analysis has furnished useful parametric representations for the general momentumspace conformal *n*-point function. Starting from the generalised simplex Feynman integral of [47, 46], we showed how all graph polynomials can be obtained from the corresponding Laplacian matrix, or the Gram matrix to which it reduces once momentum conservation has been enforced. With the graph polynomials to hand, all the usual scalar parametrisations of Feynman integrals can be adapted to represent the simplex solution. Only n(n-1)/2 integrals over Schwinger parameters remain to be performed – one for each leg of the simplex – in contrast to the (n-1)(n-2)d/2 scalar integrals we started with.

Building on the analogy between Feynman graph polynomials and those of electrical circuits, we then formulated a second class of parametric representations. For these, the integration variables represent the *effective* resistances between vertices of the simplex, rather than the conductivities (*i.e.*, the inverse Schwinger parameters) used previously. This change of variables immediately diagonalises the Schwinger exponential, expressing the *n*-point function as a standard Laplace transform of a product of polynomials raised to generalised powers. These polynomials correspond to the determinant and first minors of the Cayley-Menger matrix for the simplex, which plays an analogous role to the Gram matrix for this second class of parametrisations. From the form of these polynomials, new weight-shifting operators can immediately be constructed to raise the power of these polynomials, with further shift operators following by shadow conjugation. Besides shifting the scaling dimensions of external operators, these new weight-shifting operators raise the spacetime dimension by two. They therefore generalise the 3-point shift operators of [42, 85] to *n*-points, and constitute a distinct class of operators to those identified in [83].

Our results suggest several interesting directions for further pursuit:

• Given we now have weight-shifting operators that both preserve and raise the spacetime dimension, is it also possible to construct operators that *lower* the spacetime dimension? One approach we have explored, explained in appendix B.4, is to find so-called *Bernstein-Sato* operators which act to lower the powers to which the various polynomials of interest are raised. In this case, the relevant polynomials are the Cayley-Menger determinant and its minors appearing in the parametrisation (4.48). We found, for example, that replacing $v_{ij} \rightarrow \partial_{s_{ij}}$ in the Kirchhoff polynomial $\mathcal{U} = |g|$ yields an operator

$$\mathcal{B}_{|m|} = (|g|) \Big|_{v_{ij} \to \partial_{s_{ij}}}$$
(4.139)

which lowers by one the power to which the Cayley-Menger determinant is raised:

$$\mathcal{B}_{|m|} |m|^a = b_{|m|}(a) |m|^{a-1}, \qquad b_{|m|}(a) = -\prod_{k=1}^{n-1} (1-k-2a). \tag{4.140}$$

For the simplex representation (4.48), a is the parameter α given in (4.49) and so lowering α by one corresponds to sending $d \rightarrow d-2$ if all the operator dimensions are kept fixed. In principle, one would then integrate by parts to obtain an operator acting solely on the Schwinger exponential, which, due to its diagonal structure, could be translated into a differential operator in the external momenta. In practice, however, this approach is complicated by the presence of all the remaining powers of Cayley-Menger minors present in (4.48).

- In sections 4.4 and 4.5.1, we saw how the action of the special conformal Ward identity on the simplex reduces to a total derivative. This followed directly from the scalar parametric representation, without any recourse to the recursive arguments developed in [47, 46]. Nevertheless, these arguments, and the recursion relation between n- and (n + 1)-point simplices on which they are based, are of considerable interest in their own right and could be reformulated in the scalar-parametric language used here. The deletion/contraction relations of graph polynomials (see, e.g., [65]) and Kron reduction, corresponding to taking the Schur complement of a subset of vertices in the simplex Laplacian (see e.g., [117]), may also yield relevant identities.
- Starting from the general simplex solution, the arbitrary function of momentumspace cross ratios can be restricted by imposing additional conditions of interest: for example, dual conformal invariance [45, 95, 118, 119], or the Casimir equation for conformal blocks. For such investigations, the connection with position-space developed in section 4.5 provides a very simple link between the action of a given differential operator in the external momenta or coordinates, and its corresponding action on the arbitrary function of the simplex representation.
- For holographic *n*-point functions, bulk scalar Witten diagrams have the interesting property that their form is invariant under the action of a shadow transform on any of the external legs. In momentum space, shadow transforming the operator \mathcal{O}_i corresponds to multiplying the correlator by $p_i^{-2\beta_i}$, where $\beta_i = \Delta_i d/2$, which has the effect of replacing $\beta_i \to -\beta_i$ in the bulk-boundary propagator $z^{d/2} p_i^{\beta_i} K_{\beta_i}(p_i z)$. It would be interesting to understand the restriction this condition places on the function of cross-ratios appearing in the simplex representation.
- Finally, the parametric representations we have developed may provide a useful starting point for the construction of general spinning *n*-point correlators via the action of spin-raising operators [83, 29, 30], and for bootstrapping cosmological correlators in de Sitter spacetime.

Chapter 5

GKZ integrals and creation operators for Feynman and Witten diagrams

5.1 Introduction

It has long been suspected that Feynman integrals represent a multi-variable generalisation of hypergeometric functions [120, 121]. Recently [122–131], this connection has been sharpened by writing Feynman integrals as Gel'fand-Kapranov-Zelevinksy (GKZ) or \mathcal{A} -hypergeometric functions [132–135]. As shown in [123, 124], this can be achieved simply by expressing Feynman integrals in Lee-Pomeransky form [112], where only a single denominator polynomial appears, followed by uplifting to a higher-dimensional space of generalised momenta. \mathcal{A} -hypergeometric functions are well-studied in the mathematics literature [136–141] and satisfy a set of linear partial differential equations whose form can be read off in systematic fashion from a certain matrix – the \mathcal{A} -matrix – which encodes both the structure of the integral as well as all kinematic and spectral singularities.

A task of great practical interest is then to construct hypergeometric *shift operators* connecting integrals of different parameter values. These operators enable a known 'seed' integral to be converted, by simple differentiation, into an entire series of new integrals. For Feynman integrals, the parameters are typically the powers of various propagators and the spacetime dimension. Here we will also study Witten diagrams in anti-de Sitter spacetime for which the relevant parameters, besides the spacetime dimension, are the scaling dimensions of operators in the holographically dual conformal field theory.

While various techniques for constructing shift operators for Feynman integrals [70, 72, 142, 143, 69] and Witten diagrams [144, 85, 83, 145, 30, 119, 36] are known, the GKZ formalism offers a more powerful and unified approach. Besides the elementary shift operators, known as 'annihilation' operators in the mathematics literature, their *inverses* – a highly non-trivial class of operators known as 'creation' operators – can be systematically constructed [146–148]. Together, these creation and annihilation operators form a full set of shift operators connecting \mathcal{A} -hypergeometric functions of different parameter values, just as the ordinary creation and annihilation (or ladder) operators connect different eigenstates of the quantum harmonic oscillator.



Figure 5.1: The Newton polytope for a 3-point contact Witten diagram in momentum space is an octahedron as shown. At *n*-points, we obtain an *n*-dimensional cross-polytope. The spectral singularities consist of an infinite series of hyperplanes parallel to the facets of the Newton polytope, while the integral is convergent for parameter values lying inside the polytope. Identification of the singularities enables a systematic construction of all creation-type shift operators.

A key aim of this chapter is to show that creation operators can be constructed directly from knowledge of the spectral singularities of an \mathcal{A} -hypergeometric function, namely, the special set of parameter values for which the corresponding GKZ integral representation diverges. These singularities can be computed directly from the \mathcal{A} -matrix of the integral. Remarkably, they correspond geometrically to an infinite series of hyperplanes parallel to the co-dimension one facets of the Newton polytope associated with the integral's denominator [149, 150]. (See figure 5.1.) Standard convex hulling algorithms exist for computing such facets allowing a simple identification of all singularities.

To construct creation operators, we start with a pair of integrals connected by an annihilation operator. As we will review, this annihilator consists of a single derivative with respect to one of the GKZ generalised momenta. Specifically, we are interested in cases with parameters such that the starting integral is *divergent* while the resulting integral is *finite*. (To regulate divergences, we assume a dimensional scheme where parameters are infinitesimally shifted away from their singular values.) The divergences are thus projected out by the action of the annihilator. As the inverse of the annihilator, the creation operator must then produce the reverse shift, from the finite integral to the divergent one. Clearly, however, this cannot be achieved directly: the result of acting with a finite differential operator on a finite integral must necessarily be finite. Instead, the outcome must be a finite product of the divergent integral multiplied by a vanishing function of the parameters. This function, whose zeros serve to cancel out the divergence, is known as the *b-function* and holds the key to the construction of creation operators.

From a knowledge of the singular parameter values, we can predict the necessary zeros of the *b*-function and hence its minimal form as a polynomial. Then, acting on an integral with both the annihilator *and* the (as yet unknown) creation operator, we must recover the original integral multiplied by the *b*-function. In the GKZ formalism, however, any polynomial in the parameters can be traded for an equivalent polynomial in Euler operators acting on the generalised momenta. Applying this procedure to the *b*-function, the resulting differential operator must thus be factorisable into a product of the annihilation and the creation operator. As the annihilator is just a single derivative, this factorisation is easily performed (with the aid of a further set of PDEs known as the toric equations) revealing the identity of the creation operator. As a final step, one then

projects back from the higher-dimensional GKZ space of generalised momenta to that of the physical variables (the external momenta and masses), with the aid of an auxiliary set of Euler equations.

We hope this simple physical approach, based on the spectral singularities of the GKZ integral, will facilitate the application of creation operators to a range of physical systems. As an initial demonstration of the possibilities, we have used the formalism to construct new shift operators for a range of simple Feynman integrals, as well as Witten diagrams encoding momentum-space correlators in holographic conformal field theories. These latter objects are intimately related to cosmological correlators in de Sitter spacetime, and the new shift operators we construct can also be applied in this context. In particular, we have found new shift operators connecting both exchange and contact 4-point Witten diagrams, with arbitrary external scaling dimensions, to corresponding diagrams with shifted scaling dimensions but the same spacetime dimension. Until now, such operators were only available in the case where diagrams with *non-derivative* vertices are mapped to those with *derivative* vertices, and for a restricted set of scaling dimensions at that [30, 36]. In contrast, the new shift operators we find can be applied for any scaling dimensions, and moreover map non-derivative to non-derivative vertices. This enlarges the available arsenal of shift operators for Witten diagrams (and by extension, cosmological correlators), and as such is a useful and nontrivial result. We believe these examples provide a first proof of principle that the creation operator method, and the GKZ formalism more generally, holds promise for a variety of physical applications.

An outline of this chapter is as follows. Section 5.2 introduces \mathcal{A} -hypergeometric functions and the GKZ formalism. We summarise the PDEs these functions obey, their construction, and their invariance under affine reparametrisations. In section 5.3, we relate the spectral singularities of GKZ integrals to the Newton polytope of the denominator. In section 5.4, we introduce creation operators and detail their construction based on the spectral singularities of the integral. In section 5.5, we construct creation operators for 3- and 4-point contact Witten diagrams in momentum space, as well as a further set of shift operators that preserve the spacetime dimension. Using these results, we then derive novel shift operators for exchange diagrams. Section 5.6 constructs creation operators for a variety of simple Feynman integrals introducing the use of Gröbner bases and convex hulling algorithms to automate the computation. We conclude in section 5.7 with a summary of results and open directions. In the appendices we discuss the conversion of Feynman to GKZ integrals, creation operators for position-space contact Witten diagrams, and an extension of the minimal construction algorithm outlined above.

5.2 *A*-hypergeometric functions

The application of the GKZ formalism to Feynman integrals has been explored in a number of recent works [122–131]. In addition, many excellent expositions are available in the mathematics literature [136–141]. Here, we focus on providing a simple and self-contained summary of the key material needed to understand the construction of creation operators.

5.2.1 GKZ integrals

An \mathcal{A} -hypergeometric function (or equivalently, GKZ integral), is a multi-variable hypergeometric function depending on a set of real parameters $\boldsymbol{\gamma} = (\gamma_0, \gamma_1, \dots, \gamma_N)$ and independent variables $\boldsymbol{x} = (x_1, \dots, x_n)$, where $n \geq N + 1$. The integral takes the form

$$\mathcal{I}_{\gamma} = \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1}\Big) \mathcal{D}^{-\gamma_{0}},\tag{5.1}$$

where the 'denominator' \mathcal{D} can be expressed as a polynomial in the integration variables z_i . Every term in this polynomial is moreover multiplied by a *nonzero* coefficient x_i :

$$\mathcal{D} = \sum_{j=1}^{n} x_j \prod_{i=1}^{N} z_i^{a_{ij}}$$
(5.2)

The parameters $a_{ij} \in \mathbb{Z}^+$ specifying the powers can be assembled into an $N \times n$ matrix A,

$$(A)_{ij} = a_{ij}.\tag{5.3}$$

Thus, the *j*th term in the denominator \mathcal{D} corresponds to the column *j* of the matrix *A*, whose entries are then the powers of the variables z_i appearing in that particular term. (We will return to the relation between this matrix *A* and the larger *A*-matrix shortly.)

For Feynman integrals, it is useful to consider the Lee-Pomeransky representation [112] in which the denominator $\mathcal{G} = \mathcal{U} + \mathcal{F}$ is formed from the sum of the first and second Symanzik polynomials \mathcal{U} and \mathcal{F} . To uplift this to the GKZ integral (5.1), we simply promote the coefficient of every term in \mathcal{G} to a generalised independent variable x_j [123, 124], as summarised in appendix C.1. The original Lee-Pomeransky integral can then be restored by returning the x_j to their physical values, namely, unity for any of the terms in \mathcal{U} , and the appropriate function of the masses and external momenta for every term in \mathcal{F} .

Example: As discussed in appendix C.1, the massless triangle Feynman integral

$$I = \int \frac{\mathrm{d}^{d} \boldsymbol{q}}{(2\pi)^{d}} \frac{1}{q^{2\gamma_{3}} |\boldsymbol{q} - \boldsymbol{p}_{1}|^{2\gamma_{2}} |\boldsymbol{q} + \boldsymbol{p}_{2}|^{2\gamma_{1}}}$$
(5.4)

has the Lee-Pomeransky representation

$$I = c_{\gamma} \left(\prod_{i=1}^{3} \int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1}\right) (p_{1}^{2} z_{2} z_{3} + p_{2}^{2} z_{1} z_{3} + p_{3}^{2} z_{1} z_{2} + z_{1} + z_{2} + z_{3})^{-d/2}$$
(5.5)

where the coefficient

$$c_{\gamma} = (4\pi)^{-d/2} \frac{\Gamma(d/2)}{\Gamma(d-\gamma_t) \prod_{i=1}^3 \Gamma(\gamma_i)}, \qquad \gamma_t = \sum_{i=1}^3 \gamma_i.$$
(5.6)

The corresponding GKZ integral is

$$\mathcal{I}_{\gamma} = \Big(\prod_{i=1}^{3} \int_{0}^{\infty} \mathrm{d}z_{i} z_{i}^{\gamma_{i}-1}\Big) \mathcal{D}^{-\gamma_{0}}$$
(5.7)

where the denominator

$$\mathcal{D} = x_1 z_2 z_3 + x_2 z_1 z_3 + x_3 z_1 z_2 + x_4 z_1 + x_5 z_2 + x_6 z_3 \tag{5.8}$$

corresponds to the matrix

$$A = \begin{pmatrix} 0 & 1 & 1 & 1 & 0 & 0 \\ 1 & 0 & 1 & 0 & 1 & 0 \\ 1 & 1 & 0 & 0 & 0 & 1 \end{pmatrix}.$$
 (5.9)

To recover the original Lee-Pomeransky integral, we project to the physical subspace

$$\boldsymbol{x} = (p_1^2, p_2^2, p_3^2, 1, 1, 1), \qquad \boldsymbol{\gamma} = (d/2, \gamma_1, \gamma_2, \gamma_3),$$
 (5.10)

after which $I = c_{\gamma} \mathcal{I}_{\gamma}$.

5.2.2 The Euler and toric equations

The primary advantage of uplifting from the original masses and momenta to the generalised GKZ space parametrised by the variables \boldsymbol{x} is that the integral now obeys a systematic set of linear partial differential equations. These can be grouped into two categories, known as the Euler equations and the toric equations.

Euler equations

The Euler equations arise from integrating by parts with respect to the variables z_i , under the assumption that all boundary terms vanish. For z_1 , for example, we have

$$0 = \int_{0}^{\infty} dz_{1} \frac{\partial}{\partial z_{1}} \left(z_{1}^{\gamma_{1}} \left(\prod_{i=2}^{N} \int_{0}^{\infty} dz_{i} z_{i}^{\gamma_{i}-1} \right) \mathcal{D}^{-\gamma_{0}} \right)$$
$$= \gamma_{1} \mathcal{I}_{\gamma} + \left(\prod_{i=1}^{N} \int_{0}^{\infty} dz_{i} z_{i}^{\gamma_{i}-1} \right) z_{1} \frac{\partial}{\partial z_{1}} \mathcal{D}^{-\gamma_{0}}.$$
(5.11)

In the second term here, we can trade derivatives with respect to the integration variable z_1 for derivatives with respect to the external variables x_j :

$$z_1 \frac{\partial}{\partial z_1} \mathcal{D}^{-\gamma_0} = -\gamma_0 \mathcal{D}^{-\gamma_0 - 1} \Big(\sum_{j=1}^n a_{1j} x_j \prod_{i=1}^N z_i^{a_{ij}} \Big) = \Big(\sum_{j=1}^n a_{1j} \theta_j \Big) \mathcal{D}^{-\gamma_0}$$
(5.12)

where, here and throughout the chapter, we define the Euler operators

$$\theta_j = x_j \frac{\partial}{\partial x_j}, \qquad j = 1, \dots, n.$$
(5.13)

Pulling these Euler operators outside the integrals, we obtain the equation

$$0 = \left(\gamma_1 + \sum_{j=1}^n a_{1j}\theta_j\right) \mathcal{I}_{\gamma}.$$
(5.14)

Repeating this exercise for the remaining z_i then leads to the set of Euler equations

$$0 = \left(\gamma_i + \sum_{j=1}^n a_{ij}\theta_j\right) \mathcal{I}_{\gamma}, \qquad i = 1, \dots, N.$$
(5.15)

We are not quite done, however, since in addition we have the general identity

$$\Big(\sum_{j=1}^{n} \theta_j\Big) \mathcal{D}^{-\gamma_0} = -\gamma_0 \mathcal{D}^{-\gamma_0}$$
(5.16)

which, when applied to the GKZ integral, yields

$$0 = \left(\gamma_0 + \sum_{j=1}^n \theta_j\right) \mathcal{I}_{\gamma}.$$
(5.17)

This equation is effectively a dilatation Ward identity (or DWI, as we will use for short) encoding the scaling behaviour of the GKZ integral under a dilatation $\boldsymbol{x} \to \lambda \boldsymbol{x}$ of the external variables.

Evidently this dilatation Ward identity can be placed on the same footing as the Euler equations (5.15) by enlarging the matrix A to include a top row consisting of all 1s. This construction defines the A-matrix mentioned in the introduction,

$$\mathcal{A} = \begin{pmatrix} \mathbf{1} \\ A \end{pmatrix},\tag{5.18}$$

where **1** is the *n*-dimensional row vector with all-1 entries, or equivalently,

$$(\mathcal{A})_{0j} = 1,$$
 $(\mathcal{A})_{ij} = a_{ij},$ $i = 1, \dots, N,$ $j = 1, \dots, n,$ (5.19)

where we henceforth adopt the convention that the top row of \mathcal{A} always carries index 0. The \mathcal{A} -matrix is thus $(N+1) \times n$ dimensional, and the Euler equations and DWI together correspond to the (N+1) equations

$$0 = \left(\gamma_i + \sum_{j=1}^n \mathcal{A}_{ij}\theta_j\right)\mathcal{I}_{\gamma}, \qquad i = 0, \dots, N.$$
(5.20)

This is in effect a single matrix equation,

$$0 = \left(\boldsymbol{\gamma} + \boldsymbol{\mathcal{A}} \cdot \boldsymbol{\theta}\right) \boldsymbol{\mathcal{I}}_{\boldsymbol{\gamma}},\tag{5.21}$$

regarding $\boldsymbol{\theta} = (\theta_1, \dots, \theta_n)^T$ and $\boldsymbol{\gamma} = (\gamma_0, \gamma_1, \dots, \gamma_N)^T$ as *n*- and (N+1)-component column vectors respectively.

Example: Returning to the massless triangle integral above, the \mathcal{A} -matrix is

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 & 1 & 1 \\ 0 & 1 & 1 & 1 & 0 & 0 \\ 1 & 0 & 1 & 0 & 1 & 0 \\ 1 & 1 & 0 & 0 & 0 & 1 \end{pmatrix}$$
(5.22)

and the GKZ integral satisfies the Euler equations

$$0 = (\gamma_1 + \theta_2 + \theta_3 + \theta_4)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_2 + \theta_1 + \theta_3 + \theta_5)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_3 + \theta_1 + \theta_2 + \theta_6)\mathcal{I}_{\gamma}$$
(5.23)

and DWI

$$0 = (\gamma_0 + \sum_{j=1}^6 \theta_j) \mathcal{I}_{\gamma}.$$
(5.24)

Notice the form of these equations can be directly read off from the rows of the \mathcal{A} -matrix.

Toric equations

The toric equations arise from vectors in the *kernel* of the \mathcal{A} -matrix, and are closely related to the corresponding toric ideal [140].¹ Their origin can easily be grasped using the example of the massless triangle integral above. Defining

$$\partial_j = \frac{\partial}{\partial x_j}, \qquad j = 1, \dots, n$$
 (5.25)

in all that follows, the denominator (5.8) obeys the relations

$$\partial_1 \partial_4 \mathcal{D}^{-\gamma_0} = \partial_2 \partial_5 \mathcal{D}^{-\gamma_0} = \partial_3 \partial_6 \mathcal{D}^{-\gamma_0} = -\gamma_0 (-\gamma_0 - 1) z_1 z_2 z_3 \mathcal{D}^{-\gamma_0 - 2}, \qquad (5.26)$$

giving rise to the two independent (toric) equations

$$0 = (\partial_1 \partial_4 - \partial_3 \partial_6) \mathcal{I}_{\gamma}, \qquad 0 = (\partial_2 \partial_5 - \partial_3 \partial_6) \mathcal{I}_{\gamma}. \tag{5.27}$$

For comparison, the kernel of the \mathcal{A} -matrix (5.22) is spanned by two independent vectors, $\boldsymbol{u}_{(1)}$ and $\boldsymbol{u}_{(2)}$, which we can choose to be

$$\boldsymbol{u}_{(1)} = (1, 0, -1, 1, 0, -1)^T, \qquad \boldsymbol{u}_{(2)} = (1, -1, 0, 1, -1, 0)^T.$$
 (5.28)

Notice that since the top row of the \mathcal{A} -matrix is all 1s, the sum of the components of any kernel vector is always zero. There is now a one-to-one match between kernel vectors and toric equations (5.27) as follows. First, for each kernel vector \boldsymbol{u} , we form a vector \boldsymbol{u}^+ composed only of the *positive* components of \boldsymbol{u} , and a vector \boldsymbol{u}^- composed of only the *negative* components. The components of \boldsymbol{u}^{\pm} , for each $j = 1, \ldots, n$, are thus

$$u_j^{\pm} = \max(\pm u_j, 0).$$
 (5.29)

¹The kernel is the space of vectors u such that $\mathcal{A} \cdot u = 0$, obtained *e.g.*, via NullSpace[\mathcal{A}] in Mathematica. The full toric ideal, though not needed here, can be constructed using *Singular* [151]: see section 5.6.2.

By inspection, the toric equation corresponding to the kernel vector $\boldsymbol{u} = \boldsymbol{u}^+ - \boldsymbol{u}^-$ is now

$$0 = \left(\prod_{j=1}^{n} \partial_j^{u_j^+} - \prod_{j=1}^{n} \partial_j^{u_j^-}\right) \mathcal{I}_{\gamma}.$$
(5.30)

For example, for $\boldsymbol{u}_{(1)}$ in (5.28), $\boldsymbol{u}_{(1)}^+ = (1, 0, 0, 1, 0, 0)^T$ while $\boldsymbol{u}_{(1)}^- = (0, 0, 1, 0, 0, 1)^T$ hence (5.30) reduces to the first equation in (5.27).

Some investigation shows this construction is a general one. First, the action of each differential operator is

$$\prod_{j=1}^{n} \partial_{j}^{u_{j}^{\pm}} \mathcal{D}^{-\gamma_{0}} = (-\gamma_{0})(-\gamma_{0}-1)\dots(-\gamma_{0}-\mathfrak{u}^{\pm}+1)\mathcal{D}^{-\gamma_{0}-\mathfrak{u}^{\pm}} \left(\prod_{i=1}^{N} z_{i}^{\sum_{j=1}^{n} a_{ij}u_{j}^{\pm}}\right)$$
(5.31)

where $\mathfrak{u}^{\pm} = \sum_{j=1}^{n} u_j^{\pm}$. Moreover, since the sum of components in any kernel vector vanishes (as the top row of the \mathcal{A} -matrix is all 1s), we have that $\mathfrak{u}^+ = \mathfrak{u}^- = \mathfrak{u}$. Thus,

$$\left(\prod_{j=1}^{n} \partial_{j}^{u_{j}^{+}} - \prod_{j=1}^{n} \partial_{j}^{u_{j}^{-}}\right) \mathcal{I}_{\gamma}$$

$$= (-\gamma_{0})(-\gamma_{0} - 1) \dots (-\gamma_{0} - \mathfrak{u} + 1) \mathcal{D}^{-\gamma_{0} - \mathfrak{u}} \left(\prod_{i=1}^{N} z_{i}^{\sum_{j=1}^{n} a_{ij}u_{j}^{+}} - \prod_{i=1}^{N} z_{i}^{\sum_{j=1}^{n} a_{ij}u_{j}^{-}}\right).$$
(5.32)

However, for any kernel vector we have $\mathcal{A} \cdot \boldsymbol{u} = \mathcal{A} \cdot (\boldsymbol{u}^+ - \boldsymbol{u}^-) = \boldsymbol{0}$ and hence

$$\sum_{j=1}^{n} a_{ij} u_j^+ = \sum_{j=1}^{n} a_{ij} u_j^-, \qquad i = 1, \dots, N.$$
(5.33)

The two terms appearing within the final factor of (5.32) are thus exactly equal producing a cancellation. In general, as the \mathcal{A} -matrix is $(N+1) \times n$, there are (n-N-1) independent vectors in the kernel, and hence this same number of independent toric equations.

To summarise, given a GKZ integral defined by an \mathcal{A} -matrix and parameters γ , we have two sets of linear partial differential equations: the Euler equations (and DWI) (5.20), and the toric equations (5.30). We can also go in reverse: the Euler equations and DWI fix γ and the \mathcal{A} -matrix, and hence the toric equations and the GKZ integral. Note the Euler equations all commute among themselves, as do the toric equations, but an Euler and a toric equation do not in general commute.

5.2.3 Projection to physical variables

The systematic structure of the Euler and toric equations above is a consequence of uplifting from the Lee-Pomeransky to the GKZ denominator (5.2). To recover a set of PDEs satisfied by the original Lee-Pomeransky integral we need to reverse this process. This requires projecting the Euler and toric equations back to the *physical hypersurface* where the \boldsymbol{x} variables take their true physical values. Derivatives in directions not tangential to this hypersurface (which therefore cannot be expressed purely in terms of physical variables) can be exchanged for purely tangential derivatives through use of the Euler equations
and DWI. Together these provide N + 1 equations, and so for all unphysical (*i.e.*, non-tangential) derivatives to be removable requires the original Lee-Pomeransky polynomial to contain at least n - N - 1 independent physical variables (*i.e.*, masses and external momenta). This will generally be the case for the examples we consider, but does not hold universally – particularly for higher-loop Feynman integrals – as we discuss in section 5.7.

Example: For the massless triangle integral, the physical hypersurface is the 3-dimensional subspace spanned by the momenta in (5.10), namely $x_1 = p_1^2$, $x_2 = p_2^2$ and $x_3 = p_3^2$, with $x_4 = x_5 = x_6 = 1$. On this hypersurface, the Euler equations (5.23) reduce to

$$0 = (\gamma_1 + \theta_2 + \theta_3 + \partial_4)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_2 + \theta_1 + \theta_3 + \partial_5)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_3 + \theta_1 + \theta_2 + \partial_6)\mathcal{I}_{\gamma}, \quad (5.34)$$

where, as always, $\partial_j = \partial/\partial x_j$. These equations allow us to eliminate the unphysical derivatives ∂_4 , ∂_5 and ∂_6 from all remaining equations in which they appear linearly.² For example, evaluating the first toric equation in (5.27) on the physical hypersurface,

$$0 = (\partial_1 \partial_4 - \partial_3 \partial_6) \mathcal{I}_{\gamma}$$

= $\left(\partial_1 (-\gamma_1 - \theta_2 - \theta_3) - \partial_3 (-\gamma_3 - \theta_1 - \theta_2)\right) \mathcal{I}_{\gamma}$
= $\frac{1}{4} \left[-\left(2\gamma_1 + p_2 \frac{\partial}{\partial p_2} + p_3 \frac{\partial}{\partial p_3}\right) \frac{1}{p_1} \frac{\partial}{\partial p_1} + \left(2\gamma_3 + p_1 \frac{\partial}{\partial p_1} + p_2 \frac{\partial}{\partial p_2}\right) \frac{1}{p_3} \frac{\partial}{\partial p_3} \right] \mathcal{I}_{\gamma}, \quad (5.35)$

while for the second toric equation,

$$0 = (\partial_2 \partial_5 - \partial_3 \partial_6) \mathcal{I}_{\gamma}$$

= $\left(\partial_2 (-\gamma_2 - \theta_1 - \theta_3) - \partial_3 (-\gamma_3 - \theta_1 - \theta_2)\right) \mathcal{I}_{\gamma}$
= $\frac{1}{4} \left[-\left(2\gamma_2 + p_1 \frac{\partial}{\partial p_1} + p_3 \frac{\partial}{\partial p_3}\right) \frac{1}{p_2} \frac{\partial}{\partial p_2} + \left(2\gamma_3 + p_1 \frac{\partial}{\partial p_1} + p_2 \frac{\partial}{\partial p_2}\right) \frac{1}{p_3} \frac{\partial}{\partial p_3} \right] \mathcal{I}_{\gamma}.$ (5.36)

Finally, on the physical hypersurface, the DWI (5.24) reduces to

$$0 = \left(\frac{d}{2} + \theta_1 + \theta_2 + \theta_3 + \partial_4 + \partial_5 + \partial_6\right) \mathcal{I}_{\gamma}$$

= $\left(\frac{d}{2} - \gamma_1 - \gamma_2 - \gamma_3 - \theta_1 - \theta_2 - \theta_3\right) \mathcal{I}_{\gamma}$
= $\frac{1}{2} \left(d - 2\gamma_1 - 2\gamma_2 - 2\gamma_3 - p_1 \frac{\partial}{\partial p_1} - p_2 \frac{\partial}{\partial p_2} - p_3 \frac{\partial}{\partial p_3}\right) \mathcal{I}_{\gamma}$ (5.37)

Equations (5.35)-(5.37) involve only physical variables, namely, the momentum magnitudes.

5.2.4 Affine reparametrisations

As we have seen, the set of Euler equations associated with a given GKZ integral can be read off from the rows of the \mathcal{A} -matrix: in the *i*th Euler equation (5.15), the coefficient of

²More generally, we can rewrite $\partial_4^m = x_4^{-m}\theta_4(\theta_4 - 1)\dots(\theta_4 - m + 1)$, etc., then use the full Euler equations to eliminate θ_4 , θ_5 and θ_6 before setting $x_4 = x_5 = x_6 = 1$. Alternatively, we can supplement (5.34) with derivatives of the Euler equations (and DWI) evaluated on the physical hypersurface.

the operator θ_j is $a_{ij} = (\mathcal{A})_{ij}$ where $1 \leq i \leq N$ and $1 \leq j \leq n$. (Recall we are labelling the top all-1s row of the \mathcal{A} -matrix as i = 0.) Viewed in reverse, the set of Euler equations determines both the \mathcal{A} -matrix and the set of parameters γ , and hence the GKZ integral.

What happens if we now form a *new* set of Euler equations by taking linear combinations of the old ones? In the process, we could simultaneously add to each Euler equation some multiple of the DWI. Together, these operations correspond to left-multiplying the \mathcal{A} -matrix by an $(N+1) \times (N+1)$ matrix

$$\mathcal{M} = \begin{pmatrix} 1 & \mathbf{0} \\ \mathbf{b} & M \end{pmatrix},\tag{5.38}$$

where **0** is an N-dimensional row vector of zeros, \boldsymbol{b} is an N-dimensional column vector and M an $N \times N$ matrix. This yields

$$\mathcal{A}' = \mathcal{M}\mathcal{A} = \begin{pmatrix} 1 & \mathbf{0} \\ \mathbf{b} & M \end{pmatrix} \begin{pmatrix} \mathbf{1} \\ A \end{pmatrix} = \begin{pmatrix} \mathbf{1} \\ A' \end{pmatrix}, \tag{5.39}$$

where the components of A undergo the affine transformation

$$(A')_{ij} = a'_{ij} = b_i + \sum_{k=1}^{N} m_{ik} a_{kj}.$$
(5.40)

The new set of Euler equations now corresponds to the rows of \mathcal{A}' : the *i*th new Euler equation is the sum of m_{ik} times the *k*th old Euler equation plus b_i times the DWI (for which the coefficient of every θ_j is one). In order to have $a'_{ij} \in \mathbb{Z}^+$, so as to form a new denominator polynomial \mathcal{D}' via (5.2), we will restrict the entries of \mathcal{M} to $m_{ij} \in \mathbb{Z}^+$ and $b_i \in \mathbb{Z}^+$. Note the transformation (5.39) leaves the DWI unchanged.

The new set of Euler equations now takes the form

$$0 = \left(\boldsymbol{\gamma}' + \boldsymbol{\mathcal{A}}' \cdot \boldsymbol{\theta}\right) \boldsymbol{\mathcal{I}}_{\boldsymbol{\gamma}'},\tag{5.41}$$

where

$$\boldsymbol{\gamma}' = \begin{pmatrix} \gamma_0 \\ \gamma_1' \\ \vdots \\ \gamma_N' \end{pmatrix} = \begin{pmatrix} 1 & \mathbf{0} \\ \boldsymbol{b} & M \end{pmatrix} \begin{pmatrix} \gamma_0 \\ \gamma_1 \\ \vdots \\ \gamma_N \end{pmatrix} = \mathcal{M} \boldsymbol{\gamma}$$
(5.42)

so that $\gamma'_i = \gamma_0 b_i + \sum_{k=1}^N m_{ik} \gamma_k$ for $1 \le i \le N$ while the DWI (5.17) remains unchanged. Provided that $\det(\mathcal{M})$ is nonzero, the toric equations are also unchanged since the kernel of \mathcal{A} is preserved under multiplication by an invertible matrix.

What is now the relation of this new GKZ integral, defined by \mathcal{A}' , to the original? The new integral is

$$\mathcal{I}_{\boldsymbol{\gamma}'} = \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z'_{i} (z'_{i})^{\gamma'_{i}-1} \Big) (\mathcal{D}')^{-\gamma_{0}}, \tag{5.43}$$

where

$$\mathcal{D}' = \sum_{j=1}^{n} x_j \prod_{i=1}^{N} (z'_i)^{a'_{ij}}.$$
(5.44)

Using (5.40), and making the identification

$$z_k = \prod_{i=1}^N (z'_i)^{m_{ik}},\tag{5.45}$$

we find

$$\mathcal{D}' = \sum_{j=1}^{n} x_j \prod_{i=1}^{N} (z'_i)^{b_i + \sum_{k=1}^{N} m_{ik} a_{kj}} = \left(\prod_{l=1}^{N} (z'_l)^{b_l}\right) \left(\sum_{j=1}^{n} x_j \prod_{i=1}^{N} \prod_{k=1}^{N} (z'_i)^{m_{ik} a_{kj}}\right)$$
$$= \left(\prod_{l=1}^{N} (z'_l)^{b_l}\right) \left(\sum_{j=1}^{n} x_j \prod_{k=1}^{N} z_k^{a_{kj}}\right) = \left(\prod_{l=1}^{N} (z'_l)^{b_l}\right) \mathcal{D}.$$
(5.46)

Moving the factor of $\prod_{l=1}^{N} (z'_l)^{b_l}$ from the denominator to the numerator and using (5.42) then gives

$$\mathcal{I}_{\gamma'} = \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z'_{i}(z'_{i})^{\gamma'_{i}-\gamma_{0}b_{i}-1}\Big)\mathcal{D}^{-\gamma_{0}} = \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z'_{i}(z'_{i})^{\sum_{k=1}^{N} m_{ik}\gamma_{k}-1}\Big)\mathcal{D}^{-\gamma_{0}} \\
= \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \frac{\mathrm{d}z'_{i}}{z'_{i}} \prod_{k=1}^{N} (z'_{i})^{m_{ik}\gamma_{k}}\Big)\mathcal{D}^{-\gamma_{0}} = \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \frac{\mathrm{d}z'_{i}}{z'_{i}} z^{\gamma_{i}}_{i}\Big)\mathcal{D}^{-\gamma_{0}}.$$
(5.47)

Finally, since

$$\frac{\mathrm{d}z_i}{z_i} = \sum_{j=1}^N m_{ji} \frac{\mathrm{d}z'_j}{z'_j}, \qquad \prod_i \int_0^\infty \frac{\mathrm{d}z_i}{z_i} = |\det M| \prod_i \int_0^\infty \frac{\mathrm{d}z'_i}{z'_i}, \qquad (5.48)$$

we find

$$\mathcal{I}_{\gamma'} = |\det M|^{-1} \mathcal{I}_{\gamma}. \tag{5.49}$$

Thus, choosing a new basis for the Euler equations by taking linear combinations of the old Euler equations and the DWI only rescales the GKZ integral by a constant factor. As the GKZ system of equations is linear, this overall scaling is in any case not fixed and the solution is effectively unchanged.

Example: The affine reparametrisation above can be used to show the equivalence of the massless triangle integral (5.4) with the *triple-K integral* (see also [42, 46])

$$I_{\alpha,\{\beta_1,\beta_2,\beta_3\}} = \int_0^\infty dz \, z^\alpha \prod_{i=1}^3 p_i^{\beta_i} K_{\beta_i}(p_i z).$$
(5.50)

For $\alpha = d/2 - 1$ and $\beta_i = \Delta_i - d/2$, this integral represents the momentum-space 3-point function of scalars \mathcal{O}_{Δ_i} in any *d*-dimensional CFT. The triple-*K* integral can be put into GKZ form by first Schwinger parametrising the modified Bessel functions as

$$p_i^{\beta_i} K_{\beta_i}(p_i z) = \frac{1}{2} \int_0^\infty \mathrm{d}z_i' \, (z_i')^{\beta_i - 1} \exp\left[-\frac{z}{2} \left(z_i' + \frac{p_i^2}{z_i'}\right)\right]$$
(5.51)

then performing the z integral. This gives

$$I_{\alpha,\{\beta_1,\beta_2,\beta_3\}} = 2^{\alpha-2} \Gamma(\alpha+1) \Big(\prod_{i=1}^3 \int_0^\infty \mathrm{d}z'_i \, (z'_i)^{\beta_i-1} \Big) \Big[\sum_{j=1}^3 \Big(z'_j + \frac{p_j^2}{z'_j} \Big) \Big]^{-\alpha-1} \tag{5.52}$$

which uplifts to the GKZ integral

$$I_{\alpha,\{\beta_1,\beta_2,\beta_3\}} = 2^{\alpha-2} \Gamma(\alpha+1) \Big(\prod_{i=1}^3 \int_0^\infty \mathrm{d}z_i'(z_i')^{\gamma_i'-1} \Big) (\mathcal{D}')^{-\gamma_0'}$$
(5.53)

where

$$\mathcal{D}' = \frac{x_1}{z_1'} + \frac{x_2}{z_2'} + \frac{x_3}{z_3'} + x_4 z_1' + x_5 z_2' + x_6 z_3'.$$
(5.54)

The physical hypersurface (*i.e.*, the original triple-K integral) corresponds to

$$\gamma'_i = \beta_i, \qquad \gamma'_0 = \alpha + 1, \qquad \boldsymbol{x} = (p_1^2, p_2^2, p_3^2, 1, 1, 1).$$
 (5.55)

Here, we are using primes to distinguish the parameters of the triple-K integral from those of the massless triangle integral earlier. Also, while the denominator (5.54) is not a polynomial, this simple generalisation will nevertheless turn out to be the most convenient representation for us later.³ The \mathcal{A} -matrix corresponding to the triple-K integral is then

$$\mathcal{A}_{3\mathrm{K}} = \begin{pmatrix} 1 & 1 & 1 & 1 & 1 & 1 \\ -1 & 0 & 0 & 1 & 0 & 0 \\ 0 & -1 & 0 & 0 & 1 & 0 \\ 0 & 0 & -1 & 0 & 0 & 1 \end{pmatrix}.$$
 (5.56)

Comparing with the massless triangle \mathcal{A} -matrix (5.22), we find that

$$\mathcal{MA}_{\text{triangle}} = \mathcal{A}_{3\text{K}},\tag{5.57}$$

where

$$\mathcal{M} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 1 & 0 & -1 & -1 \\ 1 & -1 & 0 & -1 \\ 1 & -1 & -1 & 0 \end{pmatrix}.$$
 (5.58)

³ Should a purely polynomial denominator be required, one can simply pull out an overall factor of $(z'_1 z'_2 z'_3)^{-1}$ from the right-hand side of (5.54) then transfer this to the numerator by shifting the γ'_i .

The parameters of the triangle integral are connected to those of the triple-K integral by

$$\mathcal{M}\boldsymbol{\gamma}_{\text{triangle}} = \mathcal{M} \begin{pmatrix} d/2\\ \gamma_1\\ \gamma_2\\ \gamma_3 \end{pmatrix} = \begin{pmatrix} d/2\\ d/2 - \gamma_2 - \gamma_3\\ d/2 - \gamma_1 - \gamma_3\\ d/2 - \gamma_1 - \gamma_2 \end{pmatrix} = \begin{pmatrix} \alpha + 1\\ \beta_1\\ \beta_2\\ \beta_3 \end{pmatrix} = \boldsymbol{\gamma}_{3K}.$$
(5.59)

Putting everything together, from (5.49) with det $\mathcal{M} = 2$ and (5.6), we have

$$I_{d/2-1,\{d/2-\gamma_2-\gamma_3,d/2-\gamma_1-\gamma_3,d/2-\gamma_1-\gamma_2\}} = C' \int \frac{\mathrm{d}^d \boldsymbol{q}}{(2\pi)^d} \frac{1}{q^{2\gamma_3} |\boldsymbol{q}-\boldsymbol{p}_1|^{2\gamma_2} |\boldsymbol{q}+\boldsymbol{p}_2|^{2\gamma_1}}$$
(5.60)

where

$$C' = \pi^{d/2} 2^{3d/2 - 4} \Gamma(d - \gamma_t) \prod_{i=1}^3 \Gamma(\gamma_i).$$
(5.61)

As we saw above, the matrix multiplication here is just a slick way of executing the change of variables

$$z_1 = \frac{1}{z'_2 z'_3}, \qquad z_2 = \frac{1}{z'_1 z'_3}, \qquad z_3 = \frac{1}{z'_1 z'_2},$$
 (5.62)

on the triangle GKZ representation, followed by moving a factor of $(z'_1 z'_2 z'_3)^{-\gamma_0}$ from the denominator to the numerator.

5.3 Spectral singularities and the Newton polytope

We now turn to examine the singularities of GKZ integrals arising for special values of the parameters γ . As we will see, these can be viewed geometrically in terms of the *Newton* polytope of the GKZ denominator \mathcal{D} .

5.3.1 The Newton polytope

A defining feature of the GKZ representation is that only a single denominator (5.2) is present:

$$\mathcal{D} = \sum_{j=1}^{n} x_j \prod_{i=1}^{N} z_i^{a_{ij}}.$$
(5.63)

The exponents of the *j*th term in this denominator define a vector a_j living in an N-dimensional space, whose components are

$$(a_j)_i = a_{ij}, \qquad i = 1, \dots N.$$
 (5.64)

Thus, a_j is the *j*th column of the \mathcal{A} -matrix after stripping off the top row of all 1s. Constructing the convex hull of these exponent vectors then defines the *N*-dimensional Newton polytope of \mathcal{D} :

Newt(
$$\mathcal{D}$$
) = $\sum_{j=1}^{n} \alpha_j \boldsymbol{a}_j$, with $\sum_{j=1}^{n} \alpha_j = 1$, $\alpha_j \ge 0 \ \forall j$. (5.65)



Figure 5.2: The Newton polytopes corresponding to the denominators of the triple-K integral (5.54) (left) and the massless triangle integral (5.8) (right).

For the denominator (5.54) of the triple-K integral, for example, we obtain the regular octahedron shown on the left of figure 5.2. For the denominator of the massless triangle integral (5.8), we also obtain an octahedron, but now with vertices as shown on the right of the figure. The vertices of each polytope are related by the affine transformation (5.40),

$$\boldsymbol{a}_{j}^{(3K)} = \boldsymbol{b} + M \boldsymbol{a}_{j}^{(\text{triangle})}, \qquad j = 1, \dots, 6$$
(5.66)

where, from (5.38) and (5.58),

$$\boldsymbol{b} = \begin{pmatrix} 1\\1\\1 \end{pmatrix}, \qquad M = \begin{pmatrix} 0 & -1 & -1\\-1 & 0 & -1\\-1 & -1 & 0 \end{pmatrix}.$$
 (5.67)

As we saw above, for any two \mathcal{A} -matrices (and hence any two Newton polytopes) related by an affine transformation, the corresponding GKZ integrals are proportional to each other and satisfy the same system of equations (*i.e.*, DWI, Euler and toric equations). Thus, Newton polytopes such as these related by affine transformations are effectively equivalent.

5.3.2 Spectral singularities

The physical significance of the Newton polytope becomes apparent when we consider the *spectral singularities* of the GKZ integral. These are the divergences that arise for special values of the parameters γ , with general kinematics, and are distinct from the *kinematic* (or Landau) singularities (discussed, *e.g.*, in [125]) which arise for general γ but special kinematics. Remarkably, it can be shown [149, 150] that the spectral singularities are closely related to the facets (*i.e.*, co-dimension one faces) of the Newton polytope. As this polytope lives in an N-dimensional space, let us first define the N-dimensional parameter vector

$$\hat{\boldsymbol{\gamma}} = (\gamma_1, \dots, \gamma_N)^T, \qquad (5.68)$$

where the hat serves to distinguish from the (N + 1)-dimensional parameter vector $\boldsymbol{\gamma} = (\gamma_0, \hat{\boldsymbol{\gamma}})^T$. In addition, we define the *rescaled* Newton polytope to be the convex hull of

the vertex vectors $\gamma_0 a_j$. This corresponds to a linear rescaling⁴ of the original Newton polytope by a factor of γ_0 . The GKZ integral is then finite for all parameter values $\hat{\gamma}$ lying *within* this rescaled Newton polytope. On the hyperplanes corresponding to the facets of the rescaled Newton polytope, as well as on an infinite set of further hyperplanes both parallel and exterior to these facets, the integral is singular.

An exact formula for all singular hyperplanes will be derived below in (5.109). The location of these singularities will then be the main ingredient in our subsequent construction of creation operators. Two key steps are needed to establish the result (5.109). The first is to show that the GKZ integral converges for all $\hat{\gamma}$ values lying in the interior of the rescaled Newton polytope. Rather than recounting the formal proof of [149, 150], we will instead pursue a more informal approach based on a tropical analysis of the GKZ integral [152, 153]. Many closely related constructions appear in sector decomposition, see *e.g.*, [154, 155]. The second step in the analysis is to construct a series of meromorphic continuations across each of the singular hyperplanes. This can be achieved by a scaling argument due to [149, 150]. Here, we present a further variation of this argument involving a special linear combination of the Euler equations and DWI.

Example: As an initial check of the picture above, we recall that the spectral singularities of the triple-K integral (5.50) are already known from conformal field theory [45].⁵ The condition for the triple-K integral $I_{\alpha,\{\beta_1,\beta_2,\beta_3\}}$ to be singular is

$$\alpha + 1 \pm \beta_1 \pm \beta_2 \pm \beta_3 = -2m, \qquad m \in \mathbb{Z}^+$$
(5.69)

where any independent choice of the three \pm signs can be made, and any value $m = 0, 1, 2, \ldots$ is permitted. (Throughout this chapter, we will take \mathbb{Z}^+ to be the set of all non-negative integers *including* zero.) Re-expressing this condition in terms of the γ parameters (5.55) appearing in the GKZ integral, and dropping the primes, this is

$$\gamma_0 \pm \gamma_1 \pm \gamma_2 \pm \gamma_3 = -2m. \tag{5.70}$$

We see immediately that the m = 0 singularities indeed correspond to the equations of the hyperplanes containing the eight facets of the regular octahedron on the left of figure 5.2, where the vertices in the figure correspond to $(\gamma_1, \gamma_2, \gamma_3) = \gamma_0(\pm 1, 0, 0), \gamma_0(0, \pm 1, 0)$ and $\gamma_0(0, 0, \pm 1)$. The remaining singularities for m > 0 then correspond to an infinite series of regularly spaced hyperplanes, both parallel, and exterior, to the facets of the octahedron.

Tropical analysis: an example

To appreciate the role of the Newton polytope, let us start with a simple example introduced in [149]. This is the GKZ integral

$$\mathcal{I}_{\gamma} = \int_{0}^{\infty} \mathrm{d}z_{1} \int_{0}^{\infty} \mathrm{d}z_{2} \, z_{1}^{\gamma_{1}-1} z_{2}^{\gamma_{2}-1} (x_{1} + x_{2}z_{2} + x_{3}z_{1}^{2} + x_{4}z_{1}z_{2}^{2})^{-\gamma_{0}}, \tag{5.71}$$

⁴The significance of this rescaling can be anticipated by noting that the Newton polytope of the GKZ denominator \mathcal{D}^{γ_0} , in the special cases where $\gamma_0 \in \mathbb{N}$ so that \mathcal{D}^{γ_0} is itself a polynomial when expanded out, is simply the Newton polytope of \mathcal{D} linearly rescaled by γ_0 .

⁵The argument in [45] involves expanding the integrand of the triple-K integral about its lower limit and looking for the appearance of z^{-1} poles.



Figure 5.3: Left: The integration sectors for the tropicalised GKZ integral (5.74), where each sector corresponds to the dominance of a different term in the denominator. The sector boundaries are simultaneously the normals to the facets of the Newton polytope shown on the right. Right: Combining the conditions on γ_1 and γ_2 for the convergence of each sector, we obtain the interior of the Newton polytope (rescaled by γ_0) as shaded.

whose \mathcal{A} -matrix is

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 \\ 0 & 0 & 2 & 1 \\ 0 & 1 & 0 & 2 \end{pmatrix}.$$
 (5.72)

The singularities of the integral derive from regions where the z_i (for i = 1, 2) either vanish or tend to infinity. Setting $z_i = e^{\tau_i}$, these regions are mapped to $|\tau_i| \to \infty$ and

$$\mathcal{I}_{\gamma} = \int_{-\infty}^{\infty} \mathrm{d}\tau_1 \int_{-\infty}^{\infty} \mathrm{d}\tau_2 \, e^{\gamma_1 \tau_1 + \gamma_2 \tau_2} (x_1 + x_2 e^{\tau_2} + x_3 e^{2\tau_1} + x_4 e^{\tau_1 + 2\tau_2})^{-\gamma_0}. \tag{5.73}$$

For large $|\tau_i|$, we can approximate this integral by its *tropicalisation* as discussed in [152],

$$\mathcal{I}_{\gamma}^{\text{trop.}} = x_j^{-\gamma_0} \int_{-\infty}^{\infty} \mathrm{d}\tau_1 \int_{-\infty}^{\infty} \mathrm{d}\tau_2 \exp\left[\gamma_1 \tau_1 + \gamma_2 \tau_2 - \gamma_0 \max(0, \tau_2, 2\tau_1, \tau_1 + 2\tau_2)\right], \quad (5.74)$$

which corresponds to retaining only the leading exponential in the GKZ denominator. Which term this is will depend on which sector of the (τ_1, τ_2) plane we are in. If the dominant term is, say, the *j*th one, then the overall prefactor is $x_j^{-\gamma_0}$ as shown. If all $x_k > 0$ for $k = 1, \ldots, 4$, the tropicalisation of the denominator in fact provides a lower bound and so, for real $\gamma_0 > 0$ and real γ_1 and γ_2 , we have $\mathcal{I}_{\gamma} < \mathcal{I}_{\gamma}^{\text{trop.}}$. The convergence of $\mathcal{I}_{\gamma}^{\text{trop.}}$ then establishes that of \mathcal{I}_{γ} . (For rigorous bounds allowing complex γ_i , see [149, 150].)

The various integration sectors, as illustrated in figure 5.3, are then as follows:

- (i) $\tau_1 < 0$ and $\tau_2 < 0$ so j = 1 and $\max(0, \tau_2, 2\tau_1, \tau_1 + 2\tau_2) = 0$.
- (ii) $\tau_1 + \tau_2 < 0$ and $\tau_2 > 0$ so j = 2 and $\max(0, \tau_2, 2\tau_1, \tau_1 + 2\tau_2) = \tau_2$.
- (iii) $\tau_1 > 0$ and $\tau_1 2\tau_2 > 0$ so j = 3 and $\max(0, \tau_2, 2\tau_1, \tau_1 + 2\tau_2) = 2\tau_1$.
- (iv) $\tau_1 + \tau_2 > 0$ and $\tau_1 2\tau_2 < 0$ so j = 4 and $\max(0, \tau_2, 2\tau_1, \tau_1 + 2\tau_2) = \tau_1 + 2\tau_2$.

Each sector forms a cone within which we can reparametrise $\boldsymbol{\tau} = (\tau_1, \tau_2)$ as

$$\boldsymbol{\tau} = \lambda_1 \boldsymbol{n}_1 + \lambda_2 \boldsymbol{n}_2, \qquad \lambda_1, \lambda_2 \ge 0 \tag{5.75}$$

where n_1 and n_2 are the outward-pointing vectors forming the boundary of that particular sector. By inspection, these are simultaneously the normal vectors to the facets of the Newton polytope shown in the right-hand panel of figure 5.3, where the two normals chosen are those for the two facets containing the leading vertex j. For the sector j = 3, for example, we have $n_1 = (0, -1)$ and $n_2 = (2, 1)$ and so $\tau_1 = 2\lambda_2$ and $\tau_2 = -\lambda_1 + \lambda_2$. This third sector of the tropicalised integral is then

$$\mathcal{I}_{\gamma}^{\text{trop.}}\Big|_{j=3} = 2x_3^{-\gamma_0} \int_0^\infty d\lambda_1 \int_0^\infty d\lambda_2 \exp\left[-\gamma_2 \lambda_1 + (2\gamma_1 + \gamma_2 - 4\gamma_0)\lambda_2\right].$$
(5.76)

The linearity of the tropicalised exponent means that the integrals over λ_1 and λ_2 factorise, and for convergence as $\lambda_i \to \infty$, both exponents must separately be negative:

$$\gamma_2 > 0, \qquad -2\gamma_1 - \gamma_2 + 4\gamma_0 > 0.$$
 (5.77)

This corresponds to the interior region bounded by the two lines intersecting the vertex $(\gamma_1, \gamma_2) = (2\gamma_0, 0)$ in the right-hand panel of figure 5.3. This vertex is precisely that corresponding to the dominant j = 3 term (namely, $x_3 z_1^2$) in the GKZ denominator, after rescaling by γ_0 . On the boundary of the convergence region, the integral has either a single or a double pole according to how many of the inequalities in (5.77) are saturated.

Repeating this exercise for the remaining sectors, we obtain the additional constraints

$$\gamma_1 > 0, \qquad \gamma_1 - \gamma_2 + \gamma_0 > 0.$$
 (5.78)

Combining all these conditions, the full integral $\mathcal{I}_{\gamma}^{\text{trop.}}$ then converges for (γ_1, γ_2) within the polytope shown in the figure. This is indeed the Newton polytope for the GKZ denominator after rescaling all vertex vectors by γ_0 .

Tropical analysis: general case

The analysis above clearly generalises. Setting again $z_i = e^{\tau_i}$, the general GKZ integral (5.1) has the tropical approximation

$$\mathcal{I}_{\gamma}^{\text{trop.}} = \int_{\mathbb{R}^N} \mathrm{d}\boldsymbol{\tau} \, \exp\Big[\sum_{i=1}^N \gamma_i \tau_i - \gamma_0 \max_k \Big(\ln x_k + \sum_{i=1}^N a_{ik} \tau_i\Big)\Big]. \tag{5.79}$$

In particular, this is a good approximation precisely for the large $|\tau_i|$ regions where any singularities of the GKZ integral must arise, and so convergence of the tropical approximation implies convergence of the full GKZ integral.⁶

The different integration sectors of the tropical integral (5.79) correspond to when different terms dominate and are selected as the maximum in the exponent. For sufficiently

⁶ For real γ_i , $\gamma_0 > 0$ and $x_j > 0$, the tropical approximation provides an upper bound on the GKZ integral as noted in the previous example. Cases where the γ_i can be complex and the x_j are not constrained to be positive can be handled by establishing a rigorous bound on the GKZ denominator, see [149, 150].

large $|\tau_i|$, this depends only on the *direction* in the $\tau = (\tau_1, \ldots, \tau_N)$ plane and we can neglect any contribution from the $\ln x_k$ terms. Let us consider then the sector where, say, the *j*th term forms the maximum. This sector can be parametrised as

$$\boldsymbol{\tau} = \sum_{J \in \Phi_j} \lambda^{(J)} \boldsymbol{n}^{(J)}, \qquad \lambda^{(J)} \ge 0, \tag{5.80}$$

where Φ_j denotes the set of facets containing the vertex j, the $\lambda^{(J)}$ are the new integration variables, and

$$\boldsymbol{n}^{(J)} = (n_1^{(J)}, \dots, n_N^{(J)})^T$$
 (5.81)

is the outward-pointing normal to the facet J. We will assume that Φ_j contains precisely N facets so that (5.80) holds.⁷ The contribution of this sector is then

$$\mathcal{I}_{\gamma}^{\text{trop.}}\Big|_{j} = x_{j}^{-\gamma_{0}} \prod_{J \in \Phi_{j}} \int_{0}^{\infty} \mathrm{d}\lambda^{(J)} \exp\left[\lambda^{(J)} \sum_{i=1}^{N} n_{i}^{(J)} (\gamma_{i} - \gamma_{0} a_{ij})\right].$$
(5.82)

As in the previous example, convergence then requires each of these exponents to be negative giving

$$\sum_{i=1}^{N} n_i^{(J)}(\gamma_i - \gamma_0 a_{ij}) < 0 \qquad \forall J \in \Phi_j.$$
(5.83)

Viewed geometrically, these conditions state that the parameter vector $\hat{\gamma}$ lies to the *inside* of the (N-1)-dimensional hyperplane containing facet J of the rescaled Newton polytope,

$$\boldsymbol{n}^{(J)} \cdot (\hat{\boldsymbol{\gamma}} - \gamma_0 \boldsymbol{a}_j) < 0, \tag{5.84}$$

and that this holds for all facets J containing the *j*th vertex vector $\gamma_0 a_j$. Convergence of the *full* tropicalised GKZ integral requires convergence in every integration sector, and hence for every vertex j of the rescaled Newton polytope. The condition (5.84) must thus hold for *all* facets J, meaning $\hat{\gamma}$ must lie completely inside the rescaled Newton polytope.

5.3.3 Meromorphic continuation

Having shown the convergence of GKZ integrals for $\hat{\gamma}$ lying within the rescaled Newton polytope, the existence of further infinite sets of singular hyperplanes parallel to each facet can be established by meromorphic continuation [149, 150]. Once again, the idea is most easily seen in the context of an example, after which we resume our general analysis.

Example

Returning the GKZ integral (5.71), let us construct a continuation across, say, the upperright facet of the Newton polytope shown on the right of figure 5.3. The relevant outward normal is $\mathbf{n} = (2, 1)$. Following [149], we perform a rescaling $z_i \to \lambda^{-n_i} z_i$, namely $z_1 \to$

 $^{^{7}}$ If there are fewer than this, we can factor out a finite integral over a transverse subspace following appendix A of [152] then apply the argument above for the remaining integral over a lower-dimensional cone.

 $\lambda^{-2}z_1$ and $z_2 \to \lambda^{-1}z_2$, where λ is some fixed parameter. The integral (5.71) becomes

$$\mathcal{I}_{\gamma} = \lambda^{-2\gamma_1 - \gamma_2 + 4\gamma_0} \int_0^\infty \mathrm{d}z_1 \int_0^\infty \mathrm{d}z_2 \, z_1^{\gamma_1 - 1} z_2^{\gamma_2 - 1} (x_1 \lambda^4 + x_2 z_2 \lambda^3 + x_3 z_1^2 + x_4 z_1 z_2^2)^{-\gamma_0}$$
(5.85)

but its value remains unchanged. We can therefore differentiate to find

$$0 = \frac{\mathrm{d}}{\mathrm{d}\lambda} \mathcal{I}_{\gamma} \Big|_{\lambda=1} = (-2\gamma_1 - \gamma_2 + 4\gamma_0) \mathcal{I}_{\gamma} - 4\gamma_0 x_1 \mathcal{I}_{\gamma'} - 3\gamma_0 x_2 \mathcal{I}_{\gamma''}$$
(5.86)

where

$$\gamma'_0 = \gamma_0 + 1, \qquad \gamma'_1 = \gamma_1, \qquad \gamma'_2 = \gamma_2
\gamma''_0 = \gamma_0 + 1, \qquad \gamma''_1 = \gamma_1, \qquad \gamma''_2 = \gamma_2 + 1.$$
(5.87)

Alternatively, (5.86) can be obtained by taking a linear combination of the Euler equations and DWI for (5.71), namely

$$0 = \left(-2(\gamma_1 + 2\theta_3 + \theta_4) - (\gamma_2 + \theta_2 + 2\theta_4) + 4(\gamma_0 + \sum_{j=1}^4 \theta_j)\right) \mathcal{I}_{\gamma}$$

= $\left(4\gamma_0 - 2\gamma_1 - \gamma_2 + 4\theta_1 + 3\theta_3\right) \mathcal{I}_{\gamma},$ (5.88)

where evaluating the action of the $\theta_i = x_i \partial_{x_i}$ yields (5.86).

As both $\mathcal{I}_{\gamma'}$ and $\mathcal{I}_{\gamma''}$ in (5.86) take the same form as the original integral \mathcal{I}_{γ} , except with shifted parameters, the convergence regions are given by (5.77) and (5.78) replacing γ with γ' or γ'' . In terms of the unshifted parameters, $\mathcal{I}_{\gamma'}$ thus converges for

$$\gamma_1 > 0, \qquad \gamma_2 > 0, \qquad \gamma_0 + \gamma_1 - \gamma_2 + 1 > 0, \qquad 4\gamma_0 - 2\gamma_1 - \gamma_2 + 4 > 0,$$
 (5.89)

while $\mathcal{I}_{\gamma''}$ converges for

$$\gamma_1 > 0, \qquad \gamma_2 + 1 > 0, \qquad \gamma_0 + \gamma_1 - \gamma_2 > 0, \qquad 4\gamma_0 - 2\gamma_1 - \gamma_2 + 3 > 0.$$
 (5.90)

In each case, the size of the Newton polytope is rescaled from $\gamma_0 \to \gamma_0 + 1$, while for $\mathcal{I}_{\gamma''}$ we also translate by the vector (0, -1) as shown in figure 5.4. Note that neither of these operations change the normals to the facets. Re-arranging (5.86), we now have

$$\mathcal{I}_{\boldsymbol{\gamma}} = (4\gamma_0 - 2\gamma_1 - \gamma_2)^{-1} \big(4\gamma_0 x_1 \mathcal{I}_{\boldsymbol{\gamma}'} + 3\gamma_0 x_2 \mathcal{I}_{\boldsymbol{\gamma}''} \big), \tag{5.91}$$

where the sum of shifted integrals on the right-hand side converges only for the *intersection* of the two shifted polytopes (5.89) and (5.90), namely

$$\gamma_1 > 0, \qquad \gamma_2 > 0, \qquad \gamma_0 + \gamma_1 - \gamma_2 > 0, \qquad 4\gamma_0 - 2\gamma_1 - \gamma_2 + 3 > 0.$$
 (5.92)

Comparing with the original polytope formed by (5.77) and (5.78), only the final inequality has changed. Now, the region of convergence extends across the facet with normal (2, 1) as shown in figure 5.4. Equation (5.91) thus gives a meromorphic continuation of \mathcal{I}_{γ} around the pole at $4\gamma_0 - 2\gamma_1 - \gamma_2 = 0$ (corresponding to the facet of the original Newton polytope



Figure 5.4: Left: Convergence regions of \mathcal{I}_{γ} (green), $\mathcal{I}_{\gamma'}$ (blue) and $\mathcal{I}_{\gamma''}$ (orange). The meromorphic continuation (5.91) holds for the intersection of the convergence regions for $\mathcal{I}_{\gamma'}$ and $\mathcal{I}_{\gamma''}$. This extends the domain of convergence of \mathcal{I}_{γ} across the upper-right facet with normal (2, 1) to form the region bounded by the solid line. *Right:* Dashed lines indicate the complete set of singular hyperplanes (5.93) of the GKZ integral (5.71).

normal to (2,1)) to the larger region (5.92).

This process can then be repeated for the boundary of the new region (5.92) by applying the same procedure (namely, rescaling $z_i \to \lambda^{-n_i} z_i$, differentiating with respect to λ then setting $\lambda = 1$) to the integrals on the right-hand side of (5.91). Alternatively, we can extend (5.91) iteratively by using shifted analogues of (5.91) to replace $\mathcal{I}_{\gamma'}$ and $\mathcal{I}_{\gamma''}$ on the right-hand side of (5.91) itself. Repeating such calculations for all the facet normals of the original Newton polytope, we obtain an infinite set of singular hypersurfaces parallel to the facets of the Newton polytope. The integral (5.71) is thus singular on the hyperplanes

$$\gamma_1 = -m_1, \qquad \gamma_2 = -m_2, \qquad \gamma_0 + \gamma_1 - \gamma_2 = -m_3, \qquad 4\gamma_0 - 2\gamma_1 - \gamma_2 = -3m_4$$
 (5.93)

for any (independent) choice of non-negative integers $m_i \in \mathbb{Z}^+$, as illustrated in the righthand panel of figure 5.4.

General analysis

The analysis in this last example readily extends to general GKZ integrals. We begin by defining a few useful quantities. First, we have the (N + 1)-dimensional vectors

$$\boldsymbol{\gamma} = \begin{pmatrix} \gamma_0 \\ \hat{\boldsymbol{\gamma}} \end{pmatrix}, \qquad \boldsymbol{\mathcal{A}}_j = \begin{pmatrix} 1, \\ \boldsymbol{a}_j \end{pmatrix}, \qquad \boldsymbol{N}^{(J)} = \begin{pmatrix} n_0^{(J)} \\ \boldsymbol{n}^{(J)} \end{pmatrix},$$
(5.94)

where γ is the usual GKZ parameter vector, \mathcal{A}_j is the *j*th column of the full \mathcal{A} -matrix (including the top row of 1s), and, as above, $\mathbf{n}^{(J)}$ is the *N*-dimensional outwards-pointing normal to facet *J* of the Newton polytope. The additional component $n_0^{(J)}$ is fixed by requiring that

$$0 = \mathbf{N}^{(J)} \cdot \mathbf{A}_j, \qquad j \in \varphi_J \tag{5.95}$$

where φ_J denotes the set of vertices lying within the facet J, giving

$$n_0^{(J)} = -\boldsymbol{n}^{(J)} \cdot \boldsymbol{a}_j, \qquad j \in \varphi_J.$$
(5.96)

The condition that $\hat{\gamma}$ lies in the hyperplane containing facet J of the rescaled Newton polytope,

$$0 = \boldsymbol{n}^{(J)} \cdot (\hat{\boldsymbol{\gamma}} - \gamma_0 \boldsymbol{a}_j) \tag{5.97}$$

can now be compactly re-expressed as

$$0 = \boldsymbol{\gamma} \cdot \boldsymbol{N}^{(J)} \tag{5.98}$$

and the domain of convergence (5.84) corresponds to $\gamma \cdot N^{(J)} < 0$ for all facets J. (From an (N+1)-dimensional perspective, the Newton polytope therefore corresponds to a cone.) In addition, we define the distance function

$$d_i^{(J)} = -\mathcal{A}_i \cdot \mathbf{N}^{(J)} = \mathbf{n}^{(J)} \cdot (\mathbf{a}_j - \mathbf{a}_i), \qquad j \in \varphi_J.$$
(5.99)

If $\mathbf{n}^{(J)}$ is a unit vector, $d_i^{(J)}$ is the normal distance from vertex *i* to facet *J* of the Newton polytope. Rather than choosing $\mathbf{n}^{(J)}$ to be a unit vector, however, it will be more convenient in practice to choose $\mathbf{n}^{(J)}$ (and hence $\mathbf{N}^{(J)}$) to have integer components.

We now proceed to construct a meromorphic continuation of the GKZ integral across a chosen facet J of the rescaled Newton polytope. To this end, we form a linear combination of $n_0^{(J)}$ times the DWI plus the sum of $n_k^{(J)}$ times the *k*th Euler equation, namely

$$0 = \left[n_0^{(J)} \left(\gamma_0 + \sum_{l=1}^n \theta_l \right) + \sum_{k=1}^N n_k^{(J)} \left(\gamma_k + \sum_{l=1}^n a_{kl} \theta_l \right) \right] \mathcal{I}_{\gamma}$$
$$= \left(\gamma \cdot \mathbf{N}^{(J)} - \sum_{l=1}^n d_l^{(J)} \theta_l \right) \mathcal{I}_{\gamma}.$$
(5.100)

The sum over l on the second line here can be restricted to values $l \notin \varphi_J$, corresponding to vertices l not in the facet J, since $d_l^{(J)}$ vanishes for all $l \in \varphi_J$. Moreover, by direct differentiation of the GKZ integral as we will discuss further in section 5.4.1, one can show that

$$\theta_l \mathcal{I}_{\gamma} = -\gamma_0 x_l \mathcal{I}_{\gamma + \mathcal{A}_l}. \tag{5.101}$$

Here, the parameter vector of the right-hand integral has been shifted from $\gamma \to \gamma + \mathcal{A}_l$. Rearranging, this immediately gives the desired meromorphic continuation:⁸

$$\mathcal{I}_{\gamma} = -\frac{\gamma_0}{\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(J)}} \Big(\sum_{l \notin \varphi_J} x_l \, d_l^{(J)} \, \mathcal{I}_{\boldsymbol{\gamma} + \boldsymbol{\mathcal{A}}_l} \Big).$$
(5.102)

The denominator $\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(J)}$ generates a pole at the hyperplane containing the facet J, while the sum of shifted integrals has a larger domain of convergence extending across the facet

⁸Alternatively, this equation can be derived by rescaling all $z_i \to \lambda^{-n_i^{(J)}} z_i$ in \mathcal{I}_{γ} and extracting a prefactor of $\lambda^{-\gamma \cdot N^{(J)}}$. We then differentiate with respect to λ and set $\lambda = 1$ analogously to in (5.86).

J of the original rescaled Newton polytope for \mathcal{I}_{γ} .

To see this, for each shifted integral labelled by an $l \notin \varphi_J$ in the sum (5.102), the domain of convergence (5.84) is

$$(\boldsymbol{\gamma} + \boldsymbol{\mathcal{A}}_l) \cdot \boldsymbol{N}^{(K)} = \boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} - d_l^{(K)} < 0 \qquad \forall K.$$
(5.103)

This is equivalent to

$$\boldsymbol{n}^{(K)} \cdot \left((\hat{\boldsymbol{\gamma}} + \boldsymbol{a}_l) - (\gamma_0 + 1) \boldsymbol{a}_k \right) < 0, \qquad k \in \varphi_K, \tag{5.104}$$

i.e., for every facet K, the shifted parameter vector $\hat{\gamma}' = \hat{\gamma} + a_l$ must lie inside the Newton polytope rescaled by $\gamma'_0 = \gamma_0 + 1$. The common overlap of these domains for every $l \notin \varphi_J$ then corresponds to

$$\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} < \delta^{(K)} \qquad \forall K, \tag{5.105}$$

where

$$\delta^{(K)} = \min_{l \notin \varphi_J} \left[d_l^{(K)} \right] \ge 0. \tag{5.106}$$

For any facet $K \neq J$, the set of vertices $l \notin \varphi_J$ includes vertices $l \in \varphi_K$ lying *in* the facet K. For such vertices, $d_l^{(K)}$ and hence $\delta^{(K)}$ is then zero. Just as in our earlier example, the domain of convergence for the sum of shifted integrals in (5.102) is therefore unchanged for all facets $K \neq J$,

$$\delta^{(K)} = 0 \quad \forall \ K \neq J. \tag{5.107}$$

The only facet across which the domain of convergence is extended is the facet K = J, for which we obtain an extension

$$\delta^{(J)} = \min_{l \notin \varphi_J} \left[d_l^{(J)} \right]. \tag{5.108}$$

Geometrically, $\delta^{(J)} > 0$ is the normal distance to the facet J of the (non-rescaled) Newton polytope starting from the *nearest* vertex l not belonging to J, multiplied by $|\mathbf{n}^{(J)}|$. If we choose $\mathbf{n}^{(J)}$ to have integer components, then as the components of the \mathcal{A} -matrix are also integer, $\delta^{(J)}$ will be a positive integer.

Equation (5.105), together with (5.107) and (5.108), thus give us the domain of convergence of the meromorphic continuation (5.102). Repeating the argument to construct further meromorphic continuations, one finds that the GKZ integral \mathcal{I}_{γ} has an infinite series of singular hyperplanes lying parallel to each facet J of the original Newton polytope. These hyperplanes are given by

$$\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(J)} = m_J \,\delta^{(J)}, \qquad m_J \in \mathbb{Z}^+, \tag{5.109}$$

where m_J is any non-negative integer $m = 0, 1, 2, \ldots$

Example: Let us check (5.108) against our previous example. Taking J to be the facet with outward normal $\mathbf{n}^{(J)} = (2, 1)$, we have $\varphi_J = \{3, 4\}$ and so using the \mathcal{A} -matrix (5.72),

$$\delta^{(J)} = \min_{l \in \{1,2\}} \sum_{i=1}^{2} n_i^{(J)} (a_{i3} - a_{il}) = \min(4,3) = 3.$$
 (5.110)

The sole shifted boundary

$$\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(J)} = \gamma_0 n_0^{(J)} + \hat{\boldsymbol{\gamma}} \cdot \boldsymbol{n}^{(J)} < \delta^{(J)}, \qquad (5.111)$$

where $n_0^{(J)} = -\sum_{i=1}^2 n_i^{(J)} a_{i3} = -4$, then evaluates to

$$-4\gamma_0 + 2\gamma_1 + \gamma_2 - 3 < 0 \tag{5.112}$$

in agreement with (5.92), and the singular hyperplanes in (5.109) match those in (5.93).

Implementation

In higher-dimensional examples, a convenient way to determine the singular hyperplanes (5.109) is to apply a convex hulling algorithm (see, *e.g.*, [156]) to identify which sets of vertex vectors \mathbf{a}_j form the facets of the Newton polytope. We will discuss this explicitly in section 5.6.3. The condition (5.98) that $\hat{\gamma}$ lies in the hyperplane containing facet J of the rescaled Newton polytope is then equivalent to

$$0 = \boldsymbol{\gamma} \cdot \boldsymbol{N}^{(J)} = \det\left(\boldsymbol{\gamma} \mid \boldsymbol{\mathcal{A}}_{j_1} \mid \dots \mid \boldsymbol{\mathcal{A}}_{j_N}\right), \tag{5.113}$$

where $j_1, \ldots, j_N \in \varphi_J$ are the N vertices belonging to facet J, and the \mathcal{A}_j are the corresponding \mathcal{A} -matrix columns. To see this, note that from (5.95) we have $\mathcal{A}_j \cdot \mathbf{N}^{(J)} = 0$ for all the N vectors $j \in \varphi_J$. As the total dimension of the vector space is N + 1, the condition $\boldsymbol{\gamma} \cdot \mathbf{N}^{(J)} = 0$ implies that $\boldsymbol{\gamma}$ lies in the span of the \mathcal{A}_j with $j \in \varphi_J$, and hence the determinant above vanishes. The components $n_i^{(J)}$ of $\mathbf{N}^{(J)}$, for $i = 0, \ldots, N$, can thus be identified by expanding out the determinant and extracting the coefficient of γ_i . This tells us that $n_i^{(J)}$ is given by the (i, 1)th cofactor of the matrix, for example

$$n_0^{(J)} = \det(\mathbf{a}_{j_1} | \dots | \mathbf{a}_{j_N}).$$
(5.114)

One must however also check that $\mathbf{n}^{(J)}$ corresponds to the outwards-pointing normal by verifying that $d_k^{(J)} = -\mathbf{A}_k \cdot \mathbf{N}^{(J)} > 0$ for some $k \notin \varphi_J$, and swapping two columns of (5.113) if not. The spacing $\delta^{(J)}$ of the singular hyperplanes can then be computed using (5.108) and (5.99).

5.4 Shift operators

Let us now examine the shift operators associated with \mathcal{A} -hypergeometric functions. Two natural classes present themselves: the 'annihilation' operators which correspond to the simple derivative $\partial_j = \partial/\partial x_j$, and the 'creation' operators which are purely polynomial differential operators (*i.e.*, operators in the Weyl algebra) that invert this operation.

5.4.1 Annihilation operators

From the GKZ integral (5.1) and denominator (5.2), we see by direct differentiation that

$$\partial_j \mathcal{I}_{\gamma} = -\gamma_0 \mathcal{I}_{\gamma'}, \qquad j = 1, \dots, n \tag{5.115}$$

where

$$\gamma'_0 = \gamma_0 + 1, \qquad \gamma'_i = \gamma_i + a_{ij}, \qquad i = 1, \dots, N.$$
 (5.116)

In other words, differentiating with respect to x_j increases the power of the denominator by one, and adds to the numerator all powers of z_i multiplying x_j in the denominator. From the \mathcal{A} -matrix perspective, the shift of the parameter vector $\boldsymbol{\gamma}$ is given by the *j*th column of the full \mathcal{A} -matrix,

$$\boldsymbol{\gamma}' = \boldsymbol{\gamma} + \boldsymbol{\mathcal{A}}_j, \tag{5.117}$$

combining the two formulae in (5.116).

One can naturally think of the toric equations (5.30) as representing the difference of two products of annihilation operators, such that the total shift generated by each product is the same leading to a cancellation. Namely, each factor

$$\prod_{j=1}^{n} \partial_j^{u_j^{\pm}} \tag{5.118}$$

produces an overall parameter shift

$$\gamma \to \gamma + \sum_{j=1}^{n} \mathcal{A}_{j} u_{j}^{\pm},$$
 (5.119)

but since

$$\sum_{j=1}^{n} \mathcal{A}_{j} u_{j}^{+} = \sum_{j=1}^{n} \mathcal{A}_{j} u_{j}^{-}$$
(5.120)

the final shifted integral is the same in both cases and the difference vanishes.

Notice also that knowledge of the full set of n annihilation operators, plus the parameter shifts they produce, is equivalent to knowledge of all columns of the \mathcal{A} -matrix and hence the full GKZ integral itself.⁹

Example: The annihilation operators for the GKZ uplift (5.53) of the triple-K integral (5.50) are ∂_j for j = 1, ..., 6. The triple-K integral itself corresponds to evaluating the GKZ integral on the physical hypersurface $\boldsymbol{x} = (p_1^2, p_2^2, p_3^2, 1, 1, 1)$ according to (5.55). The first three annihilators thus become

$$\partial_j = \frac{\partial}{\partial x_j} = \frac{\partial}{\partial p_j^2} = \frac{1}{p_j} \frac{\partial}{\partial p_j}, \qquad j = 1, 2, 3.$$
(5.121)

while for the remaining three we need to use the Euler equations following from the \mathcal{A} -matrix (5.56). These are

$$0 = \beta_1 - \theta_1 + \theta_4, \qquad 0 = \beta_2 - \theta_2 + \theta_5, \qquad 0 = \beta_3 - \theta_3 + \theta_6, \tag{5.122}$$

 $^{^{9}}$ Prior to the work of GKZ, this approach was pioneered by Miller *et al* [157, 158] for various Lauricella and Horn-type hypergeometric functions for which the annihilators can be identified from the series definition.

and projecting to the physical hypersurface by setting $x_4 = x_5 = x_6 = 1$ gives

$$\partial_4 = \theta_1 - \beta_1 = \frac{p_1}{2} \frac{\partial}{\partial p_1} - \beta_1, \quad \partial_5 = \theta_2 - \beta_2 = \frac{p_2}{2} \frac{\partial}{\partial p_2} - \beta_2, \quad \partial_6 = \theta_3 - \beta_3 = \frac{p_3}{2} \frac{\partial}{\partial p_3} - \beta_3. \tag{5.123}$$

Up to trivial numerical factors, these are the shift operators

$$\mathcal{L}_j = -\frac{1}{p_j} \frac{\partial}{\partial p_j}, \qquad \mathcal{R}_j = 2\beta_j - p_j \frac{\partial}{\partial p_j}, \qquad j = 1, 2, 3, \tag{5.124}$$

introduced in [42, 85]. The action of these operators on the triple-K integral (5.50) can be obtained from their action on the individual Bessel functions in the integrand giving

$$\mathcal{L}_{1}I_{\alpha,\{\beta_{1},\beta_{2},\beta_{3}\}} = -(\alpha+1)I_{\alpha+1,\{\beta_{1}-1,\beta_{2},\beta_{3}\}}, \qquad \mathcal{R}_{1}I_{\alpha,\{\beta_{1},\beta_{2},\beta_{3}\}} = -(\alpha+1)I_{\alpha+1,\{\beta_{1}+1,\beta_{2},\beta_{3}\}}, \tag{5.125}$$

with the others following by permutation. This is consistent with the expected action for the annihilation operators: from the columns of the \mathcal{A} -matrix (5.56), this is

$$\mathcal{L}_j: \quad \gamma'_0 \to \gamma'_0 + 1, \qquad \gamma'_j \to \gamma'_j - 1, \qquad \mathcal{R}_j: \quad \gamma'_0 \to \gamma'_0 + 1, \qquad \gamma'_j \to \gamma'_j + 1, \quad (5.126)$$

which from (5.55) is

$$\mathcal{L}_j: \quad \alpha \to \alpha + 1, \qquad \beta_j \to \beta_j - 1, \qquad \mathcal{R}_j: \quad \alpha \to \alpha + 1, \qquad \beta_j \to \beta_j + 1.$$
 (5.127)

5.4.2 Creation operators

Over the next three subsections, we present a construction of creation operators motivated by consideration of the spectral singularities. These ideas are then illustrated using the Gauss hypergeometric function. Originally, creation operators were first proposed by Saito in [146, 147]; for further discussion, see [148, 159].

By definition, when acting on a GKZ integral, the creation operator C_j produces the *inverse* parameter shift to the annihilation operator $\partial_j = \partial/\partial x_j$. If we act with one operator followed by the other, therefore, we must arrive back at the original integral up to some function of the parameters:

$$\mathcal{C}_j \partial_j \mathcal{I}_{\gamma} = b_j(\gamma) \mathcal{I}_{\gamma}. \tag{5.128}$$

As we will see shortly, this 'b-function' $b_j(\gamma)$ is a polynomial whose zeros correspond to a specific subset of the singular hyperplanes of \mathcal{I}_{γ} given in (5.109). First, however, let us sketch how knowing $b_j(\gamma)$ enables a direct construction of the creation operator \mathcal{C}_j .

The first step is to replace all the parameters γ appearing in the *b*-function with linear combinations of Euler operators using the DWI and Euler equations (5.20). This defines a new polynomial $B_j(\theta)$ in the Euler operators,

$$B_j(\theta) = b_j(\gamma) \Big|_{\gamma \to -\sum_{k=1}^n \mathcal{A}_k \theta_k}$$
(5.129)

such that

$$\mathcal{C}_j \partial_j \mathcal{I}_{\gamma} = B_j(\theta) \mathcal{I}_{\gamma}. \tag{5.130}$$

As all Euler operators commute with one another, there are no ordering ambiguities here.

Next, we expand out $B_j(\theta)$ and re-arrange so that, in every term, all factors of x_k are to the left of all derivatives ∂_k . Up to a constant coefficient, each term of $B_j(\theta)$ is thus of the form

$$\prod_{k=1}^{n} x_k^{\mathbf{b}_k} \partial_k^{\mathbf{b}_k} \tag{5.131}$$

for some set of powers \mathfrak{b}_k . In certain cases, the product $\prod_k \partial_k^{\mathfrak{b}_k}$ will already contain an explicit factor of ∂_j . Otherwise, we can use the toric equations (5.30) to replace the product $\prod_k \partial_k^{\mathfrak{b}_k}$ (which acts on the GKZ integral \mathcal{I}_{γ} as per (5.130)) with an equivalent product that *does* contain an explicit factor of ∂_j . Such a replacement will always be possible provided the *b*-function is correctly chosen. After completing this operation for every term, the right-hand side of (5.130) now matches the form of the left-hand side allowing the operator \mathcal{C}_j to be read off. Thus, with the aid of the toric equations, $B_j(\theta)$ acting on \mathcal{I}_{γ} can be explicitly factorised into the form $\mathcal{C}_j \partial_j$.

As a final step, the creation operator C_j , which is a differential operator with polynomial coefficients defined in the *n*-dimensional GKZ space, must be projected back to the physical hypersurface. For this, we restore all x_k to their physical values (noting the x_k are positioned to the left of all derivatives), and use the Euler equations evaluated on the physical hypersurface to replace derivatives in directions lying off the physical hypersurface with derivatives tangential to this hypersurface. This replacement also restores a dependence on the parameters γ . Many examples of this projection procedure will appear in subsequent sections.

5.4.3 Action of the creation operator

Returning to (5.128) and using the action of the annihilator ∂_j as given in (5.115), the action of the creation operator is

$$\mathcal{C}_j \mathcal{I}_{\gamma'} = -\gamma_0^{-1} b_j(\gamma) \mathcal{I}_{\gamma}.$$
(5.132)

As the shift here is acting in the direction $\gamma' \rightarrow \gamma = \gamma' - \mathcal{A}_j$, rearranging (5.116) we have

$$\gamma_0 = \gamma'_0 - 1, \qquad \gamma_i = \gamma'_i - a_{ij}, \qquad i = 1, \dots N.$$
 (5.133)

We will retain this allocation of prime and unprimed variables in the following for compatibility with the algorithm in the previous section based on (5.128).

Before discussing the *b*-function itself, a crucial point to notice is that the parameter shift (5.133) can potentially take us from a *finite* to a *divergent* GKZ integral. In contrast, the reverse shift (5.116) associated with the annihilation operator ∂_j , when acting on a finite integral, will always produce another finite integral.

To see this, let us start with an integral $\mathcal{I}_{\gamma'}$ for which the vector $\hat{\gamma}' = (\gamma'_1, \ldots, \gamma'_N)$ lies inside the rescaled Newton polytope with vertices $\gamma'_0 a_j$. In the notation of section 5.3.3, this means that for every facet K we have

$$\boldsymbol{\gamma}' \cdot \boldsymbol{N}^{(K)} < 0 \tag{5.134}$$

and the GKZ representation for $\mathcal{I}_{\gamma'}$ converges without meromorphic continuation. For the

shifted integral \mathcal{I}_{γ} in (5.132), we then have

$$\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} = \boldsymbol{\gamma}' \cdot \boldsymbol{N}^{(K)} + d_j^{(K)}$$
(5.135)

where

$$d_j^{(K)} = \boldsymbol{n}^{(K)} \cdot (\boldsymbol{a}_k - \boldsymbol{a}_j), \qquad k \in \varphi_K$$
(5.136)

is proportional to the normal distance from vertex j to facet K of the (non-rescaled) Newton polytope. Now, for any facet K containing the vertex j, $d_j^{(K)}$ vanishes and hence $\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} < 0$. For the remaining facets not containing the vertex j, however, $d_j^{(K)} > 0$ since $\boldsymbol{n}^{(K)}$ is the outward normal and j lies to the inside of the facet. Consequently, we cannot be sure that $\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} < 0$ for all facets K, and hence that $\mathcal{I}_{\boldsymbol{\gamma}}$ is finite. Rather, if there are any facets for which $\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} \geq 0$, the shifted integral $\mathcal{I}_{\boldsymbol{\gamma}}$ will diverge whenever the singularity condition (5.109),

$$\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} = m_K \delta^{(K)}, \qquad m_K \in \mathbb{Z}^+ \tag{5.137}$$

is satisfied for some non-negative integer m_K . Combined with (5.135), this condition allows us to identify the initial parameter values γ' for which the shifted integral \mathcal{I}_{γ} diverges.

For the annihilation operator ∂_j , the direction of the parameter shifts are reversed and so if the starting integral is finite, the shifted integral is also necessarily finite.

5.4.4 Finding the b-function

An apparent puzzle now arises for cases where the shifted integral \mathcal{I}_{γ} in (5.132) is divergent, since the action of a differential operator \mathcal{C}_j with polynomial coefficients on any finite integral $\mathcal{I}_{\gamma'}$ must clearly be finite. The resolution is that, for such cases, the *b*-function in (5.132) must have a *zero* cancelling the divergence in \mathcal{I}_{γ} such that the right-hand side is finite.¹⁰

The *b*-function for the creation operator C_j must thus have zeros corresponding to every singular hyperplane that can be reached by a single application of C_j to any finite starting integral, as illustrated in figure 5.5. The minimal *b*-function, containing just these factors alone, is

$$b_j(\boldsymbol{\gamma}) = \prod_{K \notin \Phi_j} \prod_{m_K=0}^{F_j^{(K)} - 1} (\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} - m_K \delta^{(K)})$$
(5.138)

where the first product runs over all facets K not containing the vertex j and the upper limit in the second product is set by

$$F_j^{(K)} = \frac{d_j^{(K)}}{\delta^{(K)}}.$$
(5.139)

This counts by how many steps (in units of $\delta^{(K)}$, the spacing between singular hyperplanes) the creation operator C_j raises $\gamma' \cdot N^{(K)}$ according to (5.135). Effectively, if we *define* an

¹⁰In a 'dimensional' regularisation scheme where all parameters are shifted infinitesimally $\gamma \to \gamma + \varepsilon \bar{\gamma}$, this requires $b_j(\gamma) \sim \varepsilon^k$ while $\mathcal{I}_{\gamma} \sim \varepsilon^{-k}$ for some $k \in \mathbb{Z}^+$ such that $b_j(\gamma) \mathcal{I}_{\gamma}$ is finite as $\varepsilon \to 0$.



Figure 5.5: Mapping of finite to divergent integrals under the action of the creation operator C_j as per (5.135), and construction of the corresponding *b*-functions. Left: If $d_j^{(K)} = \delta^{(K)}$, facet K contributes only the factor $\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)}$ to the *b*-function. The zero of this factor cancels the pole of the only singular integral (dashed line) that can be reached starting from a finite integral. Right: If $d_j^{(K)} = 3\delta^{(K)}$, the facet contributes three factors, $\prod_{m_K=0}^2 (\boldsymbol{\gamma} \cdot \boldsymbol{N}^{(K)} - m_K \delta^{(K)})$ whose zeros cancel the poles of the three singular integrals reachable from a finite starting integral. The shaded region indicates the rescaled Newton polytope.

initial m'_K by the relation $\gamma' \cdot \mathbf{N}^{(K)} = m'_K \delta^{(K)}$, the creation operator \mathcal{C}_j acts to raise this to $m_K = m'_K + F_j^{(K)}$. Thus, if $F_j^{(K)} = 1$ for some particular facet K, only the singularity in (5.137) with $m_K = 0$ can be reached by the action of \mathcal{C}_j on a finite starting integral (namely, that with $m'_K = -1$). The product over m_K in (5.138) is thus capped at $F_j^{(K)} - 1 = 0$. Alternatively, if $F_j^{(K)} = 2$ for some facet, both the $m_K = 0$ and $m_K = 1$ singularities can be reached by acting with \mathcal{C}_j on the finite starting integrals with $m'_K = -2$ and $m'_K = -1$ respectively. The product over m_K in (5.138) then runs up to $F_j^{(K)} - 1 = 1$, and so on.

For all the Feynman and Witten diagrams we analyse in the remainder of the chapter, $F_j^{(K)}$ is an integer for all K and the minimal b-function (5.138) (containing only the zeros necessary to cancel out the singularities of \mathcal{I}_{γ}) is sufficient to find all creation operators. These operators are moreover of the lowest possible order in derivatives, since the b-function has the fewest factors. Nevertheless, in certain exceptional cases, the factorisation step of the algorithm in section 5.4.2 can fail when using the minimal b-function. Such cases, which arise when the associated toric ideal is non-normal [146–148, 159], can be handled by supplementing (5.138) with additional factors. An example, which also features a non-integer $F_i^{(K)}$, is discussed in appendix C.3.

Despite its formal appearance, the formula (5.138) is straightforward to evaluate in practice as will become clear in the examples to follow. All that is required is to identify the singular hyperplanes (5.109) for a given GKZ integral, along with the shift produced by the creation operator C_j , and then to form the *b*-function from the product of all singular hyperplanes that can be reached by one application of C_j on any finite starting integral. We emphasise too that, while consideration of singular cases has been used to deduce the form of the *b*-function, the creation operators we subsequently obtain can be used to map finite integrals to finite integrals.

5.4.5 Example

As a simple illustration before turning our efforts to Witten diagrams and Feynman integrals in the following sections, we compute creation operators for the GKZ integral [149, 123]

$$\mathcal{I}_{\gamma} = \int_{\mathbb{R}^2_+} \mathrm{d}z_1 \mathrm{d}z_2 \frac{z_1^{\gamma_1 - 1} z_2^{\gamma_2 - 1}}{(x_1 + x_2 z_1 + x_3 z_2 + x_4 z_1 z_2)^{\gamma_0}}.$$
 (5.140)

On the hypersurface $(x_1, x_2, x_3, x_4) = (1, 1, 1, y)$, this can be directly evaluated in terms of the Gauss hypergeometric function

$$\mathcal{I}_{\gamma}(y) = \frac{\Gamma(\gamma_1)\Gamma(\gamma_2)\Gamma(\gamma_0 - \gamma_1)\Gamma(\gamma_0 - \gamma_2)}{\Gamma(\gamma_0)^2} \,_2F_1(\gamma_1, \gamma_2, \gamma_0; 1 - y). \tag{5.141}$$

Since all shift operators for the Gauss hypergeometric function are known this will allow an easy check of our calculations.

Evaluating the \mathcal{A} -matrix,

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 \\ 0 & 1 & 0 & 1 \\ 0 & 0 & 1 & 1 \end{pmatrix}, \tag{5.142}$$

we can read off the DWI and Euler equations

$$0 = (\gamma_0 + \theta_1 + \theta_2 + \theta_3 + \theta_4) \mathcal{I}_{\gamma}, \quad 0 = (\gamma_1 + \theta_2 + \theta_4) \mathcal{I}_{\gamma}, \quad 0 = (\gamma_2 + \theta_3 + \theta_4) \mathcal{I}_{\gamma}, \quad (5.143)$$

where $\theta_i = x_i \partial_i$ and, from the kernel of the \mathcal{A} -matrix, we find a single toric equation,

$$0 = (\partial_2 \partial_3 - \partial_1 \partial_4) \mathcal{I}_{\gamma}. \tag{5.144}$$

From (5.109), the singular hyperplanes are

$$\gamma_1 = -m_1, \qquad \gamma_2 = -m_2, \qquad \gamma_0 - \gamma_1 = -m_3, \qquad \gamma_0 - \gamma_2 = -m_4, \qquad m_i \in \mathbb{Z}^+$$
(5.145)

as displayed in figure 5.6. As expected, these singularities coincide with the poles of the gamma functions in the numerator of the projected integral (5.141).

The annihilation operators ∂_j send $\gamma \to \gamma'$ while the creation operators \mathcal{C}_j send $\gamma' \to \gamma$, where for each j these parameters are related by

j = 1:	$\gamma_0' = \gamma_0 + 1,$	$\gamma_1' = \gamma_1,$	$\gamma_2' = \gamma_2$	
j = 2:	$\gamma_0' = \gamma_0 + 1,$	$\gamma_1' = \gamma_1 + 1,$	$\gamma_2' = \gamma_2$	
j = 3:	$\gamma_0' = \gamma_0 + 1,$	$\gamma_1' = \gamma_1,$	$\gamma_2' = \gamma_2 + 1,$	
j = 4:	$\gamma_0' = \gamma_0 + 1,$	$\gamma_1' = \gamma_1 + 1,$	$\gamma_2' = \gamma_2 + 1.$	(5.146)

The corresponding b-functions are

$$b_1 = (\gamma_0 - \gamma_1)(\gamma_0 - \gamma_2),$$

$$b_2 = \gamma_1(\gamma_0 - \gamma_2),$$



Figure 5.6: The singular hyperplanes of (5.140).

$$b_3 = \gamma_2(\gamma_0 - \gamma_1),$$

$$b_4 = \gamma_1 \gamma_2.$$
(5.147)

In each case, the factors that appear are the zeros needed to cancel the poles arising when the creation operator moves us from a finite to a singular integral. For b_1 , for example, the shift produced by C_1 can take a finite integral with $\gamma'_0 - \gamma'_i = 1$ to a singular integral with $\gamma_0 - \gamma_i = 0$ for both i = 1 and i = 2, as we see from (5.145). The zeros of b_1 then cancel these singularities so that the action (5.132) of C_1 on a finite integral is always finite. For b_2 , the shifts produced by C_2 can take a finite integral with $\gamma'_2 = 1$ to a singular integral with $\gamma_2 = 0$, and a finite integral with $\gamma'_0 - \gamma'_1 = 1$ to a singular integral with $\gamma_0 - \gamma_1 = 0$, with these singularities again being cancelled by the zeros of b_2 . Note that the action of C_2 leaves γ'_1 and $\gamma'_0 - \gamma'_2$ unchanged hence no further singularities arise, and hence no further factors in b_2 . One can further check that the *b*-functions (5.147) are consistent with the general formula (5.138).

From (5.130) plus the DWI and Euler equations (5.143), we now have, for example,

$$\mathcal{C}_1 \partial_1 \mathcal{I}_{\gamma} = (\gamma_0 - \gamma_1)(\gamma_0 - \gamma_2) \mathcal{I}_{\gamma} = (\theta_1 + \theta_3)(\theta_1 + \theta_2) \mathcal{I}_{\gamma}$$
$$= (x_1 \partial_1 + x_1^2 \partial_1^2 + x_1 x_2 \partial_1 \partial_2 + x_1 x_3 \partial_1 \partial_3 + x_2 x_3 \partial_2 \partial_3) \mathcal{I}_{\gamma}.$$
(5.148)

By inspection, every term in the final line contains an explicit factor of ∂_1 except for the last, but this can be replaced by $x_2 x_3 \partial_1 \partial_4$ using the toric equation (5.144). This gives us the desired factorisation

$$C_1 = x_1 + x_1^2 \partial_1 + x_1 x_2 \partial_2 + x_1 x_3 \partial_3 + x_2 x_3 \partial_4$$

= $x_1 (1 + \theta_1 + \theta_2 + \theta_3) + x_2 x_3 \partial_4.$ (5.149)

In an identical fashion, we obtain

$$C_{2} = x_{2}(1 + \theta_{1} + \theta_{2} + \theta_{4}) + x_{1}x_{4}\partial_{3},$$

$$C_{3} = x_{3}(1 + \theta_{1} + \theta_{3} + \theta_{4}) + x_{1}x_{4}\partial_{2},$$

$$C_{4} = x_{4}(1 + \theta_{2} + \theta_{3} + \theta_{4}) + x_{2}x_{3}\partial_{1}.$$
(5.150)

Finally, in order to understand their action on (5.141), these creation operators can be projected to the 'physical' hypersurface $(x_1, x_2, x_3, x_4) = (1, 1, 1, y)$. For this we use the DWI and Euler equations (5.143) evaluated on this hypersurface, which can be re-arranged so as to eliminate all derivatives apart from ∂_y :

$$\partial_{1}\mathcal{I}_{\gamma'}(y) = (-\gamma'_{0} + \gamma'_{1} + \gamma'_{2} + \theta_{y})\mathcal{I}_{\gamma'}(y),$$

$$\partial_{2}\mathcal{I}_{\gamma'}(y) = -(\gamma'_{1} + \theta_{y})\mathcal{I}_{\gamma'}(y),$$

$$\partial_{3}\mathcal{I}_{\gamma'}(y) = -(\gamma'_{2} + \theta_{y})\mathcal{I}_{\gamma'}(y).$$
(5.151)

Notice here that as the creation operators act on the integral with parameters γ' by our definition (5.132), we need to use these parameters here. With the aid of these equations, the creation operators project to

$$\begin{aligned} \mathcal{C}_{1}^{\mathrm{ph}} &= 1 - \gamma_{0}' + (1 - y)\partial_{y}, \\ \mathcal{C}_{2}^{\mathrm{ph}} &= 1 - \gamma_{0}' + (1 - y)(\gamma_{2}' + \theta_{y}), \\ \mathcal{C}_{3}^{\mathrm{ph}} &= 1 - \gamma_{0}' + (1 - y)(\gamma_{3}' + \theta_{y}), \\ \mathcal{C}_{4}^{\mathrm{ph}} &= 1 - \gamma_{0}' + (1 - y)(\gamma_{1}' + \gamma_{2}' - 1 + \theta_{y}), \end{aligned}$$
(5.152)

where the 'ph' superscript indicates the operators expressed in physical variables. From (5.132), we then have, for example,

$$\mathcal{C}_{1}^{\mathrm{ph}}\mathcal{I}_{\gamma_{0}',\gamma_{1}',\gamma_{2}'}(y) = -\gamma_{0}^{-1}(\gamma_{0}-\gamma_{1})(\gamma_{0}-\gamma_{2})\mathcal{I}_{\gamma_{0},\gamma_{1},\gamma_{2}}(y)$$

$$= -(\gamma_{0}'-1)^{-1}(\gamma_{0}'-\gamma_{1}'-1)(\gamma_{0}'-\gamma_{2}'-1)\mathcal{I}_{\gamma_{0}'-1,\gamma_{1}',\gamma_{2}'}, \qquad (5.153)$$

since here the creation operator shifts $\gamma'_0 \to \gamma_0 = \gamma'_0 - 1$ while $\gamma'_i = \gamma_i$ for i = 1, 2. Noting the presence of the gamma functions in (5.141), this corresponds to

$$\left(1 - \gamma_0' + (1 - y)\partial_y\right) {}_2F_1(\gamma_1', \gamma_2', \gamma_0'; 1 - y) = (1 - \gamma_0') {}_2F_1(\gamma_1', \gamma_2', \gamma_0' - 1; 1 - y)$$
(5.154)

which indeed follows from standard relations for $_2F_1$ (see e.g., equation 15.5.4 of [160]).

Taking into account the shifts (5.146), for the remaining operators we find

$$\begin{aligned} \mathcal{C}_{2}^{\mathrm{ph}}\mathcal{I}_{\gamma'_{0},\gamma'_{1},\gamma'_{2}}(y) &= -\gamma_{0}^{-1}\gamma_{1}(\gamma_{0}-\gamma_{2})\mathcal{I}_{\gamma_{0},\gamma_{1},\gamma_{2}} \\ &= -(\gamma'_{0}-1)^{-1}(\gamma'_{1}-1)(\gamma'_{0}-\gamma'_{2}-1)\mathcal{I}_{\gamma'_{0}-1,\gamma'_{1}-1,\gamma'_{2}}(y), \\ \mathcal{C}_{3}^{\mathrm{ph}}\mathcal{I}_{\gamma'_{0},\gamma'_{1},\gamma'_{2}}(y) &= -\gamma_{0}^{-1}\gamma_{2}(\gamma_{0}-\gamma_{1})\mathcal{I}_{\gamma_{0},\gamma_{1},\gamma_{2}} \\ &= -(\gamma'_{0}-1)^{-1}(\gamma'_{2}-1)(\gamma'_{0}-\gamma'_{1}-1)\mathcal{I}_{\gamma'_{0}-1,\gamma'_{1},\gamma'_{2}-1}(y), \\ \mathcal{C}_{4}^{\mathrm{ph}}\mathcal{I}_{\gamma'_{0},\gamma'_{1},\gamma'_{2}}(y) &= -\gamma_{0}^{-1}\gamma_{1}\gamma_{2}\mathcal{I}_{\gamma_{0},\gamma_{1},\gamma_{2}} \\ &= -(\gamma'_{0}-1)^{-1}(\gamma'_{1}-1)(\gamma'_{2}-1)\mathcal{I}_{\gamma'_{0}-1,\gamma'_{1}-1,\gamma'_{2}-1}(y). \end{aligned}$$
(5.155)

These can again be verified using standard shift identities and contiguity relations for the Gauss hypergeometric function.

5.5 Creation operators for Witten diagrams

The correlators of holographic conformal field theories at strong coupling can be computed via Witten diagrams in anti-de Sitter spacetime. As the evaluation of these diagrams is nontrivial, particularly in momentum space, it is important to identify classes of shift operators connecting known 'seed' solutions to a broader family of correlators.

In this section, we construct novel creation operators for Witten diagrams in momentum space. (Results for the position-space contact diagram, or holographic *D*-function, are given in appendix C.2.) Starting with the contact diagram, we derive explicit creation operators at 3- and 4-points, though the method extends to higher points. We also show, again at 3- and 4-points, how to construct operators that shift the scaling dimensions while preserving the spacetime dimension.

A case of particular interest, given the close connection to cosmological correlators, is the 4-point exchange diagram. Here, a class of weight-shifting operators is known connecting exchange diagrams with different external scaling dimensions [83, 30], but subject to two restrictions [36]: first, these operators map an exchange diagram with *non-derivative* vertices to one with *derivative* vertices; and second, they work only for a special set of initial scaling dimensions. While these results are sufficient for cosmologies where the inflaton is a derivatively-coupled massless scalar, finding further generalisations is highly desirable.

A key problem, therefore, is to find a shift operator connecting exchange diagrams with *non-derivative* vertices to new exchange diagrams, with shifted operator dimensions, but still with *non-derivative* vertices. This operator should moreover be applicable for diagrams of arbitrary initial scaling dimensions. In section 5.5.6, we derive such an operator.

5.5.1 Definitions

In momentum space, the n-point contact Witten diagram is

$$i_{[d;\,\Delta_1,\,\dots,\,\Delta_n]} = \int_0^\infty \mathrm{d}z \, z^{-d-1} \prod_{i=1}^n \mathcal{K}_{[\Delta_i]}(z,p_i)$$
(5.156)

where d is the boundary spacetime dimension of the CFT, Δ_i is the scaling dimension of the scalar operator \mathcal{O}_i , and the bulk-to-boundary propagator

$$\mathcal{K}_{[\Delta_i]}(z, p_i) = \frac{z^{\frac{d}{2}} p_i^{\beta_i} K_{\beta_i}(p_i z)}{2^{\beta_i - 1} \Gamma(\beta_i)}, \qquad \beta_i = \Delta_i - \frac{d}{2}.$$
(5.157)

Since the modified Bessel-K function is invariant under changing the sign of its index, note we have the shadow relation

$$i_{[d;\,\Delta_1,\,\dots,\,\Delta_n]}\Big|_{\Delta_i\to d-\Delta_i} = \frac{4^{\beta_i}\Gamma(\beta_i)}{\Gamma(-\beta_i)} p_i^{-2\beta_i} i_{[d;\,\Delta_1,\,\dots,\,\Delta_n]}.$$
(5.158)



Figure 5.7: Witten diagrams representing the contact and exchange 4-point diagram $i_{[\Delta_1 \Delta_2 \Delta_3 \Delta_4]}$ and $i_{[\Delta_1 \Delta_2, \Delta_3 \Delta_4 x \Delta_x]}$ given by the integrals (5.156) and (5.159).

In addition to the contact diagram, we will discuss the 4-point s-channel exchange diagram shown in figure 5.7,

$$i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = \int_0^\infty \mathrm{d}z \, z^{-d-1} \mathcal{K}_{[\Delta_1]}(z,p_1) \mathcal{K}_{[\Delta_2]}(z,p_2)$$

$$\times \int_0^\infty \mathrm{d}\zeta \, \zeta^{-d-1} \mathcal{G}_{[\Delta_x]}(z,s;\zeta) \mathcal{K}_{[\Delta_3]}(\zeta,p_3) \mathcal{K}_{[\Delta_4]}(\zeta,p_4),$$
(5.159)

where Δ_x is the dimension of the exchanged operator and $s^2 = (\mathbf{p}_1 + \mathbf{p}_2)^2$. The bulk-tobulk propagator in this expression is

$$\mathcal{G}_{[\Delta_x]}(z,s;\zeta) = \begin{cases} (z\zeta)^{\frac{d}{2}} I_{\beta_x}(sz) K_{\beta_x}(s\zeta) & \text{for } z < \zeta, \\ (z\zeta)^{\frac{d}{2}} K_{\beta_x}(sz) I_{\beta_x}(s\zeta) & \text{for } z > \zeta, \end{cases}$$
(5.160)

with I_{β} and K_{β} representing modified Bessel functions and $\beta_x = \Delta_x - d/2$. Where necessary, these integrals can be regulated by infinitesimally shifting the operator dimensions and spacetime dimension d so as to ensure convergence [36].

5.5.2 GKZ representation of the contact diagram

The GKZ representation for the *n*-point momentum-space contact diagram can be evaluated analogously to that for the triple-K integral (see page 99). This yields the GKZ integral

$$\mathcal{I}_{\gamma} = \left(\prod_{i=1}^{n} \int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1}\right) \left[\sum_{j=1}^{n} \left(\frac{x_{j}}{z_{j}} + \bar{x}_{j} z_{j}\right)\right]^{-\gamma_{0}}$$
(5.161)

with the contact diagram being

$$i_{[d;\,\Delta_1,\,\dots,\,\Delta_n]} = 2^{\gamma_0} \Gamma(\gamma_0) \Big(\prod_{i=1}^n \frac{1}{2^{\gamma_i} \Gamma(\gamma_i)}\Big) \mathcal{I}_{\gamma}$$
(5.162)

with parameters

$$\gamma_0 = \left(\frac{n}{2} - 1\right)d, \qquad \gamma_i = \Delta_i - \frac{d}{2} = \beta_i, \tag{5.163}$$

and physical hypersurface

$$x_i = p_i^2, \qquad \bar{x}_i = 1, \qquad i = 1, \dots, n.$$
 (5.164)

Our notation $\boldsymbol{x} = (x_i, \bar{x}_i)$ for the GKZ variables here and in (5.161) is simply a convenience designed to simplify the form of the Euler and toric equations as we will see below; \bar{x}_i should be regarded as an independent dynamical variable equivalent to x_{i+n} in the notation of the previous section.

The $(n+1) \times 2n$ dimensional *A*-matrix for the integral (5.161) is now

$$\mathcal{A} = \begin{pmatrix} \mathbf{1} & \mathbf{1} \\ -\mathbb{I}_n & \mathbb{I}_n \end{pmatrix}$$
(5.165)

where **1** is the *n*-dimensional row vector of 1s and \mathbb{I}_n is the $n \times n$ identity matrix. (Again, we are departing from the notation of the previous section where *n* referred to the number of columns in the \mathcal{A} -matrix, reserving *n* now for the number of points.) Writing

$$\partial_i = \frac{\partial}{\partial x_i}, \qquad \bar{\partial}_i = \frac{\partial}{\partial \bar{x}_i}, \qquad \theta_i = x_i \partial_i, \qquad \bar{\theta}_i = \bar{x}_i \bar{\partial}_i, \qquad (5.166)$$

the Euler equations are

$$0 = (\gamma_i - \theta_i + \bar{\theta}_i)\mathcal{I}_{\gamma}, \qquad i = 1, \dots, n,$$
(5.167)

while the DWI is

$$0 = \left(\gamma_0 + \sum_{i=1}^n (\theta_i + \bar{\theta}_i)\right) \mathcal{I}_{\gamma}.$$
(5.168)

In addition, we have the toric equations

$$0 = (\partial_i \bar{\partial}_i - \partial_j \bar{\partial}_j) \mathcal{I}_{\gamma}, \qquad i \neq j = 1, \dots, n.$$
(5.169)

These can easily be verified by noting that $\partial_i \partial_i$ sends $\gamma_0 \to \gamma_0 + 2$ but makes no change to the power of z_i appearing in the numerator of (5.161), hence the two terms in (5.169) cancel.

It is well known that the contact diagram satisfies the equation,

$$0 = (K_i - K_j)i_{[d, \Delta_1, \dots, \Delta_n]} \qquad \forall \ i \neq j$$
(5.170)

where K_i is the Bessel operator

$$K_i = \partial_{p_i}^2 + \frac{(1 - 2\gamma_i)}{p_i} \partial_{p_i} = \partial_i (\theta_i - \gamma_i).$$
(5.171)

To see this from a GKZ perspective, we use the Euler and toric equations to show that

$$(K_i - K_j)\mathcal{I}_{\gamma} = (\partial_i \bar{\theta}_i - \partial_j \bar{\theta}_j)\mathcal{I}_{\gamma} = (\bar{x}_i \partial_i \bar{\partial}_i - \bar{x}_j \partial_j \bar{\partial}_j)\mathcal{I}_{\gamma} = (\bar{x}_i - \bar{x}_j)\partial_i \bar{\partial}_i \mathcal{I}_{\gamma}.$$
 (5.172)

Upon projecting to the physical hypersurface (5.164), the right-hand side now vanishes. Finally, we observe that the shadow relation (5.158) uplifts to

$$\mathcal{I}_{\gamma}\Big|_{\gamma_i \to -\gamma_i} = \left(\frac{\bar{x}_i}{x_i}\right)^{\gamma_i} \mathcal{I}_{\gamma} \tag{5.173}$$

for any i = 1, ..., n in GKZ variables. This can be seen by evaluating the right-hand side with the substitution $z_i = x_i/(\bar{x}_i z'_i)$ in (5.161).

5.5.3 Creation and annihilation operators

The action of the annihilation operators is

$$\partial_{i}\mathcal{I}_{\gamma} = -\gamma_{0}\mathcal{I}_{\gamma}\Big|_{\gamma_{0}\to\gamma_{0}+1,\,\gamma_{i}\to\gamma_{i}-1}, \qquad \bar{\partial}_{i}\mathcal{I}_{\gamma} = -\gamma_{0}\mathcal{I}_{\gamma}\Big|_{\gamma_{0}\to\gamma_{0}+1,\,\gamma_{i}\to\gamma_{i}+1} \tag{5.174}$$

for any i = 1, ..., n. After projecting to the physical hypersurface (5.164), up to numerical factors ∂_i and $\overline{\partial}_i$ become the operators \mathcal{L}_i and \mathcal{R}_i respectively, as defined in (5.124). Note that due to the shadow relation (5.173) (or re-arranging the Euler equation (5.167)), we have

$$\bar{\partial}_i \mathcal{I}_{\gamma} = \left(\frac{\bar{x}_i}{x_i}\right)^{-\gamma_i - 1} \partial_i \left(\frac{\bar{x}_i}{x_i}\right)^{\gamma_i} \mathcal{I}_{\gamma}.$$
(5.175)

In physical variables, this projects to

$$\mathcal{R}_i = p_i^{2(\beta_i+1)} \mathcal{L}_i \, p_i^{-2\beta_i}. \tag{5.176}$$

The action of the creation operators is the inverse of that in (5.174), namely

$$\mathcal{C}_i: \quad \gamma_0 \to \gamma_0 - 1, \quad \gamma_i \to \gamma_i + 1, \qquad \bar{\mathcal{C}}_i: \quad \gamma_0 \to \gamma_0 - 1, \quad \gamma_i \to \gamma_i - 1, \tag{5.177}$$

where all remaining γ_j for $j \neq i$ stay the same. By virtue of the shadow relation (5.158), however, it suffices to construct only C_i since

$$\bar{\mathcal{C}}_i \mathcal{I}_{\gamma} = \left(\frac{\bar{x}_i}{x_i}\right)^{1-\gamma_i} \mathcal{C}_i \Big|_{\gamma_i \to -\gamma_i} \left(\frac{\bar{x}_i}{x_i}\right)^{\gamma_i} \mathcal{I}_{\gamma}.$$
(5.178)

To construct C_i , we first need to identify the singular hyperplanes of \mathcal{I}_{γ} . These can be found either by expanding the integrand of (5.156) about the lower limit z = 0 and looking for the appearance of z^{-1} pole terms (see [45], and the example on page 103), or by using the formula (5.109) based on the Newton polytope. Here, the Newton polytope takes the form of an *n*-dimensional cross-polytope with vertices at $\pm e_j$ for every basis vector $(e_j)_k = \delta_{jk}$ and 2^n facets with outward normals $\mathbf{n} = (\sigma_1, \ldots, \sigma_n)^T$ for every possible independent choice of $\sigma_j = \pm 1$. From (5.96) and (5.108), $n_0 = -1$ and $\delta = 2$ for every facet, hence the singular hyperplanes are

$$0 = -\gamma_0 + \sum_{j=1}^n \sigma_j \gamma_j - 2m, \qquad m \in \mathbb{Z}^+.$$
(5.179)

Given the action of C_i in (5.177), the only way this operator can shift us from a finite to a singular integral is if $\sigma_i = +1$ so that *m* increases by one. The corresponding *b*-function is then

$$b_i(\boldsymbol{\gamma}) = \prod_{\{\sigma_j = \pm 1\}} \frac{1}{2} (-\gamma_0 + \gamma_i + \sum_{j \neq i} \sigma_j \gamma_j), \qquad (5.180)$$

where the product runs over every possible choice of signs for all $j \neq i$. Using the Euler equations, this gives

$$B_i(\theta,\bar{\theta}) = \prod_{\{\sigma_j=\pm 1\}} \left(\theta_i + \sum_{j\neq i} (\delta_{\sigma_j,+1}\theta_j + \delta_{\sigma_j,-1}\bar{\theta}_j) \right).$$
(5.181)

For convenience, to eliminate an overall numerical factor in this expression we inserted factors of one-half in (5.180). Overall, this is simply a trivial rescaling of both the creation operator and the *b*-function.

For the 3-point function, for example, these formulae evaluate to

$$b_1(\boldsymbol{\gamma}) = \frac{1}{16}(-\gamma_0 + \gamma_1 + \gamma_2 + \gamma_3)(-\gamma_0 + \gamma_1 - \gamma_2 + \gamma_3)(-\gamma_0 + \gamma_1 + \gamma_2 - \gamma_3)(-\gamma_0 + \gamma_1 - \gamma_2 - \gamma_3).$$
(5.182)

and

$$B_1(\theta,\bar{\theta}) = (\theta_1 + \theta_2 + \theta_3)(\theta_1 + \bar{\theta}_2 + \theta_3)(\theta_1 + \theta_2 + \bar{\theta}_3)(\theta_1 + \bar{\theta}_2 + \bar{\theta}_3).$$
(5.183)

Recalling now the creation operators obey

$$C_i \partial_i \mathcal{I}_{\gamma} = b_i(\gamma) \mathcal{I}_{\gamma} = B_i(\theta, \bar{\theta}) \mathcal{I}_{\gamma}$$
(5.184)

the idea is to expand out as^{11}

$$B_i(\theta,\bar{\theta}) = Q_i(\theta,\bar{\theta})\theta_i + \sum_{j\neq i} Q_j(\theta,\bar{\theta})\theta_j\bar{\theta}_j, \qquad (5.185)$$

where without loss of generality we can choose all $Q_j(\theta, \bar{\theta})$ for $j \neq i$ to be independent of both θ_i and $\bar{\theta}_i$. (Note from (5.181) that $B_i(\theta, \bar{\theta})$ is automatically independent of $\bar{\theta}_i$.) We then use the toric equations (5.169) to re-express

$$\theta_j \bar{\theta}_j \mathcal{I}_{\gamma} = x_j \bar{x}_j \partial_j \bar{\partial}_j \mathcal{I}_{\gamma} = x_j \bar{x}_j \partial_i \bar{\partial}_i \mathcal{I}_{\gamma}$$
(5.186)

so that

$$B_i(\theta,\bar{\theta})\mathcal{I}_{\gamma} = \left[Q_i(\theta,\bar{\theta})x_i + \sum_{j\neq i} Q_j(\theta,\bar{\theta})x_j\bar{x}_j\bar{\partial}_i\right]\partial_i\mathcal{I}_{\gamma},\tag{5.187}$$

¹¹A factorisation of this form always exists as can be seen recursively in the number of points *n*. Once all θ_i -dependence has been gathered into $Q_i\theta_i$, let us write the remainder at *n*-points as $b_i^{(n)} = B_i^{(n)}|_{\theta_i \to 0}$. Multiplying out all the factors containing θ_n , and, separately, all the factors containing $\bar{\theta}_n$, we obtain $b_i^{(n)} = ((\ldots)\theta_n + b_i^{(n-1)})((\ldots)\bar{\theta}_n + b_i^{(n-1)}) = (\ldots)\theta_n\bar{\theta}_n + (\ldots)b_i^{(n-1)}$, where $b_i^{(n-1)}$ is independent of θ_n and $\bar{\theta}_n$. Thus, if the decomposition $b_i^{(n-1)} = \sum_{j\neq i}^{n-1} Q_j^{(n-1)}\theta_j\bar{\theta}_j$ exists at (n-1)-points, then it also exists at *n*-points: $b_i^{(n)} = \sum_{j\neq i}^n Q_j^{(n)}\theta_j\bar{\theta}_j$.

This yields the creation operators

$$C_i = Q_i(\theta, \bar{\theta}) x_i + \sum_{j \neq i} Q_j(\theta, \bar{\theta}) x_j \bar{x}_j \bar{\partial}_i.$$
(5.188)

To project from the GKZ space to the physical space spanned by the momenta, we first re-write (suppressing arguments for clarity)

$$Q_i x_i = x_i Q_i \Big|_{\theta_i \to \theta_i + 1} \tag{5.189}$$

$$Q_j x_j \bar{x}_j \bar{\partial}_i = x_j \bar{x}_j \bar{\partial}_i Q_j \Big|_{\theta_j \to \theta_j + 1, \, \bar{\theta}_j \to \bar{\theta}_j + 1}$$
(5.190)

where for (5.190) we recall the Q_j are independent of both θ_i and $\bar{\theta}_i$. We then project to the physical hypersurface (5.164) by using the Euler equations (5.167) to replace $\bar{\theta}_k \to \theta_k - \gamma_k$ for all $k = 1, \ldots, n$ (which is justified since after the re-arrangements (5.189)-(5.190) all $\bar{\theta}_k$ act directly on \mathcal{I}_{γ}) and set all $\bar{x}_k \to 1$. Note also that $\bar{\partial}_i = \bar{\theta}_i$ on the physical hypersurface since $\bar{x}_i = 1$, hence we can also replace $\bar{\partial}_i \to \theta_i - \gamma_i$. The result is

$$\mathcal{C}_{i}^{\mathrm{ph}} = x_{i}Q_{i}\Big|_{\theta_{i}\to\theta_{i}+1,\,\bar{\theta}_{k}\to\theta_{k}-\gamma_{k}} + (\theta_{i}-\gamma_{i})\sum_{j\neq i}x_{j}Q_{j}\Big|_{\theta_{j}\to\theta_{j}+1,\,\bar{\theta}_{k}\to\theta_{k}-\gamma_{k}+\delta_{kj}}$$
(5.191)

where the replacement on $\bar{\theta}_k$ applies to all the $\bar{\theta}$ variables present. As previously, the superscript 'ph' denotes the operator expressed in physical variables. From the shadow relation (5.158), we also have

$$\bar{\mathcal{C}}_{i}^{\mathrm{ph}} = x_{i}^{\gamma_{i}-1} \mathcal{C}_{i}^{\mathrm{ph}} \Big|_{\gamma_{i} \to -\gamma_{i}} x_{i}^{-\gamma_{i}} = Q_{i} \Big|_{\theta_{i} \to \theta_{i} - \gamma_{i} + 1, \bar{\theta}_{k} \to \theta_{k} - \gamma_{k}} + \sum_{j \neq i} x_{j} \partial_{i} Q_{j} \Big|_{\theta_{j} \to \theta_{j} + 1, \bar{\theta}_{k} \to \theta_{k} - \gamma_{k} + \delta_{kj}}.$$
(5.192)

Together, these expressions gives us the creation operators in terms of the physical variables

$$x_k = p_k^2, \qquad \theta_k = x_k \partial_k = \frac{1}{2} p_k \partial_{p_k},$$
(5.193)

From (5.132), their action is

$$\mathcal{C}_{i}\mathcal{I}_{\{\gamma_{0},\gamma_{i}\}} = -(\gamma_{0}-1)^{-1}b_{i}(\gamma_{0}-1,\gamma_{i}+1)\mathcal{I}_{\{\gamma_{0}-1,\gamma_{i}+1\}},$$

$$\bar{\mathcal{C}}_{i}\mathcal{I}_{\{\gamma_{0},\gamma_{i}\}} = -(\gamma_{0}-1)^{-1}\bar{b}_{i}(\gamma_{0}-1,\gamma_{i}-1)\mathcal{I}_{\{\gamma_{0}-1,\gamma_{i}-1\}}.$$
(5.194)

The shift in *b*-function arguments on the right-hand sides here reflects the fact that, in replacing $\bar{\theta}_k \to \theta_k - \gamma_k$ in the projection step above, we are taking the creation operator to act on the integral $\mathcal{I}_{\{\gamma_0, \gamma_i\}}$. This is equivalent to eliminating γ from (5.132) using (5.133) then relabelling $\gamma' \to \gamma$.

3-point creation operator

Let us find the creation operator C_1 for the 3-point function via the procedure outlined above. Starting from the expression for $B_1(\theta, \bar{\theta})$ in (5.183), we decompose

$$B_1(\theta,\bar{\theta}) = Q_1\theta_1 + Q_2\theta_2\bar{\theta}_2 + Q_3\theta_3\bar{\theta}_3$$
(5.195)

where

$$Q_1 = (\theta_1 + u_2 + u_3) \big((\theta_1 + u_2)(\theta_1 + u_3) + 2(v_2 + v_3) \big), \tag{5.196}$$

$$Q_2 = (u_2 + u_3)u_3 + v_2 - v_3, (5.197)$$

$$Q_3 = (u_2 + u_3)u_2 - v_2 + v_3, (5.198)$$

with

$$u_i = \theta_i + \theta_i, \qquad v_i = \theta_i \theta_i, \qquad i = 2, 3.$$
(5.199)

Note that Q_2 and Q_3 are independent of θ_1 and all coefficients are independent of $\overline{\theta}_1$. We have also chosen Q_2 and Q_3 to preserve the 2 \leftrightarrow 3 symmetry though this is not essential.

Re-iterating the steps above, making use of (5.186), we have

$$B_1(\theta,\bar{\theta})\mathcal{I}_{\gamma} = \left(Q_1x_1 + Q_2x_2\bar{x}_2\bar{\partial}_1 + Q_3x_3\bar{x}_3\bar{\partial}_1\right)\partial_1\mathcal{I}_{\gamma} = \mathcal{C}_1\partial_1\mathcal{I}_{\gamma}$$
(5.200)

yielding the creation operator C_1 in GKZ space. Moving the Q_k to the right, this can equivalently be written

$$C_{1} = x_{1}Q_{1}\Big|_{\theta_{1} \to \theta_{1}+1} + x_{2}\bar{x}_{2}\bar{\partial}_{1}Q_{2}\Big|_{\theta_{2} \to \theta_{2}+1, \bar{\theta}_{2} \to \bar{\theta}_{2}+1} + x_{3}\bar{x}_{3}\bar{\partial}_{1}Q_{3}\Big|_{\theta_{3} \to \theta_{3}+1, \bar{\theta}_{3} \to \bar{\theta}_{3}+1}$$
(5.201)

Since shifting $\theta_i \to \theta_i + 1$ and $\bar{\theta}_i \to \bar{\theta}_i + 1$ is equivalent to $u_i \to u_i + 2$ and $v_i \to 1 + u_i + v_i$, this is

$$C_{1} = x_{1}(\theta_{1} + 1 + u_{2} + u_{3}) ((\theta_{1} + 1 + u_{2})(\theta_{1} + 1 + u_{3}) + 2(v_{2} + v_{3})) + x_{2}\bar{x}_{2}\bar{\partial}_{1} (1 + u_{2} + v_{2} - v_{3} + (u_{2} + u_{3} + 2)u_{3}) + x_{3}\bar{x}_{3}\bar{\partial}_{1} (1 + u_{3} + v_{3} - v_{2} + (u_{2} + u_{3} + 2)u_{2}).$$
(5.202)

Finally, to project to the physical hypersurface, we set

 $\bar{x}_i \to 1, \qquad \bar{\theta}_i \to \theta_i - \gamma_i, \qquad \bar{\partial}_i \to \theta_i - \gamma_i$ (5.203)

which sends $u_i \to 2\theta_i - \gamma_i$ and $v_i \to \theta_i(\theta_i - \gamma_i)$, yielding

$$C_{1} = x_{1}(\theta_{1} + 1 + 2\theta_{2} + 2\theta_{3} - \gamma_{2} - \gamma_{3}) \times$$

$$\times \left[(\theta_{1} + 1 + 2\theta_{2} - \gamma_{2})(\theta_{1} + 1 + 2\theta_{3} - \gamma_{3}) + 2\theta_{2}(\theta_{2} - \gamma_{2}) + 2\theta_{3}(\theta_{3} - \gamma_{3}) \right] + (\theta_{1} - \gamma_{1}) \times$$

$$\times \left[x_{2} \left(1 + 2\theta_{2} - \gamma_{2} + \theta_{2}(\theta_{2} - \gamma_{2}) - \theta_{3}(\theta_{3} - \gamma_{3}) + (2\theta_{2} - \gamma_{2} + 2\theta_{3} - \gamma_{3} + 2)(2\theta_{3} - \gamma_{3}) \right) + x_{3} \left(1 + 2\theta_{3} - \gamma_{3} + \theta_{3}(\theta_{3} - \gamma_{3}) - \theta_{2}(\theta_{2} - \gamma_{2}) + (2\theta_{2} - \gamma_{2} + 2\theta_{3} - \gamma_{3} + 2)(2\theta_{2} - \gamma_{2}) \right) \right]$$

$$(5.204)$$

Of course, this result also follows from (5.191) directly. We can simplify somewhat further

by using the DWI evaluated on the physical hypersurface,

$$0 = \left(\theta_1 + \theta_2 + \theta_3 + \frac{1}{2}(\gamma_0 - \gamma_1 - \gamma_2 - \gamma_3)\right) \mathcal{I}_{\gamma}.$$
 (5.205)

This gives the alternative form

$$C_{1} = -\frac{x_{1}}{2}(\theta_{1} - 1 + \gamma_{0} - \gamma_{1}) \left[2\theta_{1}^{2} + 2(\gamma_{0} - \gamma_{1})\theta_{1} + (\gamma_{0} - \gamma_{1} - 1)^{2} + 1 - \gamma_{2}^{2} - \gamma_{3}^{2} \right] (5.206) + (\theta_{1} - \gamma_{1}) \times \times \left[x_{2} \left(1 + 2\theta_{2} - \gamma_{2} + \theta_{2}(\theta_{2} - \gamma_{2}) - \theta_{3}(\theta_{3} - \gamma_{3}) - (2\theta_{1} - \gamma_{1} + \gamma_{0} - 2)(2\theta_{3} - \gamma_{3}) \right) + x_{3} \left(1 + 2\theta_{3} - \gamma_{3} + \theta_{3}(\theta_{3} - \gamma_{3}) - \theta_{2}(\theta_{2} - \gamma_{2}) - (2\theta_{1} - \gamma_{1} + \gamma_{0} - 2)(2\theta_{2} - \gamma_{2}) \right) \right].$$

The action of this creation operator is

$$C_1 \mathcal{I}_{\{\gamma_0,\gamma_1\}} = -(\gamma_0 - 1)^{-1} b_1(\gamma_0 - 1, \gamma_1 + 1) \mathcal{I}_{\{\gamma_0 - 1, \gamma_1 + 1\}}$$
(5.207)

where

$$b_1(\gamma_0 - 1, \gamma_1 + 1) = b_1(\boldsymbol{\gamma})\Big|_{\gamma_0 \to \gamma_0 - 1, \gamma_1 \to \gamma_1 + 1} = \frac{1}{16} \Big[\big((2 - \gamma_0 + \gamma_1)^2 - \gamma_2^2 - \gamma_3^2 \big)^2 - 4\gamma_2^2 \gamma_3^2 \Big]$$
(5.208)

using $b_1(\boldsymbol{\gamma})$ as given in (5.182).

4-point creation operator

From (5.180), the 4-point *b*-function is

$$b_{1}(\boldsymbol{\gamma}) = 2^{-8}(-\gamma_{0} + \gamma_{1} + \gamma_{2} + \gamma_{3} + \gamma_{4})(-\gamma_{0} + \gamma_{1} - \gamma_{2} + \gamma_{3} + \gamma_{4})(-\gamma_{0} + \gamma_{1} + \gamma_{2} - \gamma_{3} + \gamma_{4}) \times (-\gamma_{0} + \gamma_{1} + \gamma_{2} + \gamma_{3} - \gamma_{4})(-\gamma_{0} + \gamma_{1} - \gamma_{2} - \gamma_{3} + \gamma_{4})(-\gamma_{0} + \gamma_{1} - \gamma_{2} + \gamma_{3} - \gamma_{4}) \times (-\gamma_{0} + \gamma_{1} + \gamma_{2} - \gamma_{3} - \gamma_{4})(-\gamma_{0} + \gamma_{1} - \gamma_{2} - \gamma_{3} - \gamma_{4})$$
(5.209)

which, after use of the Euler equations and DWI, corresponds to

$$B_{1} = (\theta_{1} + \theta_{2} + \theta_{3} + \theta_{4})(\theta_{1} + \bar{\theta}_{2} + \theta_{3} + \theta_{4})(\theta_{1} + \theta_{2} + \bar{\theta}_{3} + \theta_{4})(\theta_{1} + \theta_{2} + \theta_{3} + \bar{\theta}_{4}) \times (\theta_{1} + \bar{\theta}_{2} + \bar{\theta}_{3} + \theta_{4})(\theta_{1} + \bar{\theta}_{2} + \theta_{3} + \bar{\theta}_{4})(\theta_{1} + \theta_{2} + \bar{\theta}_{3} + \bar{\theta}_{4})(\theta_{1} + \bar{\theta}_{2} + \bar{\theta}_{3} + \bar{\theta}_{4})$$

$$(5.210)$$

consistent with (5.181). We wish to decompose this as

$$B_1(\theta,\bar{\theta}) = Q_1\theta_1 + Q_2\theta_2\bar{\theta}_2 + Q_3\theta_3\bar{\theta}_3 + Q_4\theta_4\bar{\theta}_4$$
(5.211)

where Q_2 , Q_3 and Q_4 are independent of θ_1 .

Let us deal with the Q_1 term first. Denoting the eight factors in (5.210) as R_m for $m = 1, \ldots, 8$, we have

$$B_1\Big|_{\theta_1=0} = \prod_{m=1}^8 (-\theta_1 + R_m) = \sum_{m=0}^8 \sigma_{(m)}(R)(-\theta_1)^{8-m} = B_1 + \sum_{m=0}^7 \sigma_{(m)}(R)(-\theta_1)^{8-m}$$
(5.212)

where $\sigma_{(m)}(R)$ is the *m*th elementary symmetric polynomial in the R_m . Rearranging then gives

$$Q_1 = \theta_1^{-1} \left(B_1 - B_1 \big|_{\theta_1 = 0} \right) = \sum_{m=0}^7 \sigma_{(m)}(R) (-\theta_1)^{7-m}, \tag{5.213}$$

and since θ_1 appears in each of the factors in (5.210),

$$Q_1 x_1 = x_1 Q_1 \Big|_{\theta_1 \to \theta_1 + 1} = \sum_{m=0}^7 \sigma_{(m)} (1+R) (-1-\theta_1)^{7-m}.$$
 (5.214)

When acting on the GKZ integral \mathcal{I}_{γ} , we can now use the Euler equations (5.167) and DWI (5.168) to rewrite this expression in terms of elementary symmetric polynomials of just the parameters γ alone, namely

$$x_1 Q_1 \Big|_{\theta_1 \to \theta_1 + 1} \mathcal{I}_{\gamma} = x_1 \sum_{m=0}^7 \sigma_{(m)}(r) (-1 - \theta_1)^{7-m} \mathcal{I}_{\gamma}, \qquad (5.215)$$

where the eight variables

$$r_{\{m\}} = 1 - \gamma_0 + \gamma_1 \pm \gamma_2 \pm \gamma_3 \pm \gamma_4 \tag{5.216}$$

are formed by making all possible independent choices of \pm signs.

We now turn to the remaining Q_k coefficients in (5.211) for k = 2, 3, 4. Defining the auxiliary functions

$$S(\theta) = (\theta + \theta_3 + \theta_4)(\theta + \bar{\theta}_3 + \theta_4)(\theta + \theta_3 + \bar{\theta}_4)(\theta + \bar{\theta}_3 + \bar{\theta}_4), \qquad (5.217)$$

$$T(\theta) = \theta^{-1}(S(\theta) - S(0)) = (\theta + u_3 + u_4)((\theta + u_3)(\theta + u_4) + 2v_3 + 2v_4), \quad (5.218)$$

where

$$u_k = \theta_k + \bar{\theta}_k, \qquad v_k = \theta_k \bar{\theta}_k, \qquad k = 3,4 \tag{5.219}$$

we can decompose

$$Q_2 = T(\theta_2)T(\theta_2), (5.220)$$

$$Q_3 = ((u_3 + u_4)u_4 + v_3 - v_4)(S(\theta_2) + S(\bar{\theta}_2) - S(0)), \qquad (5.221)$$

$$Q_4 = ((u_3 + u_4)u_3 + v_4 - v_3)(S(\theta_2) + S(\bar{\theta}_2) - S(0)).$$
(5.222)

Noting that for k = 3, 4,

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$$(S(\theta_2) + S(\bar{\theta}_2) - S(0))\Big|_{\theta_k \to \theta_k + 1, \,\bar{\theta}_k \to \bar{\theta}_k + 1} = (S(\theta_2 + 1) + S(\bar{\theta}_2 + 1) - S(1)), \quad (5.223)$$

and using (5.191), the creation operator is then

$$\mathcal{C}_{1}^{\text{ph}} = x_{1} \sum_{m=0}^{\prime} \sigma_{(m)}(r)(-1-\theta_{1})^{7-m} + (\theta_{1}-\gamma_{1}) \Big[x_{2} \hat{T}(\theta_{2}+1) \hat{T}(\theta_{2}-\gamma_{2}+1)$$
(5.224)

$$+\left(x_3\left((2+\hat{u}_3+\hat{u}_4)\hat{u}_4+\hat{u}_3+1+\hat{v}_3-\hat{v}_4\right)+x_4\left((2+\hat{u}_3+\hat{u}_4)\hat{u}_3+\hat{u}_4+1+\hat{v}_4-\hat{v}_3\right)\right)$$

$$\times \left(\hat{S}(\theta_2 + 1) + \hat{S}(\theta_2 - \gamma_2 + 1) - \hat{S}(1) \right) \right]$$

where all hatted quantities are defined by replacing $\bar{\theta}_k \to \theta_k - \gamma_k$ for k = 3, 4 in the corresponding unhatted quantities. Its action is

$$C_1 \mathcal{I}_{\{\gamma_0,\gamma_1\}} = -(\gamma_0 - 1)^{-1} b_1 (\gamma_0 - 1, \gamma_1 + 1) \mathcal{I}_{\{\gamma_0 - 1, \gamma_1 + 1\}}$$
(5.225)

where, using $b_1(\boldsymbol{\gamma})$ as given in (5.209),

$$b_{1}(\gamma_{0} - 1, \gamma_{1} + 1) = b_{1}(\boldsymbol{\gamma})\Big|_{\gamma_{0} \to \gamma_{0} - 1, \gamma_{1} \to \gamma_{1} + 1} \\ = \frac{1}{256} \Big[\Big((2 - \gamma_{0} + \gamma_{1})^{2} - \mathcal{S}_{(1)} \Big)^{4} - 8 \Big((2 - \gamma_{0} + \gamma_{1})^{2} - \mathcal{S}_{(1)} \Big)^{2} \mathcal{S}_{(2)} \\ + 16 \mathcal{S}_{(2)}^{2} - 64 (2 - \gamma_{0} + \gamma_{1})^{2} \mathcal{S}_{(3)} \Big]$$
(5.226)

with the $S_{(m)}$ being elementary symmetric polynomials in γ_2^2 , γ_3^2 and γ_4^2 .

5.5.4 Examples

Taking into account the additional gamma function factors in (5.162), the action of these creation operators on contact diagrams is

$$C_1 i_{[d; \Delta_1, ..., \Delta_n]} = -4\gamma_1 b_1 (\gamma_0 - 1, \gamma_1 + 1) i_{[\tilde{d}; \tilde{\Delta}_1, \Delta_2 ..., \Delta_n]}$$
(5.227)

where

$$\tilde{d} = d - \frac{2}{n-2}, \qquad \tilde{\Delta}_1 = \Delta_1 + \frac{n-3}{n-2},$$
(5.228)

Alternatively, in terms of the multiple-Bessel integral

$$I_{\gamma_0\{\gamma_1,...,\gamma_n\}} = \int_0^\infty dz \, z^{\gamma_0} \prod_{i=1}^n p_i^{2\gamma_i} K_{\gamma_i}(p_i z), \qquad (5.229)$$

from (5.156) and (5.162) we have

$$\mathcal{I}_{\gamma} = \frac{2^{n-\gamma_0}}{\Gamma(\gamma_0)} I_{\gamma_0 \{\gamma_1, \dots, \gamma_n\}}$$
(5.230)

and hence

$$\mathcal{C}_{1}I_{\gamma_{0}\{\gamma_{1},\gamma_{2},...,\gamma_{n}\}} = -2b_{1}(\gamma_{0}-1,\gamma_{1}+1)I_{\gamma_{0}-1\{\gamma_{1}+1,\gamma_{2},...,\gamma_{n}\}}.$$
(5.231)

Here we can either use (5.193) to rewrite C_1 in terms of the momenta p_i , or more easily, re-express (5.229) using $p_i = \sqrt{x_i}$ then convert back to p_i after acting with C_1 .

A quick check of these results can be obtained by examining cases where all the Bessel indices γ_i take half-integer values allowing direct evaluation of the contact diagrams. (We restrict to cases where both the initial and the shifted integral are finite; for the analysis

of renormalised cases see [36].) For example, at three points, the triple-K integrals (5.229)

$$I_{4\{\frac{1}{2}\frac{1}{2}\frac{1}{2}\}} = \frac{15\pi^2}{16\sqrt{2}} \left(p_1 + p_2 + p_3\right)^{-7/2}, \qquad I_{3\{\frac{3}{2}\frac{1}{2}\frac{1}{2}\}} = \frac{\pi^2(5p_1 + 2p_2 + 2p_3)}{8\sqrt{2}(p_1 + p_2 + p_3)^{5/2}}, \qquad (5.232)$$

and one can verify that

$$\mathcal{C}_1 I_{4\{\frac{1}{2}\frac{1}{2}\frac{1}{2}\}} = -\frac{45}{128} I_{3\{\frac{3}{2}\frac{1}{2}\frac{1}{2}\}}, \tag{5.233}$$

consistent with (5.231) using (5.208) for the 3-point *b*-function. We have performed many similar checks at both 3- and 4-points.

More non-trivially, many triple-K integrals with integer indices can be evaluated [85] by acting with the annihilators \mathcal{L}_i and \mathcal{R}_i given in (5.124) on the known 'seed' integral $I_{1\{000\}}$ which can be evaluated in terms of the Bloch-Wigner dilogarithm. These relations enable computation of all the necessary triple-K integrals arising in 3-point functions of conserved currents and stress tensors in even spacetime dimensions [34, 35]. Since the creation operators \mathcal{C}_i and $\overline{\mathcal{C}}_i$ are the inverse of \mathcal{L}_i and \mathcal{R}_i , this allows us to reverse the direction of all operations linking different triple-K integrals within the reduction scheme. Thus, for example, we find

$$\mathcal{R}_1 I_{1\{000\}} = I_{2\{100\}}, \qquad -8C_1 I_{2\{100\}} = I_{1\{000\}}$$
(5.234)

where the integrals

$$I_{1\{000\}} = \frac{1}{2p_3^2(z-\bar{z})} \Big[\operatorname{Li}_2 z - \operatorname{Li}_2 \bar{z} + \frac{1}{2} \ln(z\bar{z}) \ln\left(\frac{1-z}{1-\bar{z}}\right) \Big],$$
(5.235)

$$I_{2\{100\}} = \frac{1}{2p_3^2(z-\bar{z})^2} \left[4p_3^2 z\bar{z}(-2+z+\bar{z})I_{1\{000\}} - 2z\bar{z}\ln(z\bar{z}) - (z+\bar{z}-2z\bar{z})\ln[(1-z)(1-\bar{z})] \right]$$
(5.236)

and the variables

$$z = \frac{1}{2p_3^2} \left(p_1^2 - p_2^2 + p_3^2 + \sqrt{-J^2} \right), \qquad \bar{z} = \frac{1}{2p_3^2} \left(p_1^2 - p_2^2 + p_3^2 - \sqrt{-J^2} \right)$$
(5.237)

or equivalently

$$\frac{p_1^2}{p_3^2} = z\bar{z}, \qquad \frac{p_2^2}{p_3^2} = (1-z)(1-\bar{z}) \tag{5.238}$$

with

$$J^{2} = (p_{1} + p_{2} + p_{3})(-p_{1} + p_{2} + p_{3})(p_{1} - p_{2} + p_{3})(p_{1} + p_{2} - p_{3})$$

= $-p_{1}^{4} - p_{2}^{4} - p_{3}^{4} + 2p_{1}^{2}p_{2}^{2} + 2p_{2}^{2}p_{3}^{2} + 2p_{3}^{2}p_{1}^{2}.$ (5.239)

5.5.5 Shift operators preserving the spacetime dimension

The creation operators constructed above decrease the spacetime dimension according to (5.228). For many applications, we would prefer an operator capable of changing the operator dimensions of a contact diagram while preserving the spacetime dimension. Thus,

we seek an operator $W_{12}^{\sigma_1,\sigma_2}$ such that

$$W_{12}^{\sigma_1,\sigma_2}i_{[d;\Delta_1,\Delta_2,\Delta_3,\dots,\Delta_n]} \propto i_{[d;\Delta_1+\sigma_1,\Delta_2+\sigma_2,\Delta_3,\dots,\Delta_n]}$$
(5.240)

for any independent choice of signs $\sigma_1 = \pm 1$ and $\sigma_2 = \pm 1$. Operators of this type are known at three points [83, 30], but their analogue at four points acts on contact diagrams to generate shifted contact diagrams with derivative vertices [36]. Instead, our discussion of creation operators above can be modified to enable operators of this type to be identified.¹² At three points we will see these coincide with the operators of [83, 30], but at four points and above they are novel. Using these operators will then enable further new shift operators to be constructed for exchange diagrams.

Our starting point is the observation that, for the GKZ integral (5.161) corresponding to the contact diagram,

$$W_{12}^{--}\bar{\partial}_1 \mathcal{I}_{\gamma} = b_W(\gamma)\partial_2 \mathcal{I}_{\gamma}.$$
(5.241)

Recalling the parameter identifications (5.163), the action of the operators here is

$$\begin{split} W_{12}^{--} : & \gamma_0 \to \gamma_0, & \gamma_1 \to \gamma_1 - 1, & \gamma_2 \to \gamma_2 - 1, \\ \bar{\partial}_1 : & \gamma_0 \to \gamma_0 + 1, & \gamma_1 \to \gamma_1 + 1, & \gamma_2 \to \gamma_2, \\ \partial_2 : & \gamma_0 \to \gamma_0 + 1, & \gamma_1 \to \gamma_1, & \gamma_2 \to \gamma_2 - 1, \end{split}$$
 (5.242)

with all remaining γ_k for $k = 3, \ldots, n$ staying the same. As the shifts produced by the operators on each side of (5.241) are the same, both sides involve the same integral \mathcal{I}_{γ} . As previously, the *b*-function $b_W(\gamma)$ should be a product of linear factors that vanishes whenever W_{12}^{--} maps us from a finite to a singular integral. Taking into account the action (5.174) of the annihilators in (5.241), we have

$$W_{12}^{--}\bar{\partial}_{1}\mathcal{I}_{\gamma} = -\gamma_{0}W_{12}^{--}\mathcal{I}_{\gamma}\Big|_{\gamma_{0}\to\gamma_{0}+1,\,\gamma_{1}\to\gamma_{1}+1} = -\gamma_{0}b_{W}(\gamma)\mathcal{I}_{\gamma}\Big|_{\gamma_{0}\to\gamma_{0}+1,\,\gamma_{2}\to\gamma_{2}-1} = b_{W}(\gamma)\partial_{2}\mathcal{I}_{\gamma}$$
(5.243)

and so the zeros of $b_W(\gamma)$ must cancel the singularities of $\mathcal{I}_{\gamma}|_{\gamma_0 \to \gamma_0+1, \gamma_2 \to \gamma_2-1}$. From (5.179), this means

$$b_{W}(\boldsymbol{\gamma}) = \prod_{\sigma_{k} \in \pm 1} \frac{1}{2} \Big(-(\gamma_{0}+1) - \gamma_{1} - (\gamma_{2}-1) + \sigma_{3}\gamma_{3} + \dots \sigma_{n}\gamma_{n} \Big)$$

$$= \prod_{\sigma_{k} \in \pm 1} \frac{1}{2} \Big(-\gamma_{0} - \gamma_{1} - \gamma_{2} + \sigma_{3}\gamma_{3} + \dots \sigma_{n}\gamma_{n} \Big).$$
(5.244)

Only the singularities with $\sigma_1 = \sigma_2 = -1$ in (5.179) appear here since these are the only cases for which $\mathcal{I}_{\gamma}|_{\gamma_0 \to \gamma_0+1, \gamma_2 \to \gamma_2-1}$ is singular but the integral $\mathcal{I}_{\gamma}|_{\gamma_0 \to \gamma_0+1, \gamma_1 \to \gamma_1+1}$ on which W_{12}^{--} acts is finite. Every possible independent choice of $\sigma_k \in \pm 1$ for all $k = 3, \ldots, n$

¹²The shift operators that we identify will moreover be of minimal order, unlike the *d*-preserving combination of an annihilator ∂_i or $\bar{\partial}_i$ followed by a creation operator C_j or \bar{C}_j . For example, the combination $\bar{C}_1 \partial_2 - \bar{C}_2 \partial_1$ produces the same shift as W_{12}^{--} but is of seventh order in derivatives for the 4-point function, since each product is eighth order and taking the difference lowers the order by one. In contrast, the 4-point operator W_{12}^{--} we find will be of only fourth order.

is permitted, however, and gives rise to a corresponding factor in (5.244). Once again, we have also chosen to include trivial factors of one-half in $b_W(\gamma)$ to simplify the subsequent form of W_{12}^{--} . Replacing the parameters γ in $b_W(\gamma)$ using the Euler equations (5.167) and DWI (5.168), we find

$$W_{12}^{--}\bar{\partial}_1 \mathcal{I}_{\gamma} = \partial_2 \big(b_W(\gamma) \mathcal{I}_{\gamma} \big) = \partial_2 B_W(\theta, \bar{\theta}) \mathcal{I}_{\gamma}$$
(5.245)

where

$$B_{W}(\theta,\bar{\theta}) = \prod_{\sigma_{k}\in\pm1} \frac{1}{2} \Big(\sum_{j=1}^{n} (\theta_{j} + \bar{\theta}_{j}) - (\theta_{1} - \bar{\theta}_{1}) - (\theta_{2} - \bar{\theta}_{2}) + \sigma_{3}(\theta_{3} - \bar{\theta}_{3}) + \dots + \sigma_{n}(\theta_{n} - \bar{\theta}_{n}) \Big)$$

$$= \prod_{\sigma_{k}\in\pm1} \Big(\bar{\theta}_{1} + \bar{\theta}_{2} + (\delta_{\sigma_{3},+1}\theta_{3} + \delta_{\sigma_{3},-1}\bar{\theta}_{3}) + \dots + (\delta_{\sigma_{n},+1}\theta_{n} + \delta_{\sigma_{n},-1}\bar{\theta}_{n}) \Big).$$

(5.246)

Since $B_W(\theta, \bar{\theta})$ is in fact independent of θ_2 the ordering of ∂_2 and $B_W(\theta, \bar{\theta})$ on the righthand side of (5.245) is in fact immaterial, but had this not been the case the ordering shown would be the correct one when using the unshifted Euler equations and DWI to replace the γ parameters.

To identify W_{12}^{--} , all that is then needed is to start with $\partial_2 B_W(\theta, \bar{\theta})$ and, using the toric equations (5.169), pull out a right factor of $\bar{\partial}_1$ according to (5.245). As usual, the resulting operator can then be projected down to the physical hypersurface using the Euler equations and DWI. These procedures are illustrated for the 3- and 4-point function below. Finally, given W_{12}^{--} in physical variables, all the remaining operators in (5.240) can be found by shadow conjugation using (5.173), namely

$$(W_{12}^{+-})_{\rm ph} = x_1^{1+\gamma_1} (W_{12}^{--})_{\rm ph} x_1^{-\gamma_1}, \qquad (5.247)$$

$$(W_{12}^{-+})_{\rm ph} = x_2^{1+\gamma_2} (W_{12}^{--})_{\rm ph} x_2^{-\gamma_2}, \qquad (5.248)$$

$$(W_{12}^{++})_{\rm ph} = x_1^{1+\gamma_1} x_2^{1+\gamma_2} (W_{12}^{--})_{\rm ph} x_1^{-\gamma_1} x_2^{-\gamma_2}.$$
(5.249)

3-point function

To illustrate the above discussion, for the 3-point function we have

$$b_W(\gamma) = \frac{1}{4}(-\gamma_0 - \gamma_1 - \gamma_2 + \gamma_3)(-\gamma_0 - \gamma_1 - \gamma_2 - \gamma_3)$$
(5.250)

and

$$B_W(\theta,\bar{\theta}) = (\bar{\theta}_1 + \bar{\theta}_2 + \theta_3)(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3).$$
(5.251)

The operator W_{12}^{--} can now be extracted from

$$W_{12}^{--}\bar{\partial}_1 \mathcal{I}_{\gamma} = \partial_2 B_W(\theta, \bar{\theta}) \mathcal{I}_{\gamma}.$$
(5.252)

For this, we write

$$\begin{aligned} \partial_2(\bar{\theta}_1 + \bar{\theta}_2 + \theta_3)(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3)\mathcal{I}_{\boldsymbol{\gamma}} \\ &= \partial_2 \big[(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \theta_3)(\bar{\theta}_1 + \bar{\theta}_2) + \theta_3\bar{\theta}_3] \mathcal{I}_{\boldsymbol{\gamma}} \end{aligned}$$
$$= \left[(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \theta_3)(\bar{x}_1 \partial_2 \bar{\partial}_1 + \bar{x}_2 \partial_2 \bar{\partial}_2) + x_3 \bar{x}_3 \partial_2 \partial_3 \bar{\partial}_3 \right] \mathcal{I}_{\boldsymbol{\gamma}}$$

$$= \left[(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \theta_3)(\bar{x}_1 \partial_2 + \bar{x}_2 \partial_1) + x_3 \bar{x}_3 \partial_2 \partial_1 \right] \bar{\partial}_1 \mathcal{I}_{\boldsymbol{\gamma}}$$
(5.253)

where in the penultimate line we used the toric equations (5.169). Thus

$$W_{12}^{--} = (\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \theta_3)(\bar{x}_1\partial_2 + \bar{x}_2\partial_1) + x_3\bar{x}_3\partial_2\partial_1$$

= $(\bar{x}_1\partial_2 + \bar{x}_2\partial_1)(1 + \bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \theta_3) + x_3\bar{x}_3\partial_2\partial_1,$ (5.254)

and using the DWI (5.168) to project to the physical hypersurface (5.164), we obtain

$$(W_{12}^{--})_{\rm ph} = (\partial_2 + \partial_1)(1 - \gamma_0 - \theta_1 - \theta_2) + x_3 \partial_2 \partial_1$$

= $-(\gamma_0 + \theta_1 + \theta_2)(\partial_1 + \partial_2) + x_3 \partial_1 \partial_2$ (5.255)

where for the 3-point function $\gamma_0 = d/2$ from (5.163). A short calculation shows that

$$(W_{12}^{--})_{\rm ph} = -\frac{1}{4} \left(\partial_{p_1}^2 + \partial_{p_2}^2 + \frac{(d-1)}{p_1} \partial_{p_1} + \frac{(d-1)}{p_2} \partial_{p_2} + (p_1^2 + p_2^2 - p_3^2) \frac{1}{p_1 p_2} \partial_{p_1} \partial_{p_2} \right)$$
(5.256)

which, up to a factor of -2, is the 3-point shift operator studied in [30, 36].

The action of W_{12}^{--} is

$$W_{12}^{--}\mathcal{I}_{\gamma} = b_W(\gamma)\Big|_{\gamma_0 \to \gamma_0 - 1, \, \gamma_1 \to \gamma_1 - 1} \mathcal{I}_{\gamma}\Big|_{\gamma_1 \to \gamma_1 - 1, \, \gamma_2 \to \gamma_2 - 1}$$
(5.257)

where the shift on the *b*-function derives from the fact that, in the projection step going from (5.254) to (5.255), we have chosen that W_{12}^{--} acts on the integral \mathcal{I}_{γ} requiring us to shift the γ parameters present in (5.243). Evaluating, this gives

$$b_W(\boldsymbol{\gamma})\Big|_{\gamma_0 \to \gamma_0 - 1, \, \gamma_1 \to \gamma_1 - 1} = \frac{1}{4}(2 - \gamma_0 - \gamma_1 - \gamma_2 + \gamma_3)(2 - \gamma_0 - \gamma_1 - \gamma_2 - \gamma_3)$$
$$= \frac{1}{4}\Big((\gamma_0 + \gamma_1 + \gamma_2 - 2)^2 - \gamma_3^2\Big)$$
(5.258)

such that (5.257) is consistent with the action of \mathcal{W}_{12}^{--} obtained in [36]. Acting on the holographic contact diagram, from (5.162) we have

$$W_{12}^{--i}i_{[d,\,\Delta_1,\,\Delta_2,\,\Delta_3]} = \frac{1}{4(\gamma_1 - 1)(\gamma_2 - 1)} b_W(\boldsymbol{\gamma}) \Big|_{\gamma_0 \to \gamma_0 - 1,\,\gamma_1 \to \gamma_1 - 1} i_{[d,\,\Delta_1 - 1,\,\Delta_2 - 1,\,\Delta_3]}.$$
 (5.259)

4-point function

At 4-points, we find

$$b_W(\gamma) = \frac{1}{16} (-\gamma_0 - \gamma_1 - \gamma_2 + \gamma_3 + \gamma_4) (-\gamma_0 - \gamma_1 - \gamma_2 - \gamma_3 + \gamma_4) \\ \times (-\gamma_0 - \gamma_1 - \gamma_2 + \gamma_3 - \gamma_4) (-\gamma_0 - \gamma_1 - \gamma_2 - \gamma_3 - \gamma_4)$$
(5.260)

and hence

$$B_W(\theta,\bar{\theta}) = (\bar{\theta}_1 + \bar{\theta}_2 + \theta_3 + \theta_4)(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \theta_4)(\bar{\theta}_1 + \bar{\theta}_2 + \theta_3 + \bar{\theta}_4)(\bar{\theta}_1 + \bar{\theta}_2 + \bar{\theta}_3 + \bar{\theta}_4).$$
(5.261)

Once again, to find W^{--}_{12} we must factorise

$$W_{12}^{--}\bar{\partial}_1 \mathcal{I}_{\gamma} = \partial_2 B_W(\theta, \bar{\theta}) \mathcal{I}_{\gamma}.$$
(5.262)

As a first step, we expand

$$B_W(\theta,\bar{\theta}) = Q_0(\bar{\theta}_1 + \bar{\theta}_2) + Q_3\theta_3\bar{\theta}_3 + Q_4\theta_4\bar{\theta}_4$$
(5.263)

where the coefficients

$$Q_{0} = (u_{3} + u_{4} + \bar{\theta}_{1} + \bar{\theta}_{2}) \Big(2(v_{3} + v_{4}) + (u_{3} + \bar{\theta}_{1} + \bar{\theta}_{2})(u_{4} + \bar{\theta}_{1} + \bar{\theta}_{2}) \Big),$$

$$Q_{3} = (u_{3} + u_{4})u_{4} + v_{3} - v_{4},$$

$$Q_{4} = (u_{3} + u_{4})u_{3} - v_{3} + v_{4},$$

(5.264)

and

$$u_k = \theta_k + \bar{\theta}_k, \qquad v_k = \theta_k \bar{\theta}_k, \qquad k = 3, 4.$$
(5.265)

Now, since all coefficients are independent of θ_2 ,

$$\partial_{2}B_{W}(\theta,\bar{\theta})\mathcal{I}_{\gamma} = \left[Q_{0}(\bar{x}_{1}\partial_{2}\bar{\partial}_{1} + \bar{x}_{2}\partial_{2}\bar{\partial}_{2}) + Q_{3}x_{3}\bar{x}_{3}\partial_{2}\bar{\partial}_{3}\partial_{3} + Q_{4}x_{4}\bar{x}_{4}\partial_{2}\bar{\partial}_{4}\partial_{4} \right]\mathcal{I}_{\gamma}$$
(5.266)
$$= \left[\bar{x}_{1}Q_{0} \Big|_{\bar{\theta}_{1} \to \bar{\theta}_{1}+1} \partial_{2} + \bar{x}_{2}Q_{0} \Big|_{\bar{\theta}_{2} \to \bar{\theta}_{2}+1} \partial_{1} + x_{3}\bar{x}_{3}Q_{3} \Big|_{\theta_{3} \to \theta_{3}+1, \bar{\theta}_{3} \to \bar{\theta}_{3}+1} \partial_{2}\partial_{1} + x_{4}\bar{x}_{4}Q_{4} \Big|_{\theta_{4} \to \theta_{4}+1, \bar{\theta}_{4} \to \bar{\theta}_{4}+1} \partial_{2}\partial_{1} \right] \bar{\partial}_{1}\mathcal{I}_{\gamma}$$

where in the second line we used the toric equations (5.169). We thus have

$$W_{12}^{--} = (\bar{x}_1 \partial_2 + \bar{x}_2 \partial_1) Q_0 \Big|_{\bar{\theta}_1 \to \bar{\theta}_1 + 1} \\ + \partial_1 \partial_2 \Big(x_3 \bar{x}_3 Q_3 \Big|_{\theta_3 \to \theta_3 + 1, \bar{\theta}_3 \to \bar{\theta}_3 + 1} + x_4 \bar{x}_4 Q_4 \Big|_{\theta_4 \to \theta_4 + 1, \bar{\theta}_4 \to \bar{\theta}_4 + 1} \Big)$$
(5.267)

where in the first line we used the fact that $\bar{\theta}_1$ and $\bar{\theta}_2$ enter Q_0 only in the combination $\bar{\theta}_1 + \bar{\theta}_2$ and so the replacement $\bar{\theta}_1 \to \bar{\theta}_1 + 1$ produces the same result as $\bar{\theta}_2 \to \bar{\theta}_2 + 1$ allowing us to combine the two Q_0 terms. We have in addition moved Q_0, Q_3 and Q_4 to the right (noting that all coefficients are independent of θ_1 and θ_2) so as to be able to use the Euler equations for \mathcal{I}_{γ} to project to the physical hypersurface. For this, we set $\bar{x}_k \to 1$ and $\bar{\theta}_k \to \theta_k - \gamma_k$ inside all Q_k coefficients giving

$$(W_{12}^{--})_{\rm ph} = (\partial_1 + \partial_2)Q_0\Big|_{\bar{\theta}_k \to \theta_k - \gamma_k + \delta_{k,1}}$$

$$+ \partial_1 \partial_2 \Big(x_3 Q_3 \Big|_{\theta_3 \to \theta_3 + 1, \bar{\theta}_k \to \theta_k - \gamma_k + \delta_{k,3}} + x_4 Q_4 \Big|_{\theta_4 \to \theta_4 + 1, \bar{\theta}_k \to \theta_k - \gamma_k + \delta_{k,4}} \Big).$$
(5.268)

In all the replacements here, $\bar{\theta}_k$ stands for any index k = 1, ..., 4. Evaluating this formula explicitly using the coefficients in (5.264), we find

$$(W_{12}^{--})_{\rm ph} = (\partial_1 + \partial_2)(1 - \gamma_0 - \theta_1 - \theta_2) \Big(2\theta_3(\theta_3 - \gamma_3) + 2\theta_4(\theta_4 - \gamma_4)$$

$$+ (1 + 2\theta_3 - \gamma_3 + \theta_1 - \gamma_1 + \theta_2 - \gamma_2)(1 + 2\theta_4 - \gamma_4 + \theta_1 - \gamma_1 + \theta_2 - \gamma_2) \Big)$$

$$+ \partial_1 \partial_2 \Big[x_3 \Big((2 + 2\theta_3 - \gamma_3 + 2\theta_4 - \gamma_4)(2\theta_4 - \gamma_4) + (1 + \theta_3)(1 + \theta_3 - \gamma_3) - \theta_4(\theta_4 - \gamma_4) \Big)$$

$$+ x_4 \Big((2 + 2\theta_3 - \gamma_3 + 2\theta_4 - \gamma_4)(2\theta_3 - \gamma_3) + (1 + \theta_4)(1 + \theta_4 - \gamma_4) - \theta_3(\theta_3 - \gamma_3) \Big) \Big]$$
(5.269)

where for the 4-point function $\gamma_0 = d$ from (5.163).

Alternatively, we can use the DWI (5.168) projected to the physical hypersurface,

$$0 = \left(\frac{1}{2}(\gamma_0 - \gamma_t) + \sum_{k=1}^4 \theta_k\right) \mathcal{I}_{\gamma}, \qquad \gamma_t = \sum_{k=1}^4 \gamma_k, \tag{5.270}$$

to eliminate the factors of $\theta_1 + \theta_2$ on the second line of (5.269). After further moving all factors of ∂_1 and ∂_2 to the right, this gives the equivalent form

$$(W_{12}^{--})_{\rm ph} = -(\gamma_0 + \theta_1 + \theta_2) \Big((\theta_3 + \theta_4)(\theta_3 + \theta_4 - \gamma_3 - \gamma_4) + \frac{1}{4}(2 - \gamma_0 - \gamma_t + 2\gamma_3)(2 - \gamma_0 - \gamma_t + 2\gamma_4) \Big) (\partial_1 + \partial_2) + \Big[x_3 \Big((2 + 2\theta_3 - \gamma_3 + 2\theta_4 - \gamma_4)(2\theta_4 - \gamma_4) + (1 + \theta_3)(1 + \theta_3 - \gamma_3) - \theta_4(\theta_4 - \gamma_4) \Big) + x_4 \Big((2 + 2\theta_3 - \gamma_3 + 2\theta_4 - \gamma_4)(2\theta_3 - \gamma_3) + (1 + \theta_4)(1 + \theta_4 - \gamma_4) - \theta_3(\theta_3 - \gamma_3) \Big) \Big] \partial_1 \partial_2$$

$$(5.271)$$

The action of W_{12}^{--} is

$$W_{12}^{--}\mathcal{I}_{\gamma} = b_W(\gamma)\Big|_{\gamma_0 \to \gamma_0 - 1, \, \gamma_1 \to \gamma_1 - 1} \mathcal{I}_{\gamma}\Big|_{\gamma_1 \to \gamma_1 - 1, \, \gamma_2 \to \gamma_2 - 1}$$
(5.272)

where, once again, the shift on the *b*-function derives from the fact that in projecting from GKZ variables to the physical hypersurface we chose W_{12}^{--} to act on the unshifted integral \mathcal{I}_{γ} requiring us to shift the γ parameters present in (5.243). Explicitly, this is

$$b_W(\boldsymbol{\gamma})\Big|_{\gamma_0 \to \gamma_0 - 1, \, \gamma_1 \to \gamma_1 - 1} = \frac{1}{16} \Big(\gamma_3^2 + \gamma_4^2 - (\gamma_0 + \gamma_1 + \gamma_2 - 2)^2\Big)^2 - \frac{1}{4}\gamma_3^2\gamma_4^2.$$
(5.273)

Acting on the holographic contact diagram, from (5.162) we again have

$$W_{12}^{--}i_{[d;\,\Delta_1,\Delta_2,\Delta_3,\Delta_4]} = \frac{1}{4(\gamma_1 - 1)(\gamma_2 - 1)} b_W(\boldsymbol{\gamma}) \Big|_{\gamma_0 \to \gamma_0 - 1,\,\gamma_1 \to \gamma_1 - 1} i_{[d;\,\Delta_1 - 1,\Delta_2 - 1,\Delta_3,\Delta_4]}.$$
(5.274)

To our knowledge, this is the first time an operator that shifts the 4-point contact diagram in this fashion has been identified. We emphasise that the 3-point operator (5.256), when applied to 4-point contact diagrams, generates shifted contact diagrams but with derivative vertices and hence does not satisfy this requirement [36].

Examples: Contact diagrams for which the Bessel functions have half-integer indices can be evaluated directly. This yields many simple examples for which the action of W_{12}^{--} can be checked. For instance, with $(d, \Delta_1, \Delta_2, \Delta_3, \Delta_4) = (5, 3, 4, 3, 4)$, we find

$$i_{[5;3,4,3,4]} = \frac{1}{p_t^3} \left(p_1^2 + p_3^2 + 2(p_2^2 + p_4^2) + 3(p_1 + p_3)(p_2 + p_4) + 2p_1p_3 + 6p_2p_4 \right)$$
(5.275)

where $p_t = \sum_{j=1}^{4} p_j$, while the shifted integral with $(d, \Delta_1, \Delta_2, \Delta_3, \Delta_4) = (5, 2, 3, 3, 4)$ is

$$i_{[5;2,3,3,4]} = -\frac{(p_t + 2p_4)}{p_1 p_t^3}.$$
(5.276)

Evaluating the action of W_{12}^{--} in (5.271) using (5.193), we can verify (5.274), namely

$$W_{12}^{--}i_{[5;3,4,3,4]} = -\frac{63}{2}i_{[5;2,3,3,4]}.$$
(5.277)

Combinations of operators

To round up our discussion of shift operators for contact diagrams, we have identified operators mapping

$$\begin{aligned} \partial_i : & \gamma_0 \to \gamma_0 + 1, \quad \gamma_i \to \gamma_i - 1 \qquad \partial_i : \quad \gamma_0 \to \gamma_0 + 1, \quad \gamma_i \to \gamma_i + 1 \\ \mathcal{C}_i : & \gamma_0 \to \gamma_0 - 1, \quad \gamma_i \to \gamma_i + 1 \qquad \bar{\mathcal{C}}_i : \quad \gamma_0 \to \gamma_0 - 1, \quad \gamma_i \to \gamma_i - 1 \\ W_{ij}^{\sigma_i \sigma_j} : & \gamma_i \to \gamma_i + \sigma_i, \quad \gamma_j \to \gamma_j + \sigma_j, \qquad \{\sigma_i, \sigma_j\} \in \pm 1. \end{aligned}$$

$$(5.278)$$

Combining these allows us to construct yet further shifts, for example:

$$C_i \bar{C}_i: \gamma_0 \to \gamma_0 - 2, \qquad C_i \bar{\partial}_i: \gamma_i \to \gamma_i + 2, \qquad \bar{C}_i \partial_i: \gamma_i \to \gamma_i - 2.$$
 (5.279)

Acting on the 3-point function specifically,

$$\mathcal{C}_1 W_{23}^{++}: \quad \gamma_0 \to \gamma_0 - 1, \quad \gamma_i \to \gamma_i + 1 \quad \forall \ i = 1, 2, 3 \tag{5.280}$$

which is equivalent to shifting $d \to d-2$ while preserving all operator dimensions Δ_i .

Finally, one might wonder why all these operators produce a shift of two units: why, for example, can one not construct an operator shifting $\gamma_0 \rightarrow \gamma_0 + 1$ only, or just $\gamma_1 \rightarrow \gamma_1 + 1$? The absence of such operators can be traced to the spacing of the singular hyperplanes of the contact diagram, specifically the term -2m appearing in the singularity condition (5.179). As $m \in \mathbb{Z}^+$, this means that the singularities are effectively spaced by two units. Any operator that produced a shift of a single unit would require a *b*-function containing an infinite number of factors, since there are infinitely many finite integrals that are only one unit away from a singular integral. (Namely, those for which *m* is half-integer.) As the number of factors in the *b*-function corresponds to the order of the differential operator, there is thus no single-shift operator of finite order. In contrast, for an operator shifting by two units, the number of finite integrals that can be mapped to singular integrals is finite, and hence the *b*-functions and shift operators are also of finite order.

5.5.6 Exchange diagrams

Having analysed contact diagrams, we now turn to the *s*-channel exchange diagrams (5.159). Rather than constructing an explicit GKZ representation, here we simply note that shifts of the form

$$i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} \to i_{[d;\Delta_1+\sigma_1,\Delta_2+\sigma_2;\Delta_3,\Delta_4;\Delta_x]}$$
(5.281)

for any $\{\sigma_1, \sigma_2\} \in \pm 1$ can be obtained by combining the 3- and 4-point W_{12}^{--} operators given in (5.255) and (5.271) with the *s*-channel Casimir operator. As with contact diagrams, it is sufficient to focus on the case $\sigma_1 = \sigma_2 = -1$, since all remaining operators follow by shadow conjugation according to (5.247). We emphasise however that both the original and the shifted exchange diagrams we consider have purely *non-derivative* vertices. Moreover, any operator and spacetime dimensions are permitted, provided we work in dimensional regularisation where necessary to avoid divergences.

For purposes of disambiguation, let us define the operator

$$\mathcal{W}_{12}^{--} = (d + 2\theta_1 + 2\theta_2)(\partial_1 + \partial_2) - 2s^2 \partial_1 \partial_2 \tag{5.282}$$

where d is the boundary spacetime dimension and $\partial_i = \partial/\partial x_i$ with $x_i = p_i^2$ as usual. This is simply the 3-point operator $-2W_{12}^{--}$ in (5.256), but with p_3^2 replaced by the Mandelstam variable $s^2 = (\mathbf{p}_1 + \mathbf{p}_2)^2$ as appropriate for acting on s-channel exchange diagrams. (The factor of -2 is included for consistency with the W_{12}^{--} defined in [30, 36].) In the following, we will then use W_{12}^{--} to refer exclusively to the 4-point W_{12}^{--} operator given in (5.271).

As shown in [36], the action of \mathcal{W}_{12}^{--} on an *s*-channel exchange diagram is to produce a *linear combination* of a shifted exchange and a shifted contact diagram:

$$\mathcal{W}_{12}^{--} i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = \mathcal{N}_{exch.} i_{[d;\Delta_1-1,\Delta_2-1;\Delta_3,\Delta_4;\Delta_x]} + \mathcal{N}_{cont.} i_{[d;\Delta_1-1,\Delta_2-1,\Delta_3,\Delta_4]}$$
(5.283)

where the coefficients¹³

$$\mathcal{N}_{exch.} = \left(\frac{d}{2} - 2 + \gamma_1 + \gamma_2 + \gamma_x\right) \left(\frac{d}{2} - 2 + \gamma_1 + \gamma_2 - \gamma_x\right) \mathcal{N}_{cont.}$$
(5.284)

$$\mathcal{N}_{cont.} = -\frac{1}{8(\gamma_1 - 1)(\gamma_2 - 1)} \tag{5.285}$$

where $\gamma_i = \Delta_i - d/2$ and $\gamma_x = \Delta_x - d/2$. Thus, in order to go from an exchange diagram to shifted exchange diagram only, the shifted contact contribution in (5.283) must be subtracted.

This can be accomplished in two steps. First, the *unshifted* contact diagram is obtained by acting on the original exchange diagram with the reduced Casimir operator,

$$\mathcal{C}_{12} i_{[d;\Delta_1,\Delta_2;\Delta_3,\Delta_4;\Delta_x]} = i_{[d;\Delta_1,\Delta_2,\Delta_3,\Delta_4]}, \qquad (5.286)$$

 $^{^{13}}$ Where the shifted exchange diagram has a pole (or double pole) in dimensional regularisation, one (or both) of the factors on the right-hand side of (5.284) vanish, see [36].

where

$$\hat{\mathcal{C}}_{12} = 2s^2 \Big((\theta_1 + 1 - \gamma_1)\partial_1 + (\theta_2 + 1 - \gamma_2)\partial_2 \Big) - \Big(2\theta_1 + 2\theta_2 - \gamma_1 - \gamma_2 + \frac{d}{2}\Big)^2 + \gamma_x^2$$
(5.287)

with $\theta_i = x_i \partial_i$. The action of this operator on an *s*-channel exchange is equivalent to that of the Casimir operator plus the square of the exchanged mass [36].¹⁴ If desired, $\hat{\mathcal{C}}_{12}$ can be shorted using the identity

$$0 = \left((\theta_1 + 1 - \gamma_1) \partial_1 - (\theta_2 + 1 - \gamma_2) \partial_2 \right) i_{[d; \,\Delta_1, \Delta_2, \Delta_3, \Delta_4 \, x \, \Delta_x]}$$
(5.288)

which corresponds to the difference of the Bessel operators acting on legs 1 and 2, *i.e.*, $K_1 - K_2$ where $K_i = \partial_{p_i}^2 + (1 - 2\gamma_i)p_i^{-1}\partial_{p_i}$. However, (5.287) is symmetric under $1 \leftrightarrow 2$.

For the second step, we now construct the shifted contact diagram using the 4-point W_{12}^{--} operator defined in (5.271). From (5.274), this has the action

$$W_{12}^{--}i_{[d;\,\Delta_1,\Delta_2,\Delta_3,\Delta_4]} = \mathcal{N}_W \,\mathcal{N}_{cont.}\,i_{[d;\,\Delta_1-1,\Delta_2-1,\Delta_3,\Delta_4]}$$
(5.289)

with $\mathcal{N}_{cont.}$ from (5.285) and

$$\mathcal{N}_W = -\frac{1}{8} \Big[\left(\gamma_3^2 + \gamma_4^2 - (d + \gamma_1 + \gamma_2 - 2)^2 \right)^2 - 4\gamma_3^2 \gamma_4^2 \Big].$$
(5.290)

Putting everything together, we find the operator

$$\Omega_{12}^{--} = \mathcal{N}_W \, \mathcal{W}_{12}^{--} - W_{12}^{--} \hat{\mathcal{C}}_{12} \tag{5.291}$$

whose action is

$$\Omega_{12}^{--}i_{[d;\,\Delta_1,\Delta_2;\,\Delta_3,\Delta_4;\,\Delta_x]} = \mathcal{N}_W \mathcal{N}_{exch.} i_{[d;\,\Delta_1-1,\Delta_2-1;\,\Delta_3,\Delta_4;\,\Delta_x]}.$$
(5.292)

This is therefore the desired operator mapping an exchange to a shifted exchange diagram. Written out explicitly, with $\gamma_t = \sum_{j=1}^4 \gamma_j$, we have

 $\Omega_{12}^{--} = -\frac{1}{8} \Big[\left(\gamma_3^2 + \gamma_4^2 - (d + \gamma_1 + \gamma_2 - 2)^2 \right)^2 - 4\gamma_3^2 \gamma_4^2 \Big] \Big((d + 2\theta_1 + 2\theta_2)(\partial_1 + \partial_2) - 2s^2 \partial_1 \partial_2 \Big) \\ - \Big[- (d + \theta_1 + \theta_2) \Big((\theta_3 + \theta_4)(\theta_3 + \theta_4 - \gamma_3 - \gamma_4) \\ + \frac{1}{4} (2 - d - \gamma_t + 2\gamma_3)(2 - d - \gamma_t + 2\gamma_4) \Big) (\partial_1 + \partial_2) \Big]$

$$+ x_{3} \Big((2 + 2\theta_{3} - \gamma_{3} + 2\theta_{4} - \gamma_{4}) (2\theta_{4} - \gamma_{4}) + (1 + \theta_{3}) (1 + \theta_{3} - \gamma_{3}) - \theta_{4} (\theta_{4} - \gamma_{4}) \Big) \partial_{1} \partial_{2} \\ + x_{4} \Big((2 + 2\theta_{3} - \gamma_{3} + 2\theta_{4} - \gamma_{4}) (2\theta_{3} - \gamma_{3}) + (1 + \theta_{4}) (1 + \theta_{4} - \gamma_{4}) - \theta_{3} (\theta_{3} - \gamma_{3}) \Big) \partial_{1} \partial_{2} \Big] \times \\ \times \Big[2s^{2} \Big((\theta_{1} + 1 - \gamma_{1}) \partial_{1} + (\theta_{2} + 1 - \gamma_{2}) \partial_{2} \Big) - \Big(2\theta_{1} + 2\theta_{2} - \gamma_{1} - \gamma_{2} + \frac{d}{2} \Big)^{2} + \gamma_{x}^{2} \Big].$$
(5.293)

¹⁴Specifically, $\hat{\mathcal{C}}_{12} = \tilde{\mathcal{C}}_{12} + m_x^2$ with $\tilde{\mathcal{C}}_{12}$ as defined in (6.44) of [36] and $m_x^2 = \gamma_x^2 - d^2/4$.

Examples: All exchange diagrams involving fields of $\Delta = 2, 3$ in d = 3 were computed recently in [36] and are available in the associated Mathematica package Handbook.wl. These results enable many tests of the operator Ω_{12}^{--} in (5.293) and its shadow conjugates

$$\Omega_{12}^{+-} = x_1^{1+\gamma_1} \Omega_{12}^{--} x_1^{-\gamma_1}, \tag{5.294}$$

$$\Omega_{12}^{-+} = x_2^{1+\gamma_2} \Omega_{12}^{--}, x_2^{-\gamma_2}, \tag{5.295}$$

$$\Omega_{12}^{++} = x_1^{1+\gamma_1} x_2^{1+\gamma_2} \Omega_{12}^{--} x_1^{-\gamma_1} x_2^{-\gamma_2}.$$
(5.296)

For this, we work in the dimensionally regulated theory with $d \to d + 2\varepsilon$ and $\Delta_i \to \Delta_i + \varepsilon$ for all i = 1, 2, 3, 4, x. This scheme has the virtue of preserving the half-integer values of all Bessel function indices $\gamma_i = \Delta_i - d/2$. The simplest such example is

$$i_{[3;22;22;2]} = -\frac{1}{2s} \mathcal{D}^{(+)},$$

$$i_{[3;33;22;2]} = \frac{1}{2} (p_3 + p_4) \Gamma(2\varepsilon) p_T^{-2\varepsilon} + \frac{1}{4s} (p_1^2 + p_2^2 - s^2) \mathcal{D}^{(+)} + \frac{1}{2} (p_1 + p_2) \Big[\log \Big(\frac{l_{34+}}{p_T}\Big) + 1 \Big] + \frac{7}{8} (p_3 + p_4) + O(\varepsilon)$$
(5.297)
(5.297)
(5.297)

where

$$p_T = \sum_{i=1}^{4} p_i, \qquad l_{ij\pm} = p_i + p_j \pm s,$$
 (5.299)

and

$$\mathcal{D}^{(+)} = \text{Li}_2\left(\frac{l_{12-}}{p_T}\right) + \text{Li}_2\left(\frac{l_{34-}}{p_T}\right) + \log\left(\frac{l_{12+}}{p_T}\right)\log\left(\frac{l_{34+}}{p_T}\right) - \frac{\pi^2}{6}.$$
 (5.300)

By direct differentiation, one then finds

$$\Omega_{12}^{--}i_{[3;33;22;2]} = (90 + 261\varepsilon + O(\varepsilon^2))i_{[3;22;22;2]} + O(\varepsilon)$$
(5.301)

consistent with (5.292). Note $\mathcal{N}_W \mathcal{N}_{exch}$ on the right-hand side here is expanded to order ε since $i_{[3:22:22:2]}$ has an ε^{-1} pole. We have performed similar checks for all other values of the Δ_i and Δ_x , and for the shadow conjugated operators.

This ability to shift exchange diagrams directly to other exchange diagrams means that, instead of computing all the diagrams individually, we can compute the easiest diagram (namely, $i_{[3;22;22;2]}$) to sufficiently high order in the regulator ε , and then obtain all others by acting with $\Omega_{12}^{\sigma_1 \sigma_2}$ and $\Omega_{34}^{\sigma_3 \sigma_4}$.

Creation operators for Feynman diagrams 5.6

In this section, we analyse various Feynman integrals presenting their GKZ representations, their singularities, and the associated creation operators. Many of the examples we study have appeared in the recent works [123, 124, 128, 129]. Here, our focus will be the construction of the creation operators and ways to automate this computation using standard Gröbner basis and convex hulling algorithms.



Figure 5.8: The single-mass bubble integral (5.305), with massless and massive propagators represented by dashed and undashed lines respectively.

In all cases, we start with an L-loop scalar integral in the momentum representation

$$I = \left(\prod_{j=1}^{L} \int \frac{\mathrm{d}^d \boldsymbol{k}_j}{(2\pi)^d}\right) \frac{1}{P_1^{\gamma_1} \dots P_N^{\gamma_N}},\tag{5.302}$$

where the propagators P_i for i = 1, ..., N are raised to generalised powers γ_i . As shown in appendix C.1 (see also [123, 65]), the corresponding GKZ integral is

$$\mathcal{I}_{\gamma} = \left(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1}\right) \mathcal{D}^{-\gamma_{0}}, \qquad \gamma_{0} = \frac{d}{2}$$
(5.303)

where the denominator \mathcal{D} is formed from the Lee-Pomeransky denominator $\mathcal{G} = \mathcal{U} + \mathcal{F}$, the sum of first and second Symanzik polynomials, by replacing the coefficient of every term with an independent variable x_k . The Feynman integral (5.302) now corresponds to

$$I = c_{\gamma} \mathcal{I}_{\gamma}, \qquad c_{\gamma} = \frac{(4\pi)^{-L\gamma_0} \Gamma(\gamma_0)}{\Gamma((L+1)\gamma_0 - \gamma_t) \prod_{i=1}^N \Gamma(\gamma_i)}, \qquad \gamma_t = \sum_{i=1}^N \gamma_i$$
(5.304)

with the x_k restored to their physical (Lee-Pomeransky) values. Knowing the coefficient c_{γ} enables the action of a creation operator on the GKZ integral \mathcal{I}_{γ} to be related to its action on the Feynman integral I.

5.6.1 Bubble diagram

First, we consider the 1-loop bubble integral with propagators of mass m_1 and m_2 . To warm-up, we begin with the single-mass case $(m_1, m_2) = (0, m)$ before turning to general masses. The fully massless case $m_1 = m_2 = 0$ is trivial (evaluating to a simple power of the momentum) and will be omitted.

1-mass bubble

The single-mass bubble diagram

$$I = \int \frac{\mathrm{d}^{d} \mathbf{k}}{(2\pi)^{d}} \frac{1}{\mathbf{k}^{2\gamma_{1}} \left((\mathbf{p} - \mathbf{k})^{2} + m^{2} \right)^{\gamma_{2}}},$$
(5.305)

corresponds via (5.304) to the GKZ integral [123, 124]

$$\mathcal{I}_{\gamma} = \int_{\mathbb{R}^2_+} \mathrm{d}z_1 \mathrm{d}z_2 \frac{z_1^{\gamma_1 - 1} z_2^{\gamma_2 - 1}}{(x_1 z_1 + x_2 z_2 + x_3 z_1 z_2 + x_4 z_2^2)^{\gamma_0}}$$
(5.306)

evaluated on the physical hypersurface

$$(x_1, x_2, x_3, x_4) = (1, 1, p^2 + m^2, m^2).$$
(5.307)

In this simple case, the GKZ integral can of course be evaluated directly,

$$\mathcal{I}_{\gamma} = \frac{\Gamma(\gamma_1)\Gamma(\gamma_0 - \gamma_1)\Gamma(\gamma_1 + \gamma_2 - \gamma_0)\Gamma(2\gamma_0 - \gamma_1 - \gamma_2)}{\Gamma(\gamma_0)^2} \times m^{2(\gamma_0 - \gamma_1 - \gamma_2)} {}_2F_1\left(\gamma_1, \gamma_1 + \gamma_2 - \gamma_0; \gamma_0; -\frac{p^2}{m^2}\right),$$
(5.308)

enabling the action of all creation operators to be verified. The \mathcal{A} -matrix is

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 \\ 1 & 0 & 1 & 0 \\ 0 & 1 & 1 & 2 \end{pmatrix}, \tag{5.309}$$

and from its kernel, we find a single toric equation

$$0 = (\partial_1 \partial_4 - \partial_2 \partial_3) \mathcal{I}_{\gamma}. \tag{5.310}$$

The Euler equations can be read off from the rows of the \mathcal{A} -matrix,

$$0 = (\gamma_0 + \theta_1 + \theta_2 + \theta_3 + \theta_4)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_1 + \theta_1 + \theta_3)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_2 + \theta_2 + \theta_3 + 2\theta_4)\mathcal{I}_{\gamma}.$$
(5.311)

The (rescaled) Newton polytope derived from the column vectors of the \mathcal{A} -matrix is the parellelogram shown in figure 5.9. From (5.109), the GKZ integral is then singular for

$$2\gamma_0 - \gamma_1 - \gamma_2 = -k_1, \quad \gamma_1 + \gamma_2 - \gamma_0 = -k_2, \quad \gamma_0 - \gamma_1 = -k_3, \quad \gamma_1 = -k_4, \quad k_i \in \mathbb{Z}^+$$
(5.312)

consistent with the poles of the gamma functions in (5.308).

The annihilation operators ∂_j send $\gamma \to \gamma'$ while the creation operators \mathcal{C}_j send $\gamma' \to \gamma$ where, for each j, these parameters are related by

$$\begin{array}{ll} j = 1: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1 + 1, & \gamma'_2 = \gamma_2, \\ j = 2: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1, & \gamma'_2 = \gamma_2 + 1, \\ j = 3: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1 + 1, & \gamma'_2 = \gamma_2 + 1, \\ j = 4: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1, & \gamma'_2 = \gamma_2 + 2. \end{array}$$

$$\begin{array}{ll} (5.313) \end{array}$$

Knowing the location of the singular hyperplanes and the shifts generated by the creation operators, the *b*-functions can be constructed according to (5.138),

$$b_1 = \gamma_1(2\gamma_0 - \gamma_1 - \gamma_2),$$



Figure 5.9: The rescaled Newton polytope associated to the 1-mass bubble integral (5.306).

$$b_{2} = (\gamma_{0} - \gamma_{1})(2\gamma_{0} - \gamma_{1} - \gamma_{2}), b_{3} = \gamma_{1}(\gamma_{1} + \gamma_{2} - \gamma_{0}), b_{4} = (\gamma_{0} - \gamma_{1})(\gamma_{1} + \gamma_{2} - \gamma_{0}).$$
(5.314)

Their zeros serve to cancel the singularities that arise whenever the action of a creation operator shifts us from a finite to a singular integral. For example, C_4 shifts $k_2 \rightarrow k_2 + 1$ and $k_3 \rightarrow k_3 + 1$ which, according to (5.312), generates a singular integral when acting on finite integrals with either $k_2 = -1$ or $k_3 = -1$. These singularities, however, are cancelled by the zeros of b_4 .

Using the DWI and the Euler equations (5.311), we can now re-write

$$C_j \partial_j \mathcal{I}_{\gamma} = b_j(\gamma) \mathcal{I}_{\gamma} = B_j(\theta) \mathcal{I}_{\gamma} \tag{5.315}$$

where

$$B_{1} = (\theta_{1} + \theta_{3})(\theta_{1} + \theta_{2}) = (\theta_{1} + \theta_{2} + \theta_{3})\theta_{1} + \theta_{2}\theta_{3},$$

$$B_{2} = (\theta_{2} + \theta_{4})(\theta_{1} + \theta_{2}) = (\theta_{1} + \theta_{2} + \theta_{4})\theta_{2} + \theta_{1}\theta_{4},$$

$$B_{3} = (\theta_{1} + \theta_{3})(\theta_{3} + \theta_{4}) = (\theta_{1} + \theta_{3} + \theta_{4})\theta_{3} + \theta_{1}\theta_{4},$$

$$B_{4} = (\theta_{2} + \theta_{4})(\theta_{3} + \theta_{4}) = (\theta_{4} + \theta_{2} + \theta_{3})\theta_{4} + \theta_{2}\theta_{3}.$$
 (5.316)

By inspection, every term in B_j either contains an explicit factor of ∂_j already through θ_j , or else such a factor can be introduced using the toric equations. In B_1 and B_4 , for instance, we replace $\theta_2\theta_3 = x_2x_3\partial_2\partial_3 \rightarrow x_2x_3\partial_1\partial_4$. This enables the B_j to be factored (modulo the toric equations) in the form (5.315) yielding the creation operators

$$C_{1} = x_{1}(1 + \theta_{1} + \theta_{2} + \theta_{3}) + x_{2}x_{3}\partial_{4},$$

$$C_{2} = x_{2}(1 + \theta_{1} + \theta_{2} + \theta_{4}) + x_{1}x_{4}\partial_{3},$$

$$C_{3} = x_{3}(1 + \theta_{1} + \theta_{3} + \theta_{4}) + x_{1}x_{4}\partial_{2},$$

$$C_{4} = x_{4}(1 + \theta_{2} + \theta_{3} + \theta_{4}) + x_{2}x_{3}\partial_{1}.$$
(5.317)

These creation operators act on the full GKZ integral (5.306). To obtain their counterparts

acting on the Feynman integral (5.305), we must project to the physical hypersurface (5.307). Given the form of the operators (5.317), it is useful to first simplify using the DWI to

$$C_{1} = x_{1}(1 - \gamma_{0} - \theta_{4}) + x_{2}x_{3}\partial_{4},$$

$$C_{2} = x_{2}(1 - \gamma_{0} - \theta_{3}) + x_{1}x_{4}\partial_{3},$$

$$C_{3} = x_{3}(1 - \gamma_{0} - \theta_{2}) + x_{1}x_{4}\partial_{2},$$

$$C_{4} = x_{4}(1 - \gamma_{0} - \theta_{1}) + x_{2}x_{3}\partial_{1}.$$
(5.318)

Next, as all factors of x_j are placed to the left of all derivatives, we set

$$(x_1, x_2, x_3, x_4) \to (1, 1, m^2 + p^2, m^2)$$
 (5.319)

and replace all derivatives lying in directions off this hypersurface (namely ∂_1 and ∂_2) with those lying along the hypersurface. This can be accomplished using the Euler equations (5.311) projected according to (5.319), namely

$$\partial_1 \to -\gamma_1 - (m^2 + p^2)\partial_3, \qquad \partial_2 \to -\gamma_2 - (m^2 + p^2)\partial_3 - 2m^2\partial_4.$$
 (5.320)

In addition, we use the chain rule with $p^2 = x_3 - x_4$ and $m^2 = x_4$ to replace

$$\partial_3 = \partial_{p^2}, \qquad \partial_4 = -\partial_{p^2} + \partial_{m^2}.$$
 (5.321)

This yields

$$C_{1}^{\rm ph} = 1 - \gamma_{0} + p^{2} \partial_{m^{2}} - \theta_{p^{2}},$$

$$C_{2}^{\rm ph} = 1 - \gamma_{0} - \theta_{p^{2}},$$

$$C_{3}^{\rm ph} = (1 - \gamma_{0})m^{2} + (1 - \gamma_{0} + \gamma_{2})p^{2} + (p^{2} - m^{2})\theta_{p^{2}} + 2p^{2}\theta_{m^{2}}$$

$$C_{4}^{\rm ph} = (1 - \gamma_{0})m^{2} + \gamma_{1}p^{2} - (p^{2} + m^{2})\theta_{p^{2}}.$$
(5.322)

From (5.132), the action on the projected GKZ integral is then

$$\mathcal{C}_{1}^{\mathrm{ph}}\mathcal{I}_{\gamma_{0}',\gamma_{1}',\gamma_{2}'}(p^{2},m^{2}) = -\gamma_{0}^{-1}b_{1}\mathcal{I}_{\gamma_{0},\gamma_{1},\gamma_{2}} = -\gamma_{0}^{-1}\gamma_{1}(2\gamma_{0}-\gamma_{1}-\gamma_{2})\mathcal{I}_{\gamma_{0},\gamma_{1},\gamma_{2}}$$
(5.323)

and similarly for the other operators. When acting the original Feynman integral, there is an additional factor of $c_{\gamma'}/c_{\gamma}$ from (5.304) we must take into account giving

$$\mathcal{C}_{1}^{\mathrm{ph}}I_{\gamma_{0}',\gamma_{1}',\gamma_{2}'}(p^{2},m^{2}) = -\frac{1}{4\pi}I_{\gamma_{0},\gamma_{1},\gamma_{2}}.$$
(5.324)

All these results can be checked directly using (5.308) and the standard shift operators for the $_2F_1$ (see *e.g.*, [160]).



Figure 5.10: The massive bubble integral (5.326).

Massive bubble

Next we consider the full bubble graph with general masses m_1 and m_2 ,

$$I = \int \frac{\mathrm{d}^{d} \boldsymbol{k}}{(2\pi)^{d}} \frac{1}{(\boldsymbol{k}^{2} + m_{1}^{2})^{\gamma_{1}} \left((\boldsymbol{p} - \boldsymbol{k})^{2} + m_{2}^{2}\right)^{\gamma_{2}}}.$$
(5.325)

The corresponding GKZ integral is

$$\mathcal{I}_{\gamma} = \int_{\mathbb{R}^2_+} \mathrm{d}z_1 \mathrm{d}z_2 \frac{z_1^{\gamma_1 - 1} z_2^{\gamma_2 - 1}}{(x_1 z_1 + x_2 z_2 + x_3 z_1^2 + x_4 z_2^2 + x_5 z_1 z_2)^{\gamma_0}},\tag{5.326}$$

where $\gamma_0 = d/2$ and the physical hypersurface is

$$(x_1, x_2, x_3, x_4, x_5) = (1, 1, m_1^2, m_2^2, m_1^2 + m_2^2 + p^2).$$
 (5.327)

From the kernel of the \mathcal{A} -matrix

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 & 1 \\ 1 & 0 & 2 & 0 & 1 \\ 0 & 1 & 0 & 2 & 1 \end{pmatrix}$$
(5.328)

we obtain the toric equations

$$0 = (\partial_3 \partial_4 - \partial_5^2) \mathcal{I}_{\gamma}, \qquad 0 = (\partial_2 \partial_3 - \partial_1 \partial_5) \mathcal{I}_{\gamma}, \qquad 0 = (\partial_1 \partial_4 - \partial_2 \partial_5) \mathcal{I}_{\gamma}, \tag{5.329}$$

while the DWI and the Euler equations can be read off from the rows:

$$0 = \left(\gamma_0 + \sum_{i=1}^{5} \theta_i\right) \mathcal{I}_{\gamma}, \qquad 0 = (\gamma_1 + \theta_1 + 2\theta_3 + \theta_5) \mathcal{I}_{\gamma}, \qquad 0 = (\gamma_2 + \theta_2 + 2\theta_4 + \theta_5) \mathcal{I}_{\gamma}.$$
(5.330)

The rescaled Newton polytope corresponding to this \mathcal{A} -matrix is the quadrilateral shown in figure 5.11. The singular hyperplanes lie parallel to and outside the facets of this polytope:

$$\gamma_1 = -k_1, \quad \gamma_2 = -k_2, \quad 2\gamma_0 - \gamma_1 - \gamma_2 = -k_3, \quad -\gamma_0 + \gamma_1 + \gamma_2 = -k_4, \qquad k_i \in \mathbb{Z}^+.$$
 (5.331)

For illustration, let us now discuss the creation operator C_5 . All others can be obtained by similar computations. The annihilator ∂_5 sends $\gamma \to \gamma'$ where

$$\partial_5: \quad \gamma'_0 = \gamma_0 + 1, \quad \gamma'_1 = \gamma_1 + 1, \quad \gamma'_2 = \gamma_2 + 1,$$
(5.332)



Figure 5.11: The rescaled Newton polytope associated with the massive bubble GKZ integral (5.326).

while the creation operator C_5 acts in the opposite direction sending $\gamma' \to \gamma$. Given this shift and the location of the singular hyperplanes, we identify the *b*-function as

$$b_5 = \gamma_1 \gamma_2 (\gamma_0 - \gamma_1 - \gamma_2). \tag{5.333}$$

Using DWI and Euler equations, this can be re-written in terms of Euler operators as

$$B_5 = (\theta_1 + 2\theta_3 + \theta_5)(\theta_2 + 2\theta_4 + \theta_5)(\theta_3 + \theta_4 + \theta_5).$$
(5.334)

This expression can now be factorised as $C_5\partial_5$ by expanding out and using the toric equations to replace any terms not involving ∂_5 with equivalent terms containing this factor. Stripping off the factor of ∂_5 then yields C_5 in GKZ variables,

$$C_{5} = x_{5} \Big[2\theta_{3}^{2} + 2\theta_{4}^{2} + 8\theta_{3}\theta_{4} + 3(\theta_{3} + \theta_{4})(1 + \theta_{5}) + (1 + \theta_{5})^{2} + \theta_{2}(1 + 3\theta_{3} + \theta_{4} + \theta_{5}) \\ + \theta_{1}(1 + \theta_{2} + \theta_{3} + 3\theta_{4} + \theta_{5}) \Big] + x_{2}x_{3}\partial_{1}(1 + \theta_{1} + 2\theta_{3}) + x_{1}x_{4}\partial_{2}(1 + \theta_{2} + 2\theta_{4}) \\ + 2x_{3}x_{4}\partial_{5}(4 + \theta_{1} + \theta_{2} + 2\theta_{3} + 2\theta_{4}).$$
(5.335)

To project this operator to the physical hypersurface (5.327), we first use the Euler equations to replace

$$\theta_1 \to -\gamma_1 - 2\theta_3 - \theta_5, \qquad \theta_2 \to -\gamma_2 - 2\theta_4 - \theta_5.$$
 (5.336)

The two occurrences of ∂_1 and ∂_2 can be dealt with similarly by writing $\partial_i = (x_i)^{-1}\theta_i$ for i = 1, 2 and using (5.336). Then, setting $(x_1, x_2, x_3, x_4, x_5) \to (1, 1, x_3, x_4, x_5)$, we obtain

$$\mathcal{C}_{5}^{\text{ph}} = x_{5} \left[(1 - \gamma_{1})(1 - \gamma_{2}) + (1 - \gamma_{1} - \gamma_{2} - \theta_{5})(\theta_{3} + \theta_{4}) \right] + x_{3}(\gamma_{1} + 2\theta_{3} + \theta_{5})(\gamma_{1} - 1 + \theta_{5}) + x_{4}(\gamma_{2} + 2\theta_{4} + \theta_{5})(\gamma_{2} - 1 + \theta_{5}) - 2x_{3}x_{4}\partial_{5}(\gamma_{1} + \gamma_{2} - 4 + 2\theta_{5})$$
(5.337)

The remaining variables here are all physical since

$$(x_3, x_4, x_5) = (m_1^2, m_2^2, m_1^2 + m_2^2 + p^2)$$
(5.338)



Figure 5.12: The massive triangle graph (5.340).

and

$$\partial_3 = \partial_{m_1^2} - \partial_{p^2}, \qquad \partial_4 = \partial_{m_2^2} - \partial_{p^2}, \qquad \partial_5 = \partial_{p^2}. \tag{5.339}$$

5.6.2 Massive triangle

Since the *massless* triangle integral is equivalent [42] to the 3-point contact Witten diagram studied in sections 5.5.3 and 5.5.5, let us examine here the massive triangle integral

$$I = \int \frac{\mathrm{d}^{d} \boldsymbol{k}}{(2\pi)^{d}} \frac{1}{(\boldsymbol{k}^{2} + m_{3}^{2})^{\gamma_{3}} \left((\boldsymbol{k} - \boldsymbol{p}_{1})^{2} + m_{2}^{2}\right)^{\gamma_{2}} \left((\boldsymbol{k} + \boldsymbol{p}_{2})^{2} + m_{1}^{2}\right)^{\gamma_{1}}}.$$
 (5.340)

The corresponding GKZ integral according to (5.304) is

$$\mathcal{I}_{\gamma} = \int_{\mathbb{R}^3_+} \mathrm{d}z_1 \mathrm{d}z_2 \mathrm{d}z_3 \, z_1^{\gamma_1 - 1} z_2^{\gamma_2 - 1} z_3^{\gamma_3 - 1} \mathcal{D}^{-\gamma_0}, \qquad (5.341)$$

where

$$\mathcal{D} = x_1 z_1 + x_2 z_2 + x_3 z_3 + x_4 z_2 z_3 + x_5 z_1 z_3 + x_6 z_1 z_2 + x_7 z_1^2 + x_8 z_2^2 + x_9 z_3^2 \tag{5.342}$$

with $\gamma_0 = d/2$. The physical hypersurface is

$$\boldsymbol{x} = (1, 1, 1, p_1^2 + m_2^2 + m_3^2, p_2^2 + m_1^2 + m_3^2, p_3^2 + m_1^2 + m_2^2, m_1^2, m_2^2, m_3^2)$$
(5.343)

and the \mathcal{A} -matrix reads

For larger \mathcal{A} -matrices such as this one, it is useful to automate the calculation of creation operators using Gröbner basis algorithms. To this end, in place of the five independent toric equations spanning the kernel of the \mathcal{A} -matrix, we will use instead the full set of 17



Figure 5.13: The Newton polytope associated to the denominator of the massive triangle integral (5.342). The label {i} denotes the vector defined by the *i*th column of the \mathcal{A} -matrix.

(non-independent) toric equations forming the toric ideal:¹⁵

$$I_{toric} = \{\partial_2\partial_6 - \partial_1\partial_8, \ \partial_7\partial_8 - \partial_6^2, \ \partial_1\partial_6 - \partial_2\partial_7, \ \partial_1\partial_5 - \partial_3\partial_7, \ \partial_3^2\partial_9 - \partial_7, \ \partial_1\partial_4 - \partial_3\partial_6, \\ \partial_4^2\partial_9 - \partial_8, \ \partial_5\partial_6 - \partial_4\partial_7, \ \partial_4\partial_6 - \partial_5\partial_8, \ \partial_4\partial_5\partial_9 - \partial_6, \ \partial_3^2\partial_8\partial_9 - \partial_2^2, \ \partial_3^2\partial_7\partial_9 - \partial_1^2, \\ \partial_3^2\partial_6\partial_9 - \partial_1\partial_2, \ \partial_3\partial_5\partial_9 - \partial_1, \ \partial_3\partial_4\partial_9 - \partial_2, \ \partial_2\partial_5 - \partial_3\partial_6, \ \partial_2\partial_4 - \partial_3\partial_8\}.$$
(5.345)

Each entry here corresponds to a toric equation, for example the first is $0 = (\partial_2 \partial_6 - \partial_1 \partial_8) \mathcal{I}_{\gamma}$ and similarly for the rest. Since all the partial derivatives commute, these equations can be treated as a system of polynomial equations by mapping ∂_i to an ordinary commutative variable y_i . As we will show below, this enables the factorisation step to be handled via ordinary commutative Gröbner basis methods. (For alternative constructions of creation operators using *non-commutative* Gröbner bases over the Weyl algebra, see [148].)

The DWI and Euler equations for this \mathcal{A} -matrix are

$$0 = \left(\gamma_0 + \sum_{i=1}^9 \theta_i\right) \mathcal{I}_{\gamma},$$

$$0 = (\gamma_1 + \theta_1 + \theta_5 + \theta_6 + 2\theta_7) \mathcal{I}_{\gamma},$$

$$0 = (\gamma_2 + \theta_2 + \theta_4 + \theta_6 + 2\theta_8) \mathcal{I}_{\gamma},$$

$$0 = (\gamma_3 + \theta_3 + \theta_4 + \theta_5 + 2\theta_9) \mathcal{I}_{\gamma}$$
(5.346)

and the corresponding Newton polytope is depicted in figure 5.13. From its facets, we

¹⁵ These can be obtained using the Singular code [151]:

obtain the singularity conditions

where all $k_i \in \mathbb{Z}^+$. For the facet {123}, for example, we have the outward-pointing normal $\boldsymbol{n} = (-1, -1, -1)$ which leads via (5.108) to the spacing of singular hyperplanes $\delta^{(J)} = 1$. Let us now compute the creation operator \mathcal{C}_4 which acts on the GKZ integral to shift $\boldsymbol{\gamma}' \to \boldsymbol{\gamma}$, where

$$\gamma'_0 = \gamma_0 + 1, \qquad \gamma'_1 = \gamma_1, \qquad \gamma'_2 = \gamma_2 + 1, \qquad \gamma'_3 = \gamma_3 + 1.$$
 (5.348)

From these parameter shifts and the location of the singular hyperplanes, the corresponding b-function is

$$b_4 = \gamma_2 \gamma_3 (\gamma_0 - \gamma_1 - \gamma_2 - \gamma_3). \tag{5.349}$$

Using the DWI and Euler equations, this can be re-expressed as

$$B_4 = \left(\sum_{i=4}^{9} \theta_i\right)(\theta_2 + \theta_4 + \theta_6 + 2\theta_8)(\theta_3 + \theta_4 + \theta_5 + 2\theta_9).$$
(5.350)

Our goal is now to factorise B_4 as $C_4\partial_4$ using the toric equations. To achieve this in an automated fashion, we decompose B_4 over the Gröbner basis formed from the toric ideal (5.345) and ∂_4 . Treating the partial derivatives as ordinary commutative variables and computing this Gröbner basis, we obtain

$$\boldsymbol{g} = \{\partial_4, \ \partial_2, \ \partial_8, \ \partial_3^2 \partial_7 \partial_9 - \partial_1^2, \ \partial_6, \ \partial_1 \partial_5 - \partial_3 \partial_7, \ \partial_3 \partial_5 \partial_9 - \partial_1, \ \partial_5^2 \partial_9 - \partial_7\}.$$
(5.351)

Expanding out B_4 and rewriting all terms in the form (5.131) so that all partial derivatives ∂_i lie to the right of all x_i , we can now decompose each term of B_4 in this Gröbner basis. This yields

$$B_4 = \boldsymbol{Q} \cdot \boldsymbol{g} = Q_1 \partial_4 + \sum_{i=2}^8 Q_i g_i.$$
(5.352)

where the coefficients Q_i are polynomials in the x_j and θ_j (with j = 1, ..., 9) which can be computed automatically.¹⁶ To extract the required overall factor of ∂_4 , we now re-express those g_i (i = 2, ..., 18) that are not already complete toric equations (and hence zero) in terms of ∂_4 . For example, using the third from last toric equation in (5.345), we can

¹⁶ In Mathematica, for example, after writing B_4 in the form (5.131) with all derivatives to the right, we replace all ∂_i (both in B_4 and in the toric ideal) by commutative variables y[i]. The code

v = {y[5], y[6], y[7], y[8], y[9], y[1], y[2], y[3], y[4]}; toric = {y[2] y[6] - y[1] y[8], y[6]² - y[7] y[8], y[1] y[6] - y[2] y[7], ...}; g = GroebnerBasis[Append[toric, y[4]], v] Q = PolynomialReduce[B4, g, v][[1]]

then evaluates the Q_i coefficients with all derivatives y[i] placed to the right. These can then be re-expressed in terms of Euler operators by rewriting $y_i^n = x_i^{-n} \theta_i (\theta_i - 1) \dots (\theta_i - n + 1)$ leading to (5.356).

replace

$$g_2 = \partial_2 \quad \rightarrow \quad \partial_3 \partial_4 \partial_9. \tag{5.353}$$

In this fashion, we can replace the basis g with the equivalent basis (modulo the toric equations)

$$\tilde{\boldsymbol{g}} = \{\partial_4, \ \partial_3\partial_4\partial_9, \ \partial_4^2\partial_9, \ 0, \ \partial_4\partial_5\partial_9, \ 0, \ 0\}.$$
(5.354)

All surviving terms then have an explicit factor of ∂_4 which can be removed to obtain the creation operator

$$\mathcal{C}_4 = Q_1 + Q_2 \partial_3 \partial_9 + Q_3 \partial_4 \partial_9 + Q_5 \partial_5 \partial_9, \tag{5.355}$$

where the coefficients are

$$Q_{1} = x_{4} \Big[(1 + \theta_{4})(\theta_{5} + \theta_{7} + \theta_{9}) + \theta_{5}(\theta_{6} + \theta_{7} + 3\theta_{8} + \theta_{9}) + \theta_{6}(\theta_{7} + \theta_{8} + 3\theta_{9}) \\ + (\theta_{8} + \theta_{9})(1 + 2\theta_{7} + \theta_{8} + \theta_{9}) + 4\theta_{8}\theta_{9} + (1 + \theta_{4} + \theta_{5} + \theta_{6} + \theta_{8} + \theta_{9})^{2} \\ + \theta_{3}(1 + \theta_{4} + \theta_{5} + 2\theta_{6} + \theta_{7} + 3\theta_{8} + \theta_{9}) \\ + \theta_{2}(1 + \theta_{3} + \theta_{4} + 2\theta_{5} + \theta_{6} + \theta_{7} + \theta_{8} + 3\theta_{9}) \Big],$$

$$Q_{2} = x_{2}(\theta_{3} + \theta_{5} + 2\theta_{9})(\theta_{5} + \theta_{6} + \theta_{7} + \theta_{8} + \theta_{9}),$$

$$Q_{3} = x_{8}(\theta_{3} + \theta_{5} + 2\theta_{9})(2\theta_{5} + 3\theta_{6} + 2(1 + \theta_{7} + \theta_{8} + \theta_{9})),$$

$$Q_{5} = x_{6}(\theta_{3} + \theta_{5} + 2\theta_{9})(1 + \theta_{5} + \theta_{6} + \theta_{7} + \theta_{9}).$$
(5.356)

Finally, to project to the physical hypersurface, we use the Euler equations to eliminate the unphysical variables $\theta_1, \theta_2, \theta_3$ and set $x_1 = x_2 = x_3 = 1$. This yields the physical creation operator

$$\mathcal{C}_{4}^{\text{ph}} = x_{4} \left[(1 - \gamma_{2})(1 - \gamma_{3}) + (\theta_{5} + \theta_{6} + \theta_{7} + \theta_{8} + \theta_{9})(1 - \gamma_{2} - \gamma_{3} - \theta_{4}) \right] \\
+ (\gamma_{3} + \theta_{4}) \left[\partial_{9}(\theta_{5} + \theta_{6} + \theta_{7} + \theta_{8} + \theta_{9})(\gamma_{3} + \gamma_{4} + \gamma_{5} + 2\theta_{9}) \\
- x_{6} \partial_{5}(1 + \theta_{5} + \theta_{6} + \theta_{7} + 3\theta_{8} + \theta_{9}) - 2x_{8} \partial_{4} \partial_{9}(1 + \theta_{5} + \theta_{7} + \theta_{8} + \theta_{9}) \right] \quad (5.357)$$

where the x_i are as given in (5.343) and

$$\partial_{4} = \partial_{p_{1}^{2}}, \qquad \partial_{5} = \partial_{p_{2}^{2}}, \qquad \partial_{6} = \partial_{p_{3}^{2}}, \\ \partial_{7} = \partial_{m_{1}^{2}} - \partial_{p_{2}^{2}} - \partial_{p_{3}^{2}}, \qquad \partial_{8} = \partial_{m_{2}^{2}} - \partial_{p_{1}^{2}} - \partial_{p_{3}^{2}}, \qquad \partial_{9} = \partial_{m_{3}^{2}} - \partial_{p_{1}^{2}} - \partial_{p_{2}^{2}}.$$
(5.358)

The automated approach outlined here can be applied similarly to other examples.

5.6.3 Massless on-shell box

Next we consider the massless box integral

$$I = \int \frac{\mathrm{d}^{d} \boldsymbol{q}}{(2\pi)^{d}} \frac{1}{|\boldsymbol{q}|^{2\gamma_{1}} |\boldsymbol{q} + \boldsymbol{P}_{1}|^{2\gamma_{2}} |\boldsymbol{q} + \boldsymbol{P}_{2}|^{2\gamma_{3}} |\boldsymbol{q} + \boldsymbol{P}_{3}|^{2\gamma_{4}}},$$
(5.359)

where

$$P_k = \sum_{j=1}^k p_j$$
, for $k = 1, 2, 3$, $\sum_{i=1}^4 p_i = 0.$ (5.360)



Figure 5.14: The on-shell massless box integral (5.359).

For simplicity, we will restrict to the on-shell case¹⁷ where all $p_i^2 = 0$ for i = 1, ..., 4. According to (5.304), the corresponding GKZ integral is

$$\mathcal{I}_{\gamma} = \prod_{i=1}^{4} \left(\int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1} \right) (x_{1}z_{1} + x_{2}z_{2} + x_{3}z_{3} + x_{4}z_{4} + x_{5}z_{1}z_{3} + x_{6}z_{2}z_{4})^{-\gamma_{0}}, \quad (5.361)$$

where the physical hypersurface

$$\boldsymbol{x} = (1, 1, 1, 1, s^2, t^2) \tag{5.362}$$

with $s^2 = (\mathbf{p}_1 + \mathbf{p}_2)^2$ and $t^2 = (\mathbf{p}_2 + \mathbf{p}_3)^2$ the Mandelstam invariants. The integral can be evaluated as a linear combination of the hypergeometric function $_3F_2$ [161].

The \mathcal{A} -matrix

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 & 1 & 1 \\ 1 & 0 & 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 & 0 & 1 \end{pmatrix}$$
(5.363)

yields a single toric equation

$$0 = (\partial_1 \partial_3 \partial_6 - \partial_2 \partial_4 \partial_5) \mathcal{I}_{\gamma}, \tag{5.364}$$

along with the DWI and Euler equations

$$0 = \left(\gamma_0 + \sum_{i=1}^6 \theta_i\right) \mathcal{I}_{\gamma}, \quad 0 = (\gamma_1 + \theta_1 + \theta_5) \mathcal{I}_{\gamma}, \quad 0 = (\gamma_2 + \theta_2 + \theta_6) \mathcal{I}_{\gamma}, \\ 0 = (\gamma_3 + \theta_3 + \theta_5) \mathcal{I}_{\gamma}, \quad 0 = (\gamma_4 + \theta_4 + \theta_6) \mathcal{I}_{\gamma}.$$
(5.365)

To determine the singularities of the integral, we need to find the equations of the facets of the rescaled Newton polytope corresponding to the GKZ denominator in (5.361). As this polytope lives in four dimensions, it is convenient to use an automated hulling algorithm.

¹⁷ Creation operators for the off-shell box are also computable but the results are rather long.

Using the Mathematica package [156], for example, we can enter the vertices a_j of the non-rescaled Newton polytope (where a_j is the *j*th column of the \mathcal{A} -matrix without the top row) as row vectors:

verts = $\{\{1,0,0,0\},\{0,1,0,0\},\{0,0,1,0\},\{0,0,0,1\},\{1,0,1,0\},\{0,1,0,1\}\};$

The command CHNQuickHull[verts] then returns a list of the vertex vectors that make up the convex hull (labelled according to the numbering specified in the input), followed by a list of the facets. The latter are specified by the vertex vectors they contain. Thus, in this example, we obtain

$$\{\{1,2,3,4,5,6\}, \{\{1,2,3,5\}, \{1,2,4,3\}, \{1,2,5,6\}, \{1,2,6,4\}, \{1,3,4,5\}, \{1,4,6,5\}, \{2,3,5,6\}, \{2,3,6,4\}, \{3,4,5,6\}\}\}$$

where the first set indicates that all six vertices belong to the convex hull, while the remainder ($\{1, 2, 3, 5\}$, $\{1, 2, 4, 3\}$, *etc.*) list the facets. Here, each $\{ijkl\}$ is a co-dimension one facet containing the points (a_i, a_j, a_k, a_l) .

The equations for the facets of the *rescaled* Newton polytope with vertices $\gamma_0 a_i$ can now be computed through a determinant such as (5.113). For the facet $\{1, 2, 3, 4\}$, for example, we have

$$0 = \boldsymbol{\gamma} \cdot \boldsymbol{N} = \det\left(\boldsymbol{\gamma} \mid \boldsymbol{\mathcal{A}}_1 \mid \boldsymbol{\mathcal{A}}_2 \mid \boldsymbol{\mathcal{A}}_3 \mid \boldsymbol{\mathcal{A}}_4\right) = \gamma_0 - \gamma_1 - \gamma_2 - \gamma_3 - \gamma_4 \tag{5.366}$$

and hence $\mathbf{N} = (n_0, \mathbf{n}) = (1, -1, -1, -1, -1)$. The fact that \mathbf{n} is outwards-pointing can be verified by showing $d_i^{(J)} = -\mathcal{A}_i \cdot \mathbf{N} > 0$ for any vertex i = 5, 6 not lying in the facet. The spacing of the set of singular hyperplanes parallel to this facet is then $\delta^{(J)} = 1$ using (5.108) and (5.99), with the singular hyperplanes themselves then following from (5.109). Automating this procedure and applying it to the other facets, the singularities for the GKZ integral (5.361) are

$$\gamma_{i} = -k_{i}, \quad i = 1, 2, 3, 4, \qquad \gamma_{1} + \gamma_{2} + \gamma_{3} + \gamma_{4} - \gamma_{0} = -k_{5}, \\
\gamma_{1} + \gamma_{2} - \gamma_{0} = +k_{6}, \qquad \gamma_{2} + \gamma_{3} - \gamma_{0} = +k_{7}, \\
\gamma_{3} + \gamma_{4} - \gamma_{0} = +k_{8}, \qquad \gamma_{4} + \gamma_{1} - \gamma_{0} = +k_{9},
\end{cases}$$
(5.367)

where all $k_i \in \mathbb{Z}^+$.

We are now in a position to compute the creation operators. Let us choose C_1 , which acts on the GKZ integral to shift $\gamma' \to \gamma$ where

$$\gamma_0' = \gamma_0 + 1, \qquad \gamma_1' = \gamma_1 + 1.$$
 (5.368)

From the singularities (5.367), the corresponding *b*-function is

$$b_1 = -\gamma_1(\gamma_2 + \gamma_3 - \gamma_0)(\gamma_3 + \gamma_4 - \gamma_0), \qquad (5.369)$$

which in terms of the Euler operators reads

$$B_1 = (\theta_1 + \theta_2)(\theta_1 + \theta_4)(\theta_1 + \theta_5)$$

= $[(\theta_1 + \theta_2)(\theta_1 + \theta_4) + (\theta_1 + \theta_2 + \theta_4)\theta_5]\theta_1 + \theta_2\theta_4\theta_5$ (5.370)

Applying the toric equation (5.364) to the final term now enables us to factorise B_1 as $C_1\partial_1$ giving the creation operator

$$\mathcal{C}_1 = x_1 \left[(1 + \theta_1 + \theta_2)(1 + \theta_1 + \theta_4) + (1 + \theta_1 + \theta_2 + \theta_4)\theta_5 \right] + x_2 x_4 x_5 \partial_3 \partial_6, \qquad (5.371)$$

where we shifted the factor of x_1 to the left sending each $\theta_1 \rightarrow \theta_1 + 1$. Finally, to obtain the creation operator acting on the physical variables, we use the Euler equations (5.365) to replace θ_i for i = 1, 2, 3, 4 with θ_5 and θ_6 and project to (5.362). This yields the operator

$$\mathcal{C}_{1}^{\text{ph}} = (1 - \gamma_1 - \gamma_2 - \theta_{t^2})(1 - \gamma_1 - \gamma_4 - \theta_{t^2}) + (\gamma_1 - 1)\theta_{s^2} - s^2(\gamma_3 + \theta_{s^2})\partial_{t^2}.$$
 (5.372)

Using the automated determination of the convex hull in this example, and the factorisation of the *b*-function via Gröbner basis methods in the previous example, the calculation of any creation operator can be fully automated.

5.7 Discussion

As we have seen, the GKZ formalism enables the construction of non-trivial shift operators known as creation operators. The calculation is highly systematic. First, a Feynman or Witten diagram is represented as a GKZ or \mathcal{A} -hypergeometric function. Second, the *b*function is identified by examining the parameter shifts produced by the creation operator in conjunction with the location of all singular hyperplanes of the integral. Physically, as the *b*-function is the function of parameters that multiplies the shifted integral, its zeros serve to cancel the singularities that would otherwise occur when the creation operator maps a finite to a singular integral. Such singularities cannot arise under the action of a finite differential operator on a finite integral. Next, using the Euler equations and DWI, the *b*-function is expressed as a function of Euler operators and factorised into a product of a creation and an annihilation operator with the aid of the toric equations. The creation operator thus extracted is then re-expressed in terms of physical variables (*i.e.*, the momenta and masses) using once again the Euler equations and DWI.

This algorithm has a number of interesting features. First, the parametric singularities of the integral all lie on hyperplanes parallel to the facets of the Newton polytope associated with the integral's denominator. We derived a precise formula for the spacing of these hyperplanes in (5.109). The *b*-function therefore has a geometrical character, as originally shown by Saito in [146]. Second, the algorithm makes heavy use of the higher-dimensional GKZ space obtained by promoting the coefficient of every term in the Lee-Pomeransky denominator to an independent variable. This systematises the set of PDEs obeyed by the integral into two distinct classes: the Euler equations and DWI, and the toric equations. Using the former, we can uplift to GKZ space by exchanging all dependence on the parameters γ for dependence on the additional unphysical coordinates. Conversely, we can project back to physical variables by using the Euler equations and DWI to exchange derivatives with respect to the unphysical variables for derivatives with respect to the physical variables and dependence on the parameters γ .

This last step is however a potential weakness of the algorithm. To project a creation operator from GKZ space back to the physical hypersurface, the total number of Euler equations (including the DWI) must be equal to, or greater than, the number of unphysical coordinates. This enables every derivative in unphysical variables to be replaced by an equivalent expression in purely physical variables. For higher-loop Feynman integrals, however, the number of terms in the Lee-Pomeransky denominator, and hence the dimension of the full GKZ space, typically grows more rapidly than the number of propagators and hence Euler equations. Thus, while a full set of creation operators can be constructed in GKZ space, in general we lack sufficient Euler equations to project back to the physical hypersurface. For this reason, we have focused here on 1-loop Feynman integrals.

One possible workaround for this issue is to construct an alternative projectible GKZ system based on some representation other than the Lee-Pomeransky. For example, for higher-loop massive sunset (aka melon or banana) diagrams, one can construct a GKZ representation based on their position-space formulation as a product of Bessel functions [127, 131]. In this manner, these diagrams can be related to (analytic continuations of) the momentum-space contact Witten diagrams for which we have already constructed creation operators. For more general classes of diagrams, projectible GKZ representations can also be obtained from Mellin-Barnes representations as shown in [128, 162, 131]. A further possibility might be to develop a GKZ representation starting from the Baikov representation.

Nevertheless, using the simplest formulation based on the Lee-Pomeransky representation, we have already identified a number of useful new shift operators. In particular, for computations in AdS/CFT, we have found:

- The creation operators (5.206) and (5.224), along with their permutations and shadow conjugates, connecting 3- and 4-point momentum-space contact Witten diagrams of different operator and spacetime dimensions. These new operators are the inverse of the simple annihilators first identified in [42, 85]. The corresponding operators can also be obtained in position space as detailed in Appendix C.2.
- The creation operators (5.255) and (5.271), plus their permutations and shadow conjugates, relating 3- and 4-point momentum-space contact Witten diagrams of different operator dimensions but the same spacetime dimension. While the 3-point operator (5.255) is known [83, 30], the 4-point operator (5.271) is new.
- Using (5.271), we obtained a further new operator (5.293) connecting exchange Witten diagrams of different external operator dimensions but the same spacetime dimension. Unlike any previous construction, this operator connects exchange diagrams with purely *non-derivative vertices*. Working in dimensional regularisation where necessary to avoid divergences [36], it also applies for arbitrary operators dimensions.

There is ample scope for building on this first application of creation operators to Witten diagrams. In particular, our results for exchange Witten diagrams were obtained from our analysis of contact diagrams. It may be preferable to develop a GKZ representation for the exchange diagram directly, both in momentum and in position space, potentially enabling a more compact set of shift operators to be found, as well as operators acting to shift the dimension of the exchanged leg. Operators achieving this latter goal are at present known only for a very restricted set of external operator dimensions [30, 36]. The application of the creation operator formalism to cosmological correlators in de Sitter spacetime is also worthy of exploration. We hope to address some of these matters in future.

Part IV Conclusions

Chapter 6 Conclusions

In this thesis we presented studies of conformal field theory in momentum space, focusing on integral representations and shift operators. The mathematical and physical tools that we used were quite diverse, spanning from electrical circuits to multivariable hypergeometric systems. We believe this work shows the importance of studying an object from different perspectives. On one side, the recent formulation of conformal symmetry in momentum space enlarges the domain of possible physical applications of conformal symmetry – borrowing Whitman's words, we could say that conformal symmetry is large, it contains multitudes. And we have shown that different representations of the same function reveal different properties.

In Chapter 4 we took as input the general *n*-point solution of conformal Ward identities in the form of the simplex integral (4.1) and derived the new scalar parametrisations, (4.17), (4.48), for this integral which allowed us to find the new shift operators S_{ij} and $\mathcal S$ for the *n*-point functions. To this aim, we first parametrised the integral in terms of the inverse Schwinger parameters v_{ij} and obtained the associated Kirchhoff polynomials, expressible in terms of the Gram determinant or the Laplacian matrix. We then interpreted the parameters v_{ij} as conductances of the corresponding simplicial electrical network and computed the effective resistances s_{ij} between the nodes of the simplex. This led to a new reparametrisation of the simplex integral in terms of the Cayley-Menger matrix. This parametrisation features an exponential diagonal in the s_{ij} , and a product of powers of the determinant |m| and first minors $|m^{(i,j)}|$ of the Cayley-Menger matrix. The structure of this integral then naturally reveals the shift operators. Multiplying the integrand by either the determinant or a first minor of the Cayley-Menger matrix corresponds to shifting their powers. The diagonal exponential allowed us to translate the polynomials in s_{ij} , defining |m| and $|m^{(i,j)}|$, into differential operators with respect to $p_i \cdot p_j$. Besides finding new shift operators, we discussed other advantages of these parametrisations of the simplex. We noted that the number of integrals to be computed reduces from (n-1)(n-2)d/2 to n(n-1)/2, and the correspondence to the position-space solution simplifies evaluating the action of certain differential operators on the momentum-space integral. In particular, we showed that the special conformal Ward identity corresponds to a total derivative acting on the integrand, and we computed the action of the known weight-shifting operators $\mathcal{W}_{ij}^{\pm\pm}.$

Still, much remains to be explored about the *n*-point function. We have seen how the simplex perspective makes the recursive structure evident, and how these new parametrisa-

tions make the shift operators naturally arise. Perhaps by looking for other representations of the simplex, new properties would become manifest. For instance, we might ask whether a representation invariant under a shadow transform exists. Moreover, notice that we have only analysed scalar correlators. Having available a full set of shift operators which generalise the 3-point \mathcal{L}_i operators, together with the other known spin- and weight-shifting operators [83, 29, 30], we now have the tools to build tensorial correlators that are also of interest in cosmology.

The study of integral representations led us to deepen our knowledge of the multivariable hypergeometric functions in the form of GKZ systems. In Chapter 5 we discussed their features, giving a physical interpretation. We presented the formulation of GKZ integrals and showed how their properties are encoded in the \mathcal{A} -matrix. We discussed the spectral singularities and their geometrical interpretation via the Newton polytope associated with the \mathcal{A} -matrix. We then used this formulation to derive the creation operators of holographic contact Witten diagrams and some generalised Feynman integrals. The construction of these operators is very systematic and uses the b-function as input. The b-function is a polynomial in the parameters and the whole algorithm is based on the fact that it can be factorised in terms of annihilation and creation operators as $B(\theta) \sim C\partial$. We gave a physical interpretation of the form of the b-function. We derived it by requiring that the associated creation operator cannot send a finite integral to a divergent one. Moreover, we used an analogous algorithm to find operators that act on contact and exchange Witten diagrams to shift only the scaling dimensions and preserve the form of the functions. As discussed, this class of operators is new and applies to any values of the parameters. Our construction of creation operators for Witten diagrams is moreover valid at n points.

The GKZ formalism entered the physics literature only recently. While various studies focused on the series solutions of Feynman integrals, to our knowledge our work is the first application of creation operators to physics, therefore this is only the beginning. There are many paths to explore. Can we find a more optimal way for constructing the GKZ representations of multi-loop Feynman integrals? Can we find a GKZ representation of exchange Witten diagrams? Coming back to the simplex representation, in Chapter 4, we derived its Lee-Pomeransky representation from which a GKZ representation can be obtained. This would give us the spectral singularities of the conformal *n*-point functions. Moreover, conformal blocks in position space were found to have a description in terms of multivariable hypergeometric systems based on root systems [163, 164]. An interesting question is whether we can construct a momentum-space description of conformal blocks, and also whether a connection between GKZ and root systems exists. A more immediate consequence of knowing various families of shift operators is their application to cosmological correlators.

With these various open directions we conclude this thesis but certainly not the research ahead.

Appendix A The master integral $I_{1\{000\}}$

In this appendix we evaluate the master integral $I_{1\{000\}}$. The strategy we follow is to start with the Feynman parametrisation of the 1-loop triangle diagram (3.75), and solve the resulting integral by making suitable changes of variables and using partial fractions method. We will show that the z-variables defined in (3.57) arise naturally.

For the master integral d = 4 and Δ_j (j = 1, 2, 3) and using (3.87), $a_{jk} = -1$. Setting the parameters at these values and Feynman parametrising the triangle representation we have

$$I_{1\{000\}} = \frac{1}{4} \int_{[0,1]^3} \mathrm{d}X \frac{1}{p_1^2 x_2 x_3 + p_2^2 x_1 x_3 + p_3^2 x_1 x_2},\tag{A.1}$$

where $dX = dx_1 dx_2 dx_3 \delta(x_1 + x_2 + x_3 - 1)$. Setting $u = p_1^2 / p_3^2$ and $v = p_2^2 / p_3^2$, the integral reads:

$$4p_3^2 I_{1\{000\}} = \int_0^1 \mathrm{d}x_3 \int_0^{1-x_3} \mathrm{d}x_2 \frac{1}{x_1 x_2 + x_2 x_3 u + x_1 x_3 v},\tag{A.2}$$

where $x_1 = 1 - x_2 - x_3$. Re-parametrising

$$y_1 = \frac{x_1}{x_3}, \quad y_2 = \frac{x_2}{x_3} \quad \Rightarrow \quad x_2 = \frac{y_2}{1 + y_1 + y_2}, \quad x_3 = \frac{1}{1 + y_1 + y_2}$$
(A.3)

and computing the Jacobian $|\partial x/\partial y| = (1 + y_1 + y_2)^{-3}$, we get

$$4p_3^2 I_{1\{000\}} = \int_0^\infty \mathrm{d}y_1 \int_0^\infty \mathrm{d}y_2 \frac{1}{(1+y_1+y_2)(y_1y_2+y_1v+y_2u)}.$$
 (A.4)

We then perform a first partial fraction

$$\frac{1}{(1+y_1+y_2)(y_1y_2+y_1v+y_2u)} = \frac{1}{y_1^2 + (1+u-v)y_1 + u} \times \left(\frac{u+y_1}{(u+y_1)y_2 + vy_1} - \frac{1}{1+y_1+y_2}\right), \quad (A.5)$$

and, integrating over y_2 , we find

$$4p_3^2 I_{1\{000\}} = \int_0^\infty \mathrm{d}y_1 \frac{1}{y_1^2 + (1+u-v)y_1 + u} \log\left(\frac{(1+y_1)(u+y_1)}{vy_1}\right). \tag{A.6}$$

The denominator is quadratic in y_1 , hence we can factorise it. This, indeed, leads to defining the z-variables as in (3.57) and the integral reads

$$4p_3^2 I_{1\{000\}} = \int_0^\infty \mathrm{d}y_1 \frac{1}{(y_1 + z)(y_1 + \bar{z})} \log\left(\frac{(1 + y_1)(y_1 + z\bar{z})}{y_1(1 - z)(1 - \bar{z})}\right).$$
(A.7)

Then, we perform a second partial fraction

$$\frac{1}{(y_1+z)(y_1+\bar{z})} = \frac{1}{\bar{z}-z} \left(\frac{1}{y_1+\bar{z}} - \frac{1}{y_1+z} \right),$$
(A.8)

and, after some manipulations, the integral becomes

$$4p_3^2(z-\bar{z})I_{1\{000\}} = \int_0^\infty \mathrm{d}y_1\left(\frac{1}{y_1+z} - \frac{1}{y_1+\bar{z}}\right)\log\left(\frac{(1+y_1)^2}{(1-z)(1-\bar{z})}\right).$$
 (A.9)

We now can evaluate it in terms of the dilogarithm giving the expression in equation 3.56, with $z_{1,2} \in \mathbb{C}-] - \infty, 0] \cup [1, +\infty[$. To see this, note that there will be terms of the form

$$\int_0^\infty \mathrm{d}y \frac{\log(1+y)}{y+z},\tag{A.10}$$

which can be computed by applying the Cauchy integral formula taking into account that the discontinuity of $\text{Li}_2(z)$ across the branch cut $[1, \infty]$ is equal to $2\pi i \log |z|$. Setting t = 1 + y:

$$\int_{1}^{\infty} \mathrm{d}t \frac{\log t}{t-1+z} = \frac{1}{2\pi i} \int_{1}^{\infty} \mathrm{d}t \frac{\mathrm{disc}(\mathrm{Li}_{2}(t))}{t-1+z} = \frac{1}{2\pi i} \oint_{C} \mathrm{d}t \frac{\mathrm{Li}_{2}(t)}{t-1+z} = \mathrm{Li}_{2}(1-z), \quad (A.11)$$

where C is the "pac-man" closed contour. Then, using dilogarithm's properties one can express $\text{Li}_2(1-z)$ in terms of $\text{Li}_2(z)$ and find the explicit result anticipated in (3.56):

$$I_{1\{000\}} = \frac{-1}{2p_3^2(z-\bar{z})} \left[\text{Li}_2(\bar{z}) - \text{Li}_2(z) - \frac{1}{2}\ln(z\bar{z})\ln\left(\frac{1-z}{1-\bar{z}}\right) \right].$$
 (A.12)

Appendix B

Appendix to chapter 4

B.1 Derivation of graph polynomials

In this appendix, we show that Schwinger representation of the simplex integral (4.1) is given by (4.17) with graph polynomials given in (4.10). Our discussion builds on that in [165]. Labelling the vertices of the simplex by i = 1, ..., n, and the (directed) legs by a = 1, ..., N where N = n(n-1)/2, we introduce the incidence matrix

$$\varepsilon_i^a = \begin{cases} +1 & \text{if } \log a \text{ is ingoing to vertex } i \\ -1 & \text{if } \log a \text{ is outgoing to vertex } i \\ 0 & \text{otherwise} \end{cases}$$
(B.1)

where for clarity we will write the vertex index downstairs and the leg index upstairs. Thus, for example, if we choose $a = \{(12), (13), (14), (23), (24), (34)\}$ as the legs of the 4-point function, where the leg (i, j) runs from vertex *i* to vertex *j*, the incidence matrix is

$$\varepsilon = \begin{pmatrix} -1 & -1 & -1 & 0 & 0 & 0\\ 1 & 0 & 0 & -1 & -1 & 0\\ 0 & 1 & 0 & 1 & 0 & -1\\ 0 & 0 & 1 & 0 & 1 & 1 \end{pmatrix}.$$
 (B.2)

Momentum conservation at vertex k of the simplex can now be re-expressed as

$$0 = \boldsymbol{p}_k + \sum_{l \neq k}^{n} \boldsymbol{q}_{lk} = \boldsymbol{p}_k + \sum_{a}^{N} \varepsilon_k^a \boldsymbol{q}_a$$
(B.3)

where q_a is the internal momentum flowing along the directed leg *a*. As always, all sums are assumed to begin at one unless otherwise specified. The Laplacian matrix \tilde{g}_{ij} defined in (4.23) can now be written

$$\tilde{g}_{ij} = \sum_{a}^{N} v_a \varepsilon_i^a \varepsilon_j^a, \tag{B.4}$$

which follows by noting that for $i \neq j$ only the leg for which *a* runs between vertices *i* and *j* contributes giving $-v_{ij}$, while for i = j all legs running into this vertex contribute giving $\sum_{k\neq i} v_{ik}$ as required.

Turning to the simplex integral (4.1), we first rewrite the delta functions of momentum conservation in Fourier form

$$\prod_{k}^{n} (2\pi)^{d} \delta \left(\boldsymbol{p}_{k} + \sum_{l \neq k}^{n} \boldsymbol{q}_{lk} \right) = \prod_{k}^{n} \int \mathrm{d}^{d} \boldsymbol{y}_{k} \exp \left(-i \boldsymbol{y}_{k} \cdot \left(\boldsymbol{p}_{k} + \sum_{a}^{N} \varepsilon_{k}^{a} \boldsymbol{q}_{a} \right) \right).$$
(B.5)

Next, exponentiating all propagators of internal momenta (labelled by their leg indices) using the Schwinger representation (4.4), we find¹

$$\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle = \left(\prod_a^N \frac{1}{\Gamma(\alpha_a + d/2)} \int_0^\infty \mathrm{d} v_a \, v_a^{-d/2 - \alpha_a - 1} \right) f(\hat{\boldsymbol{v}}) \\ \times \left(\prod_k^n \int \mathrm{d}^d \boldsymbol{y}_k \, \exp(-i\boldsymbol{y}_k \cdot \boldsymbol{p}_k) \right) \int \frac{\mathrm{d}^d \boldsymbol{q}_a}{(2\pi)^d} \exp\left(-\sum_a^N \left(\frac{q_a^2}{v_a} + i\sum_l^n \varepsilon_l^a \boldsymbol{y}_l \cdot \boldsymbol{q}_a \right) \right).$$
(B.6)

Evaluating the \boldsymbol{q}_a integrals by completing the square and using (B.4) now gives

$$\langle \mathcal{O}_1(\boldsymbol{p}_1) \dots \mathcal{O}_n(\boldsymbol{p}_n) \rangle = \left(\prod_a^N \frac{\pi^{d/2}}{\Gamma(\alpha_a + d/2)} \int_0^\infty \mathrm{d} v_a \, v_a^{-\alpha_a - 1} \right) f(\hat{\boldsymbol{v}}) \\ \times \left(\prod_k^n \int \mathrm{d}^d \boldsymbol{y}_k \right) \, \exp\left(-i \sum_k^n \boldsymbol{y}_k \cdot \boldsymbol{p}_k - \frac{1}{4} \sum_{k,l}^n \tilde{g}_{kl} \, \boldsymbol{y}_k \cdot \boldsymbol{y}_l \right).$$
(B.7)

Since the Laplacian matrix has no inverse, to compute the \boldsymbol{y}_k integrals we must first shift

$$\boldsymbol{y}_n = \boldsymbol{z}_n, \qquad \boldsymbol{y}_k = \boldsymbol{z}_k + \boldsymbol{z}_n, \qquad k = 1, \dots n - 1.$$
 (B.8)

This transformation has unit Jacobian, but moreover greatly simplifies the exponent. Since all row and column sums of the Laplacian matrix vanish,

$$\sum_{l}^{n-1} \tilde{g}_{kl} = -\tilde{g}_{kn}, \qquad \sum_{k}^{n-1} \tilde{g}_{kn} = -\tilde{g}_{nn}, \qquad (B.9)$$

and using these identities we then find

$$egin{aligned} &-i\sum_{k}^{n}oldsymbol{y}_{k}\cdotoldsymbol{p}_{k}-rac{1}{4}\sum_{k,l}^{n} ilde{g}_{kl}\,oldsymbol{y}_{k}\cdotoldsymbol{y}_{l}\ &=-ioldsymbol{z}_{n}\cdotoldsymbol{p}_{n}-i\sum_{k}^{n-1}(oldsymbol{z}_{k}+oldsymbol{z}_{n})\cdotoldsymbol{p}_{k} \end{aligned}$$

¹Note the argument of the arbitrary function f changes from the momentum cross ratios \hat{q} in (4.3) to the Schwinger parameter cross ratios \hat{v} in (4.16). This can be seen by temporarily representing the arbitrary function in Mellin-Barnes form (*i.e.*, (4.18) of [46]) allowing all q_{ij} , including those from the cross ratios, to be exponentiated via the Schwinger parametrisation (4.4). Performing the Mellin-Barnes integration then generates $f(\hat{v})$, since the Schwinger parametrisation replaces powers of q_{ij} by powers of v_{ij} .

$$-\frac{1}{4}\tilde{g}_{nn}z_{n}^{2} - \frac{1}{2}\sum_{k}^{n-1}\tilde{g}_{kn}(\boldsymbol{z}_{k} + \boldsymbol{z}_{n}) \cdot \boldsymbol{z}_{n} - \frac{1}{4}\sum_{k,l}^{n-1}\tilde{g}_{kl}(\boldsymbol{z}_{k} + \boldsymbol{z}_{n}) \cdot (\boldsymbol{z}_{l} + \boldsymbol{z}_{n})$$
$$= -i\boldsymbol{z}_{n} \cdot \left(\sum_{k}^{n}\boldsymbol{p}_{k}\right) - i\sum_{k}^{n-1}\boldsymbol{z}_{k} \cdot \boldsymbol{p}_{k} - \frac{1}{4}\sum_{k,l}^{n-1}g_{kl}\boldsymbol{z}_{k} \cdot \boldsymbol{z}_{l}.$$
(B.10)

In the final line here, all the $\mathbf{z}_k \cdot \mathbf{z}_n$ and z_n^2 terms cancel while the Laplacian matrix \tilde{g}_{kl} reduces to g_{kl} for k, l = 1, ..., n. The \mathbf{z}_n integral now gives the overall delta function of momentum conservation which we strip off to obtain the reduced correlator (4.15). The remaining \mathbf{z}_k integrals can be evaluated by completing the square, given that the inverse g_{kl}^{-1} exists. This yields our desired result,

$$\langle\!\langle \mathcal{O}_1(\boldsymbol{p}_1)\dots\mathcal{O}_n(\boldsymbol{p}_n)\rangle\!\rangle = \mathcal{C}\prod_a^N \int_0^\infty \mathrm{d}v_a \, v_a^{-\alpha_a - 1} f(\hat{\boldsymbol{v}}) \, |g|^{-d/2} \exp\left(-\sum_{k,l}^{n-1} g_{kl}^{-1} \, \boldsymbol{p}_k \cdot \boldsymbol{p}_l\right) \quad (B.11)$$

where the constant

$$C = (4\pi)^{(n-1)d/2} \prod_{a}^{N} \frac{\pi^{d/2}}{\Gamma(\alpha_a + d/2)}$$
(B.12)

can simply be re-absorbed into the arbitrary function $f(\hat{v})$. Rewriting the product of legs a as a product over vertices i < j and replacing $p_k \cdot p_l$ with the Gram matrix G_{kl} , we recover precisely (4.17) with graph polynomials (4.10).

B.2 Jacobian matrix

In this appendix, we compute the Jacobian matrix for the change of variables from v_{ij} to s_{ij} . In section B.2.1 we evaluate the Jacobian determinant, then in section B.2.2 we give expressions for its matrix elements enabling conversion between partial derivatives.

B.2.1 Jacobian determinant

Our first goal is to derive the relation (4.45) for the Jacobian determinant, namely

$$\left|\frac{\partial s}{\partial v}\right| = \left|\frac{\partial^2 \ln |g|}{\partial v \, \partial v}\right| \propto |g|^{-n},\tag{B.13}$$

where the constant of proportionality is not required since it can be re-absorbed into the arbitrary function $f(\hat{v})$. For small values of n this result can be verified by direct calculation, and the exponent is simply fixed by power counting, but our aim is nevertheless to prove this relation for general n.

We start by noting

$$\frac{\partial^2 \ln |g|}{\partial v_{ij} \partial v_{kl}} = \frac{\partial g_{pq}}{\partial v_{ij}} \frac{\partial^2 \ln |g|}{\partial g_{pq} \partial g_{rs}} \frac{\partial g_{rs}}{\partial v_{kl}}$$
(B.14)

can be re-expressed as a product of three square matrices of dimension n(n-1)/2. Each of the index pairs (p,q) and (r,s) is replaced by a single index running over the n(n-1)/2

independent entries of the $(n-1) \times (n-1)$ symmetric matrix g, while (i, j) and (k, l) are each replaced by a single index running over the n(n-1)/2 edges of the simplex. Noting the elements of g are linear in the v, the matrix determinant $|\partial g/\partial v|$ evaluates to a nonzero constant. On taking the determinant of (B.14), we find

$$\left|\frac{\partial^2 \ln|g|}{\partial v \,\partial v}\right| \propto \left|\frac{\partial^2 \ln|g|}{\partial g \,\partial g}\right| \tag{B.15}$$

hence it suffices to show that

$$\left|\frac{\partial^2 \ln |g|}{\partial g \, \partial g}\right| \propto |g|^{-n}.\tag{B.16}$$

This relation in fact holds for any invertible symmetric square matrix g of dimension n-1.

To see this, from Jacobi's relation we have

$$\frac{\partial^2 \ln |g|}{\partial g_{pq} \partial g_{rs}} = \frac{\partial}{\partial g_{rs}} \left(\frac{1}{|g|} (\operatorname{adj} g)_{pq} \right) = \frac{\partial (g^{-1})_{pq}}{\partial g_{rs}} = -2(g^{-1})_{p(r}(g^{-1})_{s)q}.$$
(B.17)

Diagonalising g via an orthogonal matrix O,

$$\Lambda = OgO^{-1},\tag{B.18}$$

since $g^{-1} = O^{-1} \Lambda^{-1} O$ the chain rule gives

$$\frac{\partial g^{-1}}{\partial g} = \frac{\partial (O^{-1}\Lambda^{-1}O)}{\partial \Lambda^{-1}} \frac{\partial \Lambda^{-1}}{\partial \Lambda} \frac{\partial (OgO^{-1})}{\partial g}$$
(B.19)

where the last factor is just $\partial \Lambda / \partial g$. Regarding this as a matrix product, the first and last matrices depend only on O and are inverses of each other. On taking the determinant of the right-hand side, their contributions therefore cancel giving

$$\left|\frac{\partial g^{-1}}{\partial g}\right| = \left|\frac{\partial \Lambda^{-1}}{\partial \Lambda}\right|.$$
 (B.20)

We thus only need to evaluate the latter determinant for the diagonal matrix Λ .

From (B.17), the Hessian is nonzero only when the index pairs are equal (p,q) = (r,s), and is thus diagonal when regarded as a square matrix of dimension n(n-1)/2:

$$\frac{\partial^2 \ln |\Lambda|}{\partial \Lambda_{pq} \partial \Lambda_{rs}} = \frac{\partial (\Lambda^{-1})_{pq}}{\partial \Lambda_{rs}} = \begin{cases} -\Lambda_{pp}^{-1} \Lambda_{qq}^{-1} & \text{if } (p,q) = (r,s) \\ 0 & \text{otherwise} \end{cases}$$
(B.21)

The determinant is now

$$\left|\frac{\partial^2 \ln |\Lambda|}{\partial \Lambda_{pq} \Lambda_{rs}}\right| \propto \prod_{p=1}^{n-1} (\Lambda_{pp})^{-n} = |\Lambda|^{-n} = |g|^{-n}$$
(B.22)

since each eigenvalue Λ_{pp} appears a total of *n* times along the diagonal: for example, Λ_{11} appears quadratically in the position (1, 1) and then linearly in each of the (n - 2) entries indexed by (1, q) for $q = 2, \ldots n - 1$. We have thus established (B.16), and hence (B.13).

B.2.2 Matrix elements

We now compute the elements of the Jacobian matrix required to establish the relation

$$\partial_{v_{kl}} = -\frac{1}{4} \sum_{i(B.23)$$

which we used in (4.92). Starting with (4.27),

$$\frac{\partial s_{ij}}{\partial v_{kl}} = \frac{\partial}{\partial v_{kl}} \left((g^{-1})_{ab} \frac{\partial g_{ab}}{\partial v_{ij}} \right) = \frac{\partial (g^{-1})_{ab}}{\partial v_{kl}} \frac{\partial g_{ab}}{\partial v_{ij}}$$
(B.24)

where since g_{ab} is linear in the v_{ij} its second derivative vanishes. Using

$$\frac{\partial (g^{-1})_{ab}}{\partial v_{kl}} = -(g^{-1})_{ae}(g^{-1})_{bf}\frac{\partial g_{ef}}{\partial v_{kl}}$$
(B.25)

then gives

$$\frac{\partial s_{ij}}{\partial v_{kl}} = -\mathrm{tr}\left(g^{-1} \cdot \frac{\partial g}{\partial v_{ij}} \cdot g^{-1} \cdot \frac{\partial g}{\partial v_{jk}}\right) = -\sum_{a,b,e,f}^{n-1} (g^{-1})_{ae} \frac{\partial g_{ef}}{\partial v_{kl}} (g^{-1})_{fb} \frac{\partial g_{ba}}{\partial v_{ij}}.$$
 (B.26)

For $i, j, k, l \neq n$, we can evaluate this as

$$\frac{\partial s_{ij}}{\partial v_{kl}} = -\sum_{a,b,e,f}^{n-1} (g^{-1})_{ae} (-2\delta_{k(e}\delta_{f)l} + \delta_{ek}\delta_{fk} + \delta_{el}\delta_{fl}))(g^{-1})_{fb} (-2\delta_{i(a}\delta_{b)j} + \delta_{ai}\delta_{bi} + \delta_{aj}\delta_{bj})
= -2(g^{-1})_{ik}(g^{-1})_{jl} - 2(g^{-1})_{il}(g^{-1})_{jk} + 2(g^{-1})_{ik}(g^{-1})_{jk} + 2(g^{-1})_{il}(g^{-1})_{jl}
+ 2(g^{-1})_{ik}(g^{-1})_{il} + 2(g^{-1})_{jk}(g^{-1})_{jl} - ((g^{-1})_{ik})^2 - ((g^{-1})_{il})^2 - ((g^{-1})_{jk})^2 - ((g^{-1})_{jl})^2
= -((g^{-1})_{ik} - (g^{-1})_{il} - (g^{-1})_{jk} + (g^{-1})_{jl}))^2
= -\frac{1}{4}(s_{ik} - s_{il} - s_{jk} + s_{jl})^2$$
(B.27)

where we used the symmetry of the inverse matrix g_{ij}^{-1} , and in the last line we used (4.28). For j = n but $i, k, l \neq n$,

$$\frac{\partial s_{in}}{\partial v_{kl}} = -\sum_{a,b,e,f}^{n-1} (g^{-1})_{ae} (-2\delta_{k(e}\delta_{f)l} + \delta_{ek}\delta_{fk} + \delta_{el}\delta_{fl}))(g^{-1})_{fb} (\delta_{ai}\delta_{bi})$$
$$= -((g^{-1})_{ik} - (g^{-1})_{il})^2 = -\frac{1}{4} (s_{ik} - s_{il} - s_{kn} + s_{ln})^2$$
(B.28)

which is equivalent to (B.27) setting j = n. The same also holds for l = n but $i, j, k \neq n$ due to the symmetry of (B.26). Finally

$$\frac{\partial s_{in}}{\partial v_{kn}} = -\sum_{a,b,e,f}^{n-1} (g^{-1})_{ae} (\delta_{ek} \delta_{fk})) (g^{-1})_{fb} (\delta_{ai} \delta_{bi})$$

$$= -((g^{-1})_{ik})^2 = -\frac{1}{4}(s_{ik} - s_{in} - s_{kn})^2,$$
(B.29)

also equivalent to (B.27) since $s_{nn} = 0$. Thus (B.27) in fact holds for all values of the indices and we obtain (B.23).

For completeness, we can also calculate the inverse Jacobian by similar means:

$$\frac{\partial v_{ij}}{\partial s_{kl}} = \frac{\partial}{\partial s_{kl}} \left((m^{-1})_{ab} \frac{\partial m_{ab}}{\partial s_{ij}} \right) = \frac{\partial (m^{-1})_{ab}}{\partial s_{kl}} \frac{\partial m_{ab}}{\partial s_{ij}} = -(m^{-1})_{ae} \frac{\partial m_{ef}}{\partial s_{kl}} (m^{-1})_{fb} \frac{\partial m_{ba}}{\partial s_{ij}}
= -(m^{-1})_{ae} (\delta_{ek} \delta_{fl} + \delta_{el} \delta_{fk}) (m^{-1})_{fb} (\delta_{bi} \delta_{aj} + \delta_{bj} \delta_{ai})
= -2 \left((m^{-1})_{ik} (m^{-1})_{jl} + (m^{-1})_{li} (m^{-1})_{jk} \right).$$
(B.30)

Apart from the final $(n+1)^{\text{th}}$ row and column, the inverse Cayley-Menger matrix is minus one half the Laplacian matrix \tilde{g}_{ij} as we showed in (4.37) and (4.39). This gives

$$\frac{\partial v_{ij}}{\partial s_{kl}} = -\frac{1}{2} \Big(\tilde{g}_{ik} \tilde{g}_{jl} + \tilde{g}_{il} \tilde{g}_{jk} \Big), \tag{B.31}$$

where $\tilde{g}_{ij} = -v_{ij}$ for $i \neq j$ and $\tilde{g}_{ii} = \sum_{a=1}^{n} v_{ia}$. For i, j, k, l all different, we therefore have

$$\frac{\partial v_{ij}}{\partial s_{kl}} = -\frac{1}{2} \Big(v_{ik} v_{jl} + v_{il} v_{jk} \Big), \qquad i \neq j \neq k \neq l$$
(B.32)

while if j = l,

$$\frac{\partial v_{ij}}{\partial s_{kj}} = -\frac{1}{2} \left(v_{ij} v_{jk} - v_{ik} \left(\sum_{a=1}^{n} v_{ja} \right) \right)$$
(B.33)

and if i = k and j = l,

$$\frac{\partial v_{ij}}{\partial s_{ij}} = -\frac{1}{2} \left(v_{ij}^2 + \left(\sum_{a=1}^n v_{ia} \right) \left(\sum_{b=1}^n v_{jb} \right) \right). \tag{B.34}$$

B.3 Landau singularities

The Landau singularities of the simplex integral are best studied in the Lee-Pomeransky representation (4.18). They follow from solving simultaneously for all v_{ij} the conditions

$$0 = \mathcal{U} + \mathcal{F}, \qquad 0 = v_{ij} \frac{\partial}{\partial v_{ij}} (\mathcal{U} + \mathcal{F}).$$
(B.35)

Here, the first Landau equation stipulates the vanishing of the Lee-Pomeransky denominator, while the second requires that this vanishing is either a double zero (for $v_{ij} \neq 0$), corresponding to a pinching of the v_{ij} integration contour between two converging singularities of the integrand, or else a pinch of the integration contour between a singularity and the end-point of the integration ($v_{ij} = 0$). The second condition thus ensures the singularity generated by the vanishing denominator cannot be avoided by a deformation of the integration contour. Where the Landau conditions have more than one solution, the solution with the greatest number of $v_{ij} \neq 0$ is referred to as the *leading* singularity. An important feature of the \mathcal{U} polynomial (4.10) is that it is *multilinear* in the v_{kl} : from the determinant structure one sees that all the quadratic v_{kl}^2 terms cancel, and that no higher powers can appear since v_{kl} enters only in the row/columns (k, k), (k, l), (l, k) and (l, l). Alternatively, this result follows from the matrix tree theorem where the Kirchhoff polynomial \mathcal{U} is the generator of spanning trees on the simplex. Since \mathcal{U} is also homogeneous of degree (n - 1), it follows that

$$\sum_{k < l} \frac{\partial \mathcal{U}}{\partial v_{kl}} v_{kl} = (n-1)\mathcal{U}.$$
 (B.36)

We now find

$$\mathcal{U} + \mathcal{F} = \mathcal{U} + \sum_{k < l} \frac{\partial \mathcal{U}}{\partial v_{kl}} V_{kl} = \sum_{k < l} \frac{\partial \mathcal{U}}{\partial v_{kl}} \left(\frac{v_{kl}}{n-1} + V_{kl} \right)$$
(B.37)

and so a solution of the first Landau condition for all k < l is

 $v_{kl} = \lambda V_{kl} = -\lambda \boldsymbol{p}_k \cdot \boldsymbol{p}_l$ and $|G| = |\boldsymbol{p}_k \cdot \boldsymbol{p}_l| = 0$ (B.38)

for some constant λ . Evaluating the second Landau condition on this solution (*) of the first gives

$$\left[v_{ij}\frac{\partial}{\partial v_{ij}}(\mathcal{U}+\mathcal{F})\right]_{*} = \left[v_{ij}\frac{\partial\mathcal{U}}{\partial v_{ij}} + v_{ij}\sum_{k(B.39)$$

using again the homogeneity of \mathcal{U} . The second Landau condition is thus solved for all i, j when $\lambda = 2 - n$, and indeed this is the leading singularity since the v_{kl} are generically nonzero. Returning to (B.37), on the solution (*) we have

$$(\mathcal{U} + \mathcal{F})_* = (2 - n)^{n-2} |G| = 0, \tag{B.40}$$

so to solve the first Landau condition we do indeed need the Gram determinant to vanish. Generally this requires analytic continuation to non-physical momentum configurations, since the only physical configurations (in Euclidean signature) for which the Gram determinant vanishes are collinear ones, and on physical grounds there are no collinear singularities. There is no contradiction here since the Landau equations are necessary, but not sufficient, conditions for a singularity.

B.4 Bernstein-Sato operators

In this appendix, we construct a Cayley-Menger analogue of the classic identity

$$\det(\partial)(\det X)^{a} = a(a+1)\dots(a+n-1)(\det X)^{a-1},$$
(B.41)

where $X = (x_{ij})$ is an $n \times n$ matrix of independent variables and $\partial = (\partial/\partial x_{ij})$ is the corresponding matrix of partial derivatives. For proofs and variants of this identity, traditionally attributed to Cayley, see, *e.g.*, [166, 167]. From a modern perspective, (B.41) is an example of a Bernstein-Sato operator, a differential operator whose action lowers the power *a* to which some polynomial of interest is raised, generating in the process an auxiliary polynomial in a known as the b-function [168]. Thus we have

$$\mathcal{B}_f f(x_{ij})^a = b_f(a) f(x_{ij})^{a-1}$$
(B.42)

where for (B.41), $\mathcal{B}_f = \det(\partial)$, $f = \det X$ and $b_f(a) = a(a+1)\dots(a+n-1)$. In the following, we construct analogous operators for the Cayley-Menger determinant and other polynomials arising in our parametric representations (4.48) and (4.17). Such relations are potentially a source of new weight-shifting operators, see *e.g.*, [169, 69].

Our starting point is the observation that

$$\mathcal{B}_{|m|} = (|g|) \Big|_{v_{ij} \to \partial_{s_{ij}}} \tag{B.43}$$

is a Bernstein-Sato operator for the Cayley-Menger determinant,

$$\mathcal{B}_{|m|} |m|^a = b_{|m|}(a) |m|^{a-1}, \qquad b_{|m|}(a) = -\prod_{k=1}^{n-1} (1-k-2a).$$
(B.44)

The operator $\mathcal{B}_{|m|}$ thus corresponds to evaluating the Kirchhoff polynomial $\mathcal{U} = |g|$ and replacing all $v_{ij} \to \partial_{s_{ij}}$ to generate a polynomial differential operator in the $\partial_{s_{ij}}$. We have verified (B.44) by direct calculation for matrices up to and including n = 5. Moreover, the leading behaviour at order a^{n-1} follows by noting that such terms can only arise from all n-1 partial derivatives in $\mathcal{B}_{|m|}$ hitting a power of |m| rather than a derivative of |m|. Using (4.35) in the form $\partial_{s_{ij}}|m|^a = av_{ij}|m|^a$ along with (4.34), then gives

$$\mathcal{B}_{|m|} |m|^a = a^{n-1} |m|^a |g| + O(a^{n-2}) = (-1)^n 2^{n-1} a^{n-1} |m|^{a-1} + O(a^{n-2})$$
(B.45)

in agreement with (B.44).²

Similarly, we find

$$\mathcal{B}_{|g|} = (|m|) \Big|_{s_{ij} \to \partial_{v_{ij}}} \tag{B.46}$$

(*i.e.*, the Cayley-Menger determinant replacing each $s_{ij} \rightarrow \partial_{v_{ij}}$) is the Bernstein-Sato operator for the Kirchhoff polynomial $\mathcal{U} = |g|$,

$$\mathcal{B}_{|g|}|g|^a = b_{|g|}(a)|g|^{a-1}, \qquad b_{|g|}(a) = -\prod_{k=1}^{n-1} (1-k-2a). \tag{B.47}$$

The *b*-function here is the same as that in (B.44), and the leading a^{n-1} behaviour can be understood via the analogous argument to that in (B.45). We note the result (B.47) is equivalent to Theorem 2.15 of [166], since $|m| = |m^{(n+1,n+1)}| - |m^{(n+1,n+1)} + J|$ where *J* is the $n \times n$ all-1s matrix, and $|m^{(n+1,n+1)}|$ is the Cayley-Menger minor formed by deleting the final row and column consisting of 1s and 0s. In addition, we find

$$\mathcal{B}_{|g|}(\partial_{v_{ij}}|g|)^a = 0. \tag{B.48}$$

Some further results worth recording are the following. For the second minors of the

 $^{^{2}}$ A full proof of (B.44) likely follows via the methods of [166], though we will not pursue this here.

Laplacian matrix, $|\tilde{g}^{(ij,ij)}| = \partial_{v_{ij}}|g|$, we find the operator

$$\mathcal{B}_{\partial_{v_{ij}}|g|} = (\partial_{s_{ij}}|m|)\Big|_{s_{kl} \to \partial_{v_{kl}}}$$
(B.49)

satisfies

$$\mathcal{B}_{\partial_{v_{ij}}|g|} \left(\partial_{v_{ij}}|g| \right)^a = b(a) \left(\partial_{v_{ij}}|g| \right)^{a-1}, \qquad \mathcal{B}_{\partial_{v_{ij}}|g|} \left|g\right|^a = b(a)v_{ij}|g|^{a-1}$$
(B.50)

where the b-function is proportional to that in (B.44) but is missing the final factor,

$$b(a) = 2 \prod_{k=1}^{n-2} (1 - k - 2a).$$
(B.51)

Again, we have verified these identities for values up to and including n = 5. This operator further annihilates all $(\partial_{v_{kl}}|g|)^a$ corresponding to other legs, *i.e.*,

$$\mathcal{B}_{\partial_{v_{ij}}|g|}(\partial_{v_{kl}}|g|)^a = 0 \quad \text{for all} \quad (i,j) \neq (k,l).$$
(B.52)

Similarly,

$$\left(\partial_{v_{ij}}|g|\right)\Big|_{v_{kl}\to\partial_{s_{kl}}}|m|^a = b(a)s_{ij}|m|^{a-1} \tag{B.53}$$

with the same *b*-function (B.51), but this operator does not appear to act simply (for n > 3) on $(\partial_{s_{ij}}|m|)^a$, in contrast to (B.50). Finding a Bernstein-Sato operator for $(\partial_{s_{ij}}|m|)^a$ would be useful since by (4.47) this corresponds to the Cayley-Menger minors featuring in (4.48).

In principle, given a Bernstein-Sato relation such as (B.44), one might hope to apply it inside the parametric representation (4.48) and integrate by parts to obtain an operator acting solely on the Schwinger exponential. Since the exponential is diagonal in the representation (4.48), the result could then be translated to a differential operator in the external momenta. This would then yield a new weight-shifting operator.

In practice, however, we must account for all the other powers of Cayley-Menger minors present in (4.48), as well as the arbitrary function. Either we must find a modified Bernstein-Sato operator that acts appropriately on the *entire* non-exponential prefactor in (4.48), which seems hard to do, or else we must find some means of removing and then restoring these other factors. The Cayley-Menger minors, for example, can be removed and then restored via a conjugation $\Omega \mathcal{B}_{|m|} \Omega^{-1}$ where $\Omega = \prod_{i<j}^{n} |m^{(i,j)}|^{-\alpha_{ij}-1}$. After multiplying out, however, this conjugated operator is not in the Weyl algebra (*i.e.*, is non-polynomial in the s_{ij} and their derivatives) and so does not trivially translate into an operator in the external momenta. On the other hand, if we include additional powers of the $|m^{(i,j)}|$ on the left, so as to recover an operator in the Weyl algebra, besides lowering α in (4.48) we also lower some of the α_{ij} . The operator then does not lower the spacetime dimension d. Thus we have not succeeded in finding new weight-shifting operators via this route, though with some variation the method might yet be successful.

Appendix C

Appendix to chapter 5

C.1 GKZ representation of Feynman integrals

In this appendix we relate a generic L-loop Feynman integral of the form (5.302) to the corresponding GKZ integral (5.303). Related discussions can be found in, *e.g.*, [123, 65].

After exponentiating the propagators and integrating out the loop momenta, (5.302) has the Schwinger parametrisation

$$I = (4\pi)^{-\gamma_0 L} \Big(\prod_{i=1}^N \frac{1}{\Gamma(\gamma_i)} \int_0^\infty \mathrm{d}t_i \, t_i^{\gamma_i - 1} \Big) \, \mathcal{U}[t]^{-\gamma_0} \exp\Big(-\frac{\mathcal{F}[t]}{\mathcal{U}[t]}\Big), \qquad \gamma_0 = \frac{d}{2}, \qquad (C.1)$$

where $\mathcal{U}[t]$ and $\mathcal{F}[t]$ are the first and second Symanzik polynomials respectively, which are homogeneous of weights L and L + 1 in the Schwinger parameters t_i . The prefactor of $(4\pi)^{-\gamma_0 L}$ is simply that in (5.302) multiplied by L factors of $\pi^{d/2}$ from integrating out the loop momenta. The corresponding Feynman representation is obtained by reparameterising

$$t_i = \sigma y_i, \qquad y_t = \sum_{i=1}^N y_i = 1$$
 (C.2)

and integrating out the variable σ . Using the Jacobian¹

$$\prod_{i=1}^{N} \mathrm{d}t_i = \sigma^{N-1} \mathrm{d}\sigma \prod_{i=1}^{N} \mathrm{d}y_i \,\delta(1-y_t),\tag{C.3}$$

as well as the homogeneity of the Symanzik polynomials, we find

$$I = (4\pi)^{-\gamma_0 L} \Big(\prod_{i=1}^N \frac{1}{\Gamma(\gamma_i)} \int_0^1 \mathrm{d}y_i \, y_i^{\gamma_i - 1} \Big) \delta(1 - y_t) \mathcal{U}[y]^{-\gamma_0} \int_0^\infty \mathrm{d}\sigma \, \sigma^{\gamma_t - \gamma_0 L - 1} \exp\left(-\sigma \frac{\mathcal{F}[y]}{\mathcal{U}[y]}\right)$$

$$= (4\pi)^{-\gamma_0 L} \Gamma(\gamma_t - \gamma_0 L) \Big(\prod_{i=1}^N \frac{1}{\Gamma(\gamma_i)} \int_0^1 \mathrm{d}y_i \, y_i^{\gamma_i - 1} \Big) \delta(1 - y_t) \mathcal{U}[y]^{\gamma_t - \gamma_0 (L+1)} \mathcal{F}[y]^{-\gamma_t + \gamma_0 L}.$$
(C.4)

¹See, e.g., Appendix B of [46].
In special cases where $\gamma_t - \gamma_0(L+1)$ vanishes (e.g., d = 2 multi-loop sunsets with standard propagators) one can use the \mathcal{F} polynomial alone to construct a GKZ representation [122]. More generally, one can use the Lee-Pomeransky representation [112, 123] obtained by combining the two Symanzik polynomial factors using the Euler beta identity

$$\mathcal{U}[y]^{-a}\mathcal{F}[y]^{a-b} = \frac{\Gamma(b)}{\Gamma(a)\Gamma(b-a)} \int_0^\infty \mathrm{d}s \, s^{a-1} (\mathcal{F}[y] + s \, \mathcal{U}[y])^{-b} \tag{C.5}$$

with $a = \gamma_0(L+1) - \gamma_t$ and $b = \gamma_0$ giving

$$I = c_{\gamma} \left(\prod_{i=1}^{N} \int_{0}^{1} \mathrm{d}y_{i} \, y_{i}^{\gamma_{i}-1} \right) \delta(1-y_{t}) \int_{0}^{\infty} \mathrm{d}s \, s^{\gamma_{0}(L+1)-\gamma_{t}-1} (\mathcal{F}[y] + s \, \mathcal{U}[y])^{-\gamma_{0}}$$
(C.6)

where

$$c_{\gamma} = \frac{(4\pi)^{-L\gamma_0} \Gamma(\gamma_0)}{\Gamma((L+1)\gamma_0 - \gamma_t) \prod_{i=1}^N \Gamma(\gamma_i)}.$$
 (C.7)

Setting $y_i = sz_i$ and using once again the homogeneity of the Symanzik polynomials, we can eliminate the s integral since

$$\int_0^\infty \frac{\mathrm{d}s}{s} \,\delta(1-sz_t) = \int_0^\infty \frac{\mathrm{d}s}{sz_t} \delta(z_t^{-1}-s) = 1 \tag{C.8}$$

after which

$$I = c_{\gamma} \left(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1}\right) (\mathcal{F}[z] + \mathcal{U}[z])^{-\gamma_{0}}.$$
 (C.9)

Finally, this Lee-Pomeransky representation is upgraded to the GKZ representation by replacing the coefficient of every term in the denominator $\mathcal{F}[z] + \mathcal{U}[z]$ with an independent variable x_k . For the massless triangle integral, for example,

$$\mathcal{U}[z] = z_1 + z_2 + z_3, \qquad \mathcal{F}[z] = p_1^2 z_2 z_3 + p_2^2 z_3 z_1 + p_3^2 z_1 z_2$$
(C.10)

and so we replace the Lee-Pomeransky denominator

$$\mathcal{G} = \mathcal{F}[z] + \mathcal{U}[z] = p_1^2 z_2 z_3 + p_2^2 z_3 z_1 + p_3^2 z_1 z_2 + z_1 + z_2 + z_3 \tag{C.11}$$

with the GKZ denominator

$$\mathcal{D} = x_1 z_2 z_3 + x_2 z_3 z_1 + x_3 z_1 z_2 + x_4 z_1 + x_5 z_2 + x_6 z_3.$$
(C.12)

The GKZ integral

$$\mathcal{I}_{\gamma} = \Big(\prod_{i=1}^{N} \int_{0}^{\infty} \mathrm{d}z_{i} \, z_{i}^{\gamma_{i}-1} \Big) \mathcal{D}^{-\gamma_{0}} \tag{C.13}$$

is then related to the massless triangle integral by

$$I = c_{\gamma} \mathcal{I}_{\gamma} \tag{C.14}$$

evaluated on the physical hypersurface

$$x_i = p_i^2, \qquad x_{i+3} = 1, \qquad i = 1, 2, 3.$$
 (C.15)

C.2 Creation operators for the position-space contact Witten diagram

In position space, the *n*-point AdS contact Witten diagram

$$I_n = \int_0^\infty \frac{dz}{z^{d+1}} \int d^d \boldsymbol{x}_0 \prod_{i=1}^n C_{\Delta_i} \left(\frac{z}{z^2 + x_{i0}^2}\right)^{\Delta_i}, \qquad C_{\Delta_i} = \frac{\Gamma(\Delta_i)}{\pi^{d/2} \Gamma(\Delta_i - \frac{d}{2})}, \tag{C.16}$$

has the parametric representation²

$$I_n = C_n \left(\prod_{i=1}^n \int_0^\infty dz_i \, z_i^{\Delta_i - 1}\right) \delta\left(1 - \sum_{i=1}^n \kappa_i z_i\right) \left(\sum_{i < j} z_i z_j x_{ij}^2\right)^{-\Delta_t / 2}.$$
 (C.17)

where

$$C_n = \frac{\pi^{d/2}}{2} \Gamma\left(\frac{\Delta_t}{2}\right) \Gamma\left(\frac{\Delta_t - d}{2}\right) \prod_{i=1}^n \frac{C_{\Delta_i}}{\Gamma(\Delta_i)}, \qquad \Delta_t = \sum_{i=1}^n \Delta_i, \qquad \boldsymbol{x}_{ij} = \boldsymbol{x}_i - \boldsymbol{x}_j. \quad (C.18)$$

The parameters $\kappa_i \geq 0$ can be chosen arbitrarily provided they are not all zero. For the 4-point function specifically, choosing $\kappa_i = \delta_{i4}$ and eliminating y_4 using the delta function leads to the GKZ representation

$$I_4 = C_4 \mathcal{I}_{\gamma}, \qquad \mathcal{I}_{\gamma} = \Big(\prod_{i=1}^3 \int_0^\infty \mathrm{d} z_i \, z_i^{\gamma_i - 1}\Big) \mathcal{D}^{-\gamma_0} \tag{C.19}$$

where

$$\mathcal{D} = x_1 z_2 z_3 + x_2 z_1 z_3 + x_3 z_1 z_2 + x_4 z_1 + x_5 z_2 + x_6 z_3, \tag{C.20}$$

the parameters

$$\gamma_1 = \Delta_1, \qquad \gamma_2 = \Delta_2, \qquad \gamma_3 = \Delta_3, \qquad \gamma_0 = \frac{1}{2}(\Delta_1 + \Delta_2 + \Delta_3 + \Delta_4), \qquad (C.21)$$

and the GKZ variables are related to the physical coordinate separations by

$$(x_1, x_2, x_3, x_4, x_5, x_6) = (x_{23}^2, x_{13}^2, x_{12}^2, x_{14}^2, x_{24}^2, x_{34}^2).$$
(C.22)

Comparing with (5.8), the position-space 4-point contact diagram, also known as the holographic D-function [170], is thus equivalent to the massless triangle integral (see also [46]). As shown on page 99, the massless triangle integral is itself equivalent to the triple-K integral (or momentum-space 3-point contact diagram) under affine reparametrisation of the GKZ integral. The creation operators for the position-space contact diagram are thus

²See, e.g., equations (5.46)-(5.51) and (B.1)-(B.11) of [46].

those analysed in section 5.5.3 and 5.5.5, except that no final projection to the physical hypersurface is required as all the GKZ variables in (C.22) are physical.

Concretely, the \mathcal{A} -matrix (5.22) leads to the Euler equations (5.23) and DWI (5.24), and the toric equations (5.27). The Newton polytope corresponds to the right-hand panel in figure 5.2. From its facets we obtain the singularity conditions

$$\gamma_{i} = -n_{i}, \qquad \gamma_{0} - \gamma_{i} = -m_{i}, \qquad i = 1, 2, 3, \gamma_{1} + \gamma_{2} + \gamma_{3} - \gamma_{0} = -n, \qquad 2\gamma_{0} - \gamma_{1} - \gamma_{2} - \gamma_{3} = -m,$$
(C.23)

where $n_i, m_i, n, m \in \mathbb{Z}^+$. The action of the annihilator $\partial_1 = \partial/\partial x_{23}^2$ is to raise γ_0, γ_2 and γ_3 by one which corresponds to raising Δ_2 and Δ_3 by one, and the action of the creation operator C_1 is the reverse of this. The corresponding *b*-function

$$b_1 = \gamma_2 \gamma_3 (\gamma_0 - \gamma_1) (\gamma_1 + \gamma_2 + \gamma_3 - \gamma_0),$$
(C.24)

when re-expressed in terms of Euler operator is

$$B_1 = (\theta_1 + \theta_3 + \theta_5)(\theta_1 + \theta_2 + \theta_6)(\theta_1 + \theta_5 + \theta_6)(\theta_1 + \theta_2 + \theta_3).$$
(C.25)

As expected, this is simply (5.183) under the mapping $\bar{\theta}_i = \theta_{i+3}$ since the affine reparametrisation from the \mathcal{A} -matrix (5.22) to (5.56) leaves the creation operators unchanged. Expanding out and using the toric equations to factorise $B_1 = C_1 \partial_1$, we recover the creation operator (5.202) in GKZ variables. In our present variables (C.22), this is

$$C_{1} = x_{1}(\theta_{1} + 1 + u_{2} + u_{3})((\theta_{1} + 1 + u_{2})(\theta_{1} + 1 + u_{3}) + 2(v_{2} + v_{3})) + x_{2}x_{5}\partial_{4}(1 + u_{2} + v_{2} - v_{3} + (u_{2} + u_{3} + 2)u_{3}) + x_{3}x_{6}\partial_{4}(1 + u_{3} + v_{3} - v_{2} + (u_{2} + u_{3} + 2)u_{2})$$
(C.26)

where $u_i = \theta_i + \theta_{i+3}$ and $v_i = \theta_i \theta_{i+3}$. One likewise obtains the operator (5.254), namely

$$W_{12}^{--} = (\theta_4 + \theta_5 + \theta_6 + \theta_3)(x_4\partial_2 + x_5\partial_1) + x_3x_6\partial_2\partial_1.$$
 (C.27)

Both these operators can be rewritten in various equivalent forms using the DWI and Euler equations. Their action on the position-space contact diagram follows from (5.59), namely

$$\mathcal{C}_1: \ \Delta_2 \to \Delta_2 - 1, \quad \Delta_3 \to \Delta_3 - 1, \qquad W_{12}^{--}: \ \Delta_3 \to \Delta_3 + 1, \quad \Delta_4 \to \Delta_4 - 1.$$
 (C.28)

C.3 Non-minimal b-functions

As we have seen, creation operators are constructed starting from a polynomial $b(\gamma)$ in the spectral parameters known as the *b*-function. In section 5.4.4, we argued that $b(\gamma)$ must possess a certain *minimal* set of zeros, namely, those required to cancel the singularities arising when a creation operator shifts us from a finite to a singular integral. Notice however that this argument does not preclude the existence of *additional* zeros besides this minimal set. For all the Feynman and Witten diagram examples in the main text, the minimal *b*-functions were sufficient for the construction of all creation operators. As these

b-functions contain the fewest factors, the resulting creation operators were moreover of lowest possible order in derivatives. Nevertheless, there are instances where the minimal *b*-function is not sufficient: a simple example, which we analyse in this appendix, is the GKZ integral (5.71). As we will show, additional factors must be appended to the minimal *b*-functions in order to be able to apply the toric equations and factorise into a product of creation and annihilation operators. The zeros of these additional factors are all parallel to the facets of the rescaled Newton polytope, and in most (though not all) cases correspond to additional singular hyperplanes of the GKZ integral.

Let us recall the necessary analysis of section 5.3. The integral (5.71), namely

$$\mathcal{I}_{\gamma} = \int_{0}^{\infty} \mathrm{d}z_{1} \int_{0}^{\infty} \mathrm{d}z_{2} \, z_{1}^{\gamma_{1}-1} z_{2}^{\gamma_{2}-1} (x_{1} + x_{2}z_{2} + x_{3}z_{1}^{2} + x_{4}z_{1}z_{2}^{2})^{-\gamma_{0}}, \tag{C.29}$$

corresponds to the \mathcal{A} -matrix

$$\mathcal{A} = \begin{pmatrix} 1 & 1 & 1 & 1 \\ 0 & 0 & 2 & 1 \\ 0 & 1 & 0 & 2 \end{pmatrix}$$
(C.30)

with DWI and Euler equations

$$0 = (\gamma_0 + \theta_1 + \theta_2 + \theta_3 + \theta_4)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_1 + 2\theta_3 + \theta_4)\mathcal{I}_{\gamma}, \quad 0 = (\gamma_2 + \theta_2 + 2\theta_4)\mathcal{I}_{\gamma}, \quad (C.31)$$

and a single toric equation

$$0 = (\partial_1^3 \partial_4^2 - \partial_2^4 \partial_3) \mathcal{I}_{\gamma}. \tag{C.32}$$

The singularities of this integral, derived in (5.93), are

$$\gamma_1 = -m_1, \quad \gamma_2 = -m_2, \quad \gamma_0 + \gamma_1 - \gamma_2 = -m_3, \quad 4\gamma_0 - 2\gamma_1 - \gamma_2 = -3m_4,$$
 (C.33)

for all $m_j \in \mathbb{Z}^+$. The annihilation operators ∂_j send $\gamma \to \gamma'$ while the creation operators \mathcal{C}_j send $\gamma' \to \gamma$, where for each j these parameters are related by

$$\begin{array}{ll} j = 1: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1, & \gamma'_2 = \gamma_2 \\ j = 2: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1, & \gamma'_2 = \gamma_2 + 1 \\ j = 3: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1 + 2, & \gamma'_2 = \gamma_2, \\ j = 4: & \gamma'_0 = \gamma_0 + 1, & \gamma'_1 = \gamma_1 + 1, & \gamma'_2 = \gamma_2 + 2. \end{array}$$

$$(C.34)$$

According to (5.138), the minimal *b*-functions containing only the zeros necessary to cancel the singularities produced by the action of the C_j are

$$b_1^{\min} = (\gamma_0 + \gamma_1 - \gamma_2) \prod_{m_4=0}^{1} (4\gamma_0 - 2\gamma_1 - \gamma_2 + 3m_4),$$

$$b_2^{\min} = \gamma_2 (4\gamma_0 - 2\gamma_1 - \gamma_2),$$

$$b_3^{\min} = \prod_{m_1=0}^{1} (\gamma_1 + m_1) \prod_{m_3=0}^{2} (\gamma_0 + \gamma_1 - \gamma_2 + m_3),$$

$$b_4^{\min} = \gamma_1 \prod_{m_2=0}^1 (\gamma_2 + m_2).$$
 (C.35)

For example, C_3 shifts $m_1 \to m_1 + 2$ and $m_3 \to m_3 + 3$, and so the five singular integrals with $m_1 = 0, 1$ and $m_3 = 0, 1, 2$ in (C.33) are accessible starting from finite integrals. This means b_3^{\min} has the five zeros shown, which act to cancel these singularities. The operator C_1 is however a special case: this sends $m_3 \to m_3 + 1$ and $m_4 \to m_4 + 4/3$, corresponding to a non-integer $F_1^{(4)} = 4/3$ in (5.139). Only the singularities with $m_3 = 0$ and $m_4 = 0, 1$ are then accessible starting from finite integrals for which all $m_j < 0$. (In other words, integrals for which the GKZ representation (C.29) converges without meromorphic continuation.)

Using the DWI and Euler equations to rewrite these *b*-functions in terms of Euler operators, we then find

$$B_1^{\min} = -(\theta_1 + 3\theta_3) \prod_{m_4=0}^1 (4\theta_1 + 3\theta_2 - 3m_4),$$

$$B_2^{\min} = (\theta_2 + 2\theta_4)(4\theta_1 + 3\theta_2),$$

$$B_3^{\min} = -(2\theta_3 + \theta_4)(2\theta_3 + \theta_4 - 1) \prod_{m_3=0}^2 (\theta_1 + 3\theta_3 - m_3),$$

$$B_4^{\min} = -(2\theta_3 + \theta_4)(\theta_2 + 2\theta_4)(\theta_2 + 2\theta_4 - 1).$$
 (C.36)

At this point a problem appears: to extract a creation operator requires factorising

$$B_j \mathcal{I}_{\gamma} = \mathcal{C}_j \partial_j \mathcal{I}_{\gamma}, \tag{C.37}$$

however the only toric equation we have available for this purpose, (C.32), is of *fifth* order in derivatives. While B_3^{\min} is indeed of fifth order, the remaining B_j^{\min} are of at most third order. Upon expanding out and ordering terms according to (5.131), we find

$$B_1^{\min} = (\dots)\partial_1 - 27x_2^2 x_3 \partial_2^2 \partial_3,$$

$$B_2^{\min} = (\dots)\partial_2 + 8x_1 x_4 \partial_1 \partial_4,$$

$$B_3^{\min} = (\dots)\partial_3 - x_1^3 x_4^2 \partial_1^3 \partial_4^2,$$

$$B_4^{\min} = (\dots)\partial_4 - 2x_2^2 x_3 \partial_2^2 \partial_3.$$
 (C.38)

For B_3^{\min} , we obtain the necessary factorisation (C.37) upon using (C.32) allowing a successful construction of C_3 . For the others, the order in derivatives is too low to apply (C.32).

To find C_1 , C_2 and C_4 , therefore, we look for new (non-minimal) B_j of the form:

$$B_{1} = (...)\partial_{1} + (...)x_{2}^{4}x_{3}\partial_{2}^{4}\partial_{3},$$

$$B_{2} = (...)\partial_{2} + (...)x_{1}^{3}x_{4}^{2}\partial_{1}^{3}\partial_{4}^{2},$$

$$B_{4} = (...)\partial_{4} + (...)x_{2}^{4}x_{3}\partial_{2}^{4}\partial_{3}.$$
 (C.39)

By construction, these are all of fifth order and can be factorised into the desired form

(C.37) using (C.32). Since the B_j must be functions of the Euler operators, and

$$x_i^n \partial_i^n = \theta_i(\theta_i - 1) \dots (\theta_i - n + 1), \tag{C.40}$$

this is equivalent to seeking

$$B_{1} = (\dots)\theta_{1} + (\dots)\theta_{2}(\theta_{2} - 1)(\theta_{2} - 2)(\theta_{2} - 3)\theta_{3},$$

$$B_{2} = (\dots)\theta_{2} + (\dots)\theta_{1}(\theta_{1} - 1)(\theta_{1} - 2)\theta_{4}(\theta_{4} - 1),$$

$$B_{4} = (\dots)\theta_{4} + (\dots)\theta_{2}(\theta_{2} - 1)(\theta_{2} - 2)(\theta_{2} - 3)\theta_{3}.$$
 (C.41)

For comparison, the singular hyperplanes in (C.33), when translated to Euler operators via (C.31), correspond to the zeros of

$$(2\theta_3 + \theta_4 - m_1), \quad (\theta_2 + 2\theta_4 - m_2), \quad (\theta_1 + 3\theta_3 - m_3), \quad (4\theta_1 + 3\theta_2 - 3m_4).$$
 (C.42)

As the non-minimal B_j in (C.41) must still contain the factors present in the minimal B_j^{\min} in (C.36), we see that for B_1 , and B_4 it suffices simply to append factors corresponding to additional singular hyperplanes:

$$B_{1} = -(\theta_{1} + 3\theta_{3}) \prod_{m_{4}=0}^{3} (4\theta_{1} + 3\theta_{2} - 3m_{4}),$$

$$B_{4} = -(2\theta_{3} + \theta_{4}) \prod_{m_{2}=0}^{3} (\theta_{2} + 2\theta_{4} - m_{2}).$$
 (C.43)

Each of these non-minimal B_j contain the factors already present in the minimal B_j^{\min} . Moreover, they are of the form (C.41) since they correspond to performing a linear shift in θ_j on each of the factors present in the second term of each B_j in (C.41). (Equivalently, setting θ_j to zero in each of the B_j in (C.43) yields the second term of each B_j in (C.41).) This also shows that they are of the smallest order in derivatives consistent with (C.41).

For B_2 , the additional factors we must append to B_2^{\min} are parallel to the singular hyperplanes in (C.42) but have different spacing. Explicitly, we require

$$B_2 = -\prod_{m_1=0}^{1} (\theta_2 + 2\theta_4 - 2m_1) \prod_{m_2=0}^{2} (4\theta_1 + 3\theta_2 - 4m_2)$$
(C.44)

so that, when expanded in θ_2 , we obtain an expression of the form given in (C.41). Note this is not possible using the spacings in (C.42).³

By construction, the non-minimal B_j in (C.43) and (C.44) all derive from corresponding non-minimal *b*-functions which are polynomials in the spectral parameters,

$$b_1 = (\gamma_0 + \gamma_1 - \gamma_2) \prod_{m_4=0}^3 (4\gamma_0 - 2\gamma_1 - \gamma_2 + 3m_4),$$

³The zeros of (C.44), and of the corresponding of b_2 in (C.45), do however coincide with the singular hyperplanes of the integral obtained by deleting the second column of the \mathcal{A} -matrix. This removes a vertex from the Newton polytope changing the spacings of the singular hyperplanes; a procedure consistent with confining all θ_2 dependence in B_2 to the first factor in (C.41).



Figure C.1: For purposes of illustration, a two-dimensional example of a non-normal lattice can be obtained by projecting to the (γ_1, γ_2) plane. The points (2k + 1, 0) and (2k + 1, 1) for $k \in \mathbb{Z}^+$ (shown in red) lie within the positive cone $\sum_j \mathbb{R}^+ a_j$ (shaded region) spanned by the vertex vectors a_j of the Newton polytope (shown in blue). These points also belong to $\sum_j \mathbb{Z} a_j$ since (2k + 1, 0) =k(2, 0) + (1, 2) - 2(0, 1) and (2k + 1, 1) = k(2, 0) + (1, 2) - (0, 1). However, these points cannot be expressed in the form $\sum_j \mathbb{Z}^+ a_j$ and hence the lattice generated by the a_j is non-normal.

$$b_{2} = \prod_{m_{1}=0}^{1} (\gamma_{2} + 2m_{1}) \prod_{m_{2}=0}^{2} (4\gamma_{0} - 2\gamma_{1} - \gamma_{2} + 4m_{2}),$$

$$b_{4} = \gamma_{1} \prod_{m_{2}=0}^{3} (\gamma_{2} + m_{2}).$$
 (C.45)

and all lead to valid creation operators via (C.37). From B_2 , for example, we find

$$\mathcal{C}_{2} = x_{2} \Big[12\theta_{4}^{2} \Big(48\theta_{1}^{2} + 12\theta_{1}(3\theta_{2} - 5) + 9\theta_{2}(\theta_{2} - 2) + 5 \Big) \\ + 4\theta_{4} \Big(64\theta_{1}^{3} + 48\theta_{1}^{2}(3\theta_{2} - 4) + 4\theta_{1} \Big(9\theta_{2}(3\theta_{2} - 5) + 32 \Big) + 3\theta_{2} \Big(9\theta_{2}(\theta_{2} - 2) + 5 \Big) \Big) \\ + (\theta_{2} - 1)(4\theta_{1} + 3\theta_{2} - 5)(4\theta_{1} + 3\theta_{2} - 1)(4\theta_{1} + 3\theta_{2} + 3) \Big] + 256x_{1}^{3}x_{4}^{2}\partial_{2}^{3}\partial_{3}. \quad (C.46)$$

Having solved this example, let us note that the failure of the minimal *b*-functions in (C.35) can also be understood geometrically. For B_1^{\min} , B_2^{\min} and B_4^{\min} in (C.36) to be factorisable as $C_j \partial_j$, we would need each of $\partial_2^2 \partial_3 \partial_1^{-1}$, $\partial_1 \partial_4 \partial_2^{-1}$ and $\partial_2^2 \partial_3 \partial_4^{-1}$ to be expressible as $\prod_{k=1}^N \partial_k^{c_k}$ for some set of $c_k \in \mathbb{Z}^+$. (The inverses here are purely formal: we mean that $\partial_2^2 \partial_3 = \partial_1 \prod_{k=1}^N \partial_k^{c_k}$, etc.) In terms of the shifts produced by these operators on the spectral parameters $\boldsymbol{\gamma} = (\gamma_0, \gamma_1, \gamma_2)$, this is equivalent to requiring that

$$2\boldsymbol{\mathcal{A}}_{2} + \boldsymbol{\mathcal{A}}_{3} - \boldsymbol{\mathcal{A}}_{1} = \begin{pmatrix} 2\\2\\2 \end{pmatrix}, \quad \boldsymbol{\mathcal{A}}_{1} + \boldsymbol{\mathcal{A}}_{4} - \boldsymbol{\mathcal{A}}_{2} = \begin{pmatrix} 1\\1\\1 \end{pmatrix}, \quad 2\boldsymbol{\mathcal{A}}_{2} + \boldsymbol{\mathcal{A}}_{3} - \boldsymbol{\mathcal{A}}_{4} = \begin{pmatrix} 2\\1\\0 \end{pmatrix} \quad (C.47)$$

are all expressible as $\sum_{k=1}^{N} c_k \mathcal{A}_k$ for some set of $c_k \in \mathbb{Z}^+$, where \mathcal{A}_k denotes the *k*th column of the \mathcal{A} -matrix including the top row of ones. Clearly this is not possible, although these vectors do all lie in the positive cone corresponding to solutions with $c_k \in \mathbb{R}^+$ since

$$2\mathcal{A}_{2} + \mathcal{A}_{3} - \mathcal{A}_{1} = 2(\mathcal{A}_{1} + \mathcal{A}_{4} - \mathcal{A}_{2}) = \frac{2}{3}(\mathcal{A}_{2} + \mathcal{A}_{3} + \mathcal{A}_{4}),$$

$$2\mathcal{A}_{2} + \mathcal{A}_{3} - \mathcal{A}_{4} = \frac{1}{2}(3\mathcal{A}_{1} + \mathcal{A}_{3}).$$
 (C.48)

Mathematically, this is precisely the condition that lattice generated by the \mathcal{A}_k (or equivalently, the toric ideal associated with the \mathcal{A} -matrix) is *non-normal*. Conversely, when the normality condition

$$\left(\sum_{k} \mathbb{R}^{+} \mathcal{A}_{k}\right) \cap \left(\sum_{k} \mathbb{Z} \mathcal{A}_{k}\right) = \left(\sum_{k} \mathbb{Z}^{+} \mathcal{A}_{k}\right)$$
(C.49)

is satisfied, it can be shown that the minimal b-functions (5.138) generate valid creation operators [146, 147]. The non-triviality of the normality condition is illustrated in figure C.1.

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